

# N-BEARING MOLECULES IN THE PROTOSTELLAR ENVELOPE OF AFGL 2591

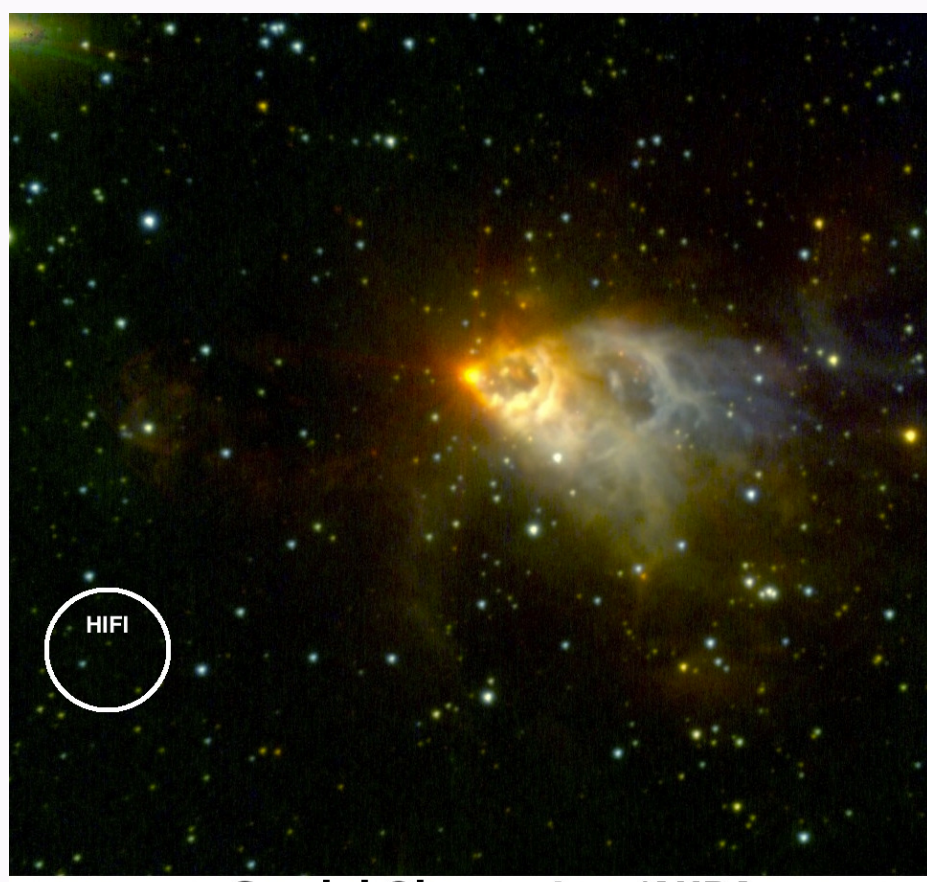


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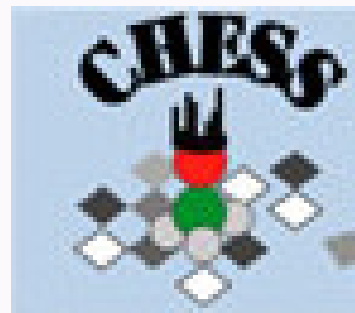
## 1. AFGL 2591



