



### Herschel Spectrometers Workshop PACS Spectrometer Session #4 23 April 2014 E. Puga, Herschel Science Centre





- Session #4: Combining PACS and SPIRE spectra
  - 1. Point sources
  - 2. Slightly extended sources (modify useful script to contain both parts, but just show figure for photometry)

Outline





## PACS and SPIRE for Point Sources

23 Apr 2014

Herschel Spectrometers Workshop: PACS #4

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## Point Source Calibrated Spectra



- PACS: After level 2 + post-processing steps
- SPIRE: level 2 already contains the point source calibration
- Stitching PACS+SPIRE spectrum is contained in a useful script in HIPE





- Stitching PACS and SPIRE spectra of a point source
- Scripts >> PACS Useful Scripts >> Spectroscopy: Combine PACS and SPIRE spectra
- Source: (default) CRL 618





## PACS and SPIRE for Semi-extended Sources

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Semi-extended Sources



- Semi-extended source corrections in HIPE:
  - PACS: specExtendedToPointCorrection (E2P)
  - SPIRE: Semi-Extended Correction Tool (SECT), see full presentation tomorrow by Ed Polehampton







PACS

**SPIRE** 





## Are these two corrections comparable?



### • SPIRE-FTS

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Following the details of SECT are in Wu et al. 2013:

$$F_{s} = I_{ext} \cdot C_{point} \cdot \eta_{c}(\nu, \Omega_{source}) \cdot \frac{\iint\limits_{2\pi} P_{\nu}(\Psi) \delta_{\nu}(\Psi) d\Psi}{\iint\limits_{2\pi} P_{\nu}(\Psi - \Omega_{0}) D_{\nu}(\Psi) d\Psi} \cdot \iint\limits_{2\pi} P_{ref}(\Psi) D_{\nu}(\Psi) d\Psi$$

...SECT introduces a limitation of the flux density to that within a reference beam (default Gaussian beamSize=42" for SECT)

$$F_{s} = I_{ext} \cdot C_{point} \cdot \eta_{c}(\nu, \Omega_{source}) \cdot \frac{\iint_{i,j} P_{\nu}(\Psi) \delta_{\nu}(\Psi) d\Psi}{\iint_{i,j} P_{\nu}(\Psi - \Omega_{0}) D_{\nu}(\Psi) d\Psi} \cdot \iint_{i,j} D_{\nu}(\Psi) d\Psi$$



A reference beam that does not encompass the entire source spatial extent, preserves the corrected shape of the SPIRE-FTS spectrum, but at a smaller absolute level  $\rightarrow$  when comparing with PACS spectra, make reference beam large enough, compared to the source extent.

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## •esa Hands-on #2



- Merging PACS and SPIRE spectra of a semi-extended source
- Source: NGC7027 protoplanetary nebula
- PACS
  - Local data, Ipipe processing: Telescope Background Normalization, flatfield(excludeLeaks=1)
  - Obsids: 1342186968, 1342186969
  - Source model: Gaussian FWHM 12"
- SPIRE-FTS
  - Local data: cal\_spire\_11\_0
  - Obsid: 1342259592
  - Source model: Gaussian FWHM 12" and gaussRefBeamDiameter=60.0"

# Cesa Additionally



- Photometric points, with and without color correction
- ISO LWS