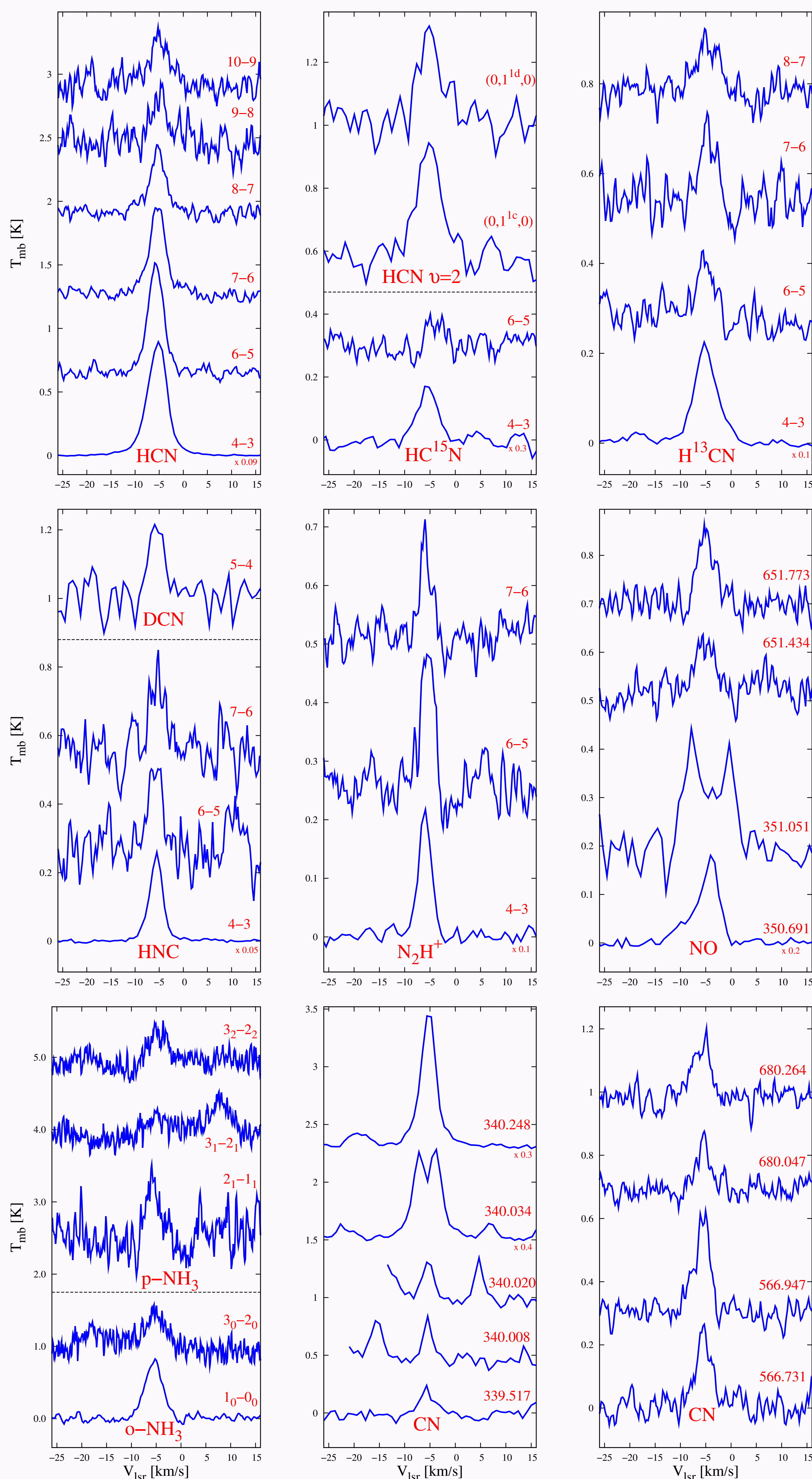
Gemini Observatory/AURA

- ▶ AFGL 2591 – high mass protostellar object with a bipolar outflow (Van der Tak et al. 1999)
- ▶ located in the Cygnus X region ( $l, b$ )=78.°9, 0.°71
- ▶ relatively isolated massive star-forming region

- ▶ one of the targets of the CHESs project (Ceccarelli et al. 2010)
- ▶ luminosity =  $2 \times 10^5 L_{\odot}$  (Sanna et al. 2012)
- ▶ distance = 3 kpc (Rygl et al. 2012)



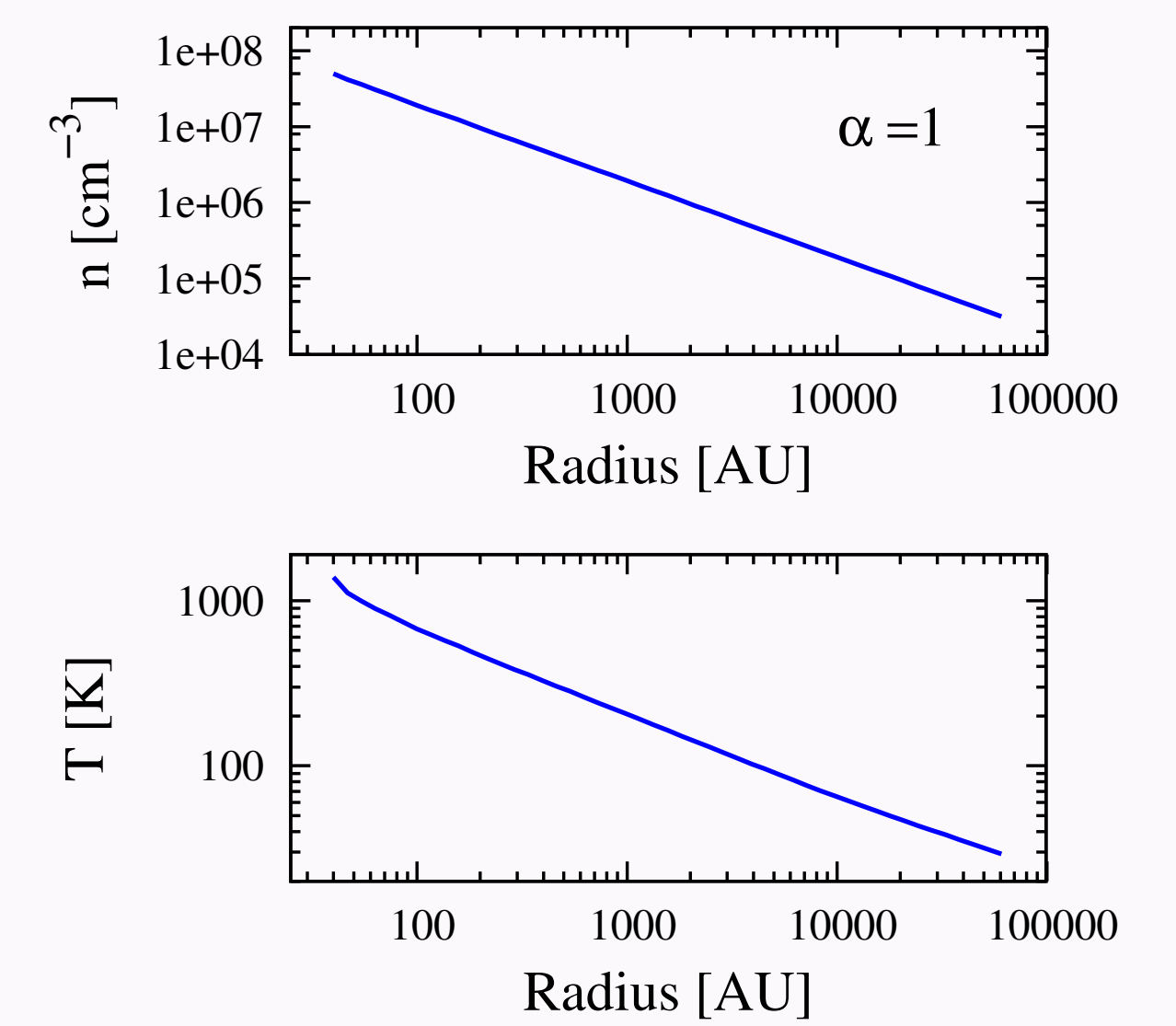
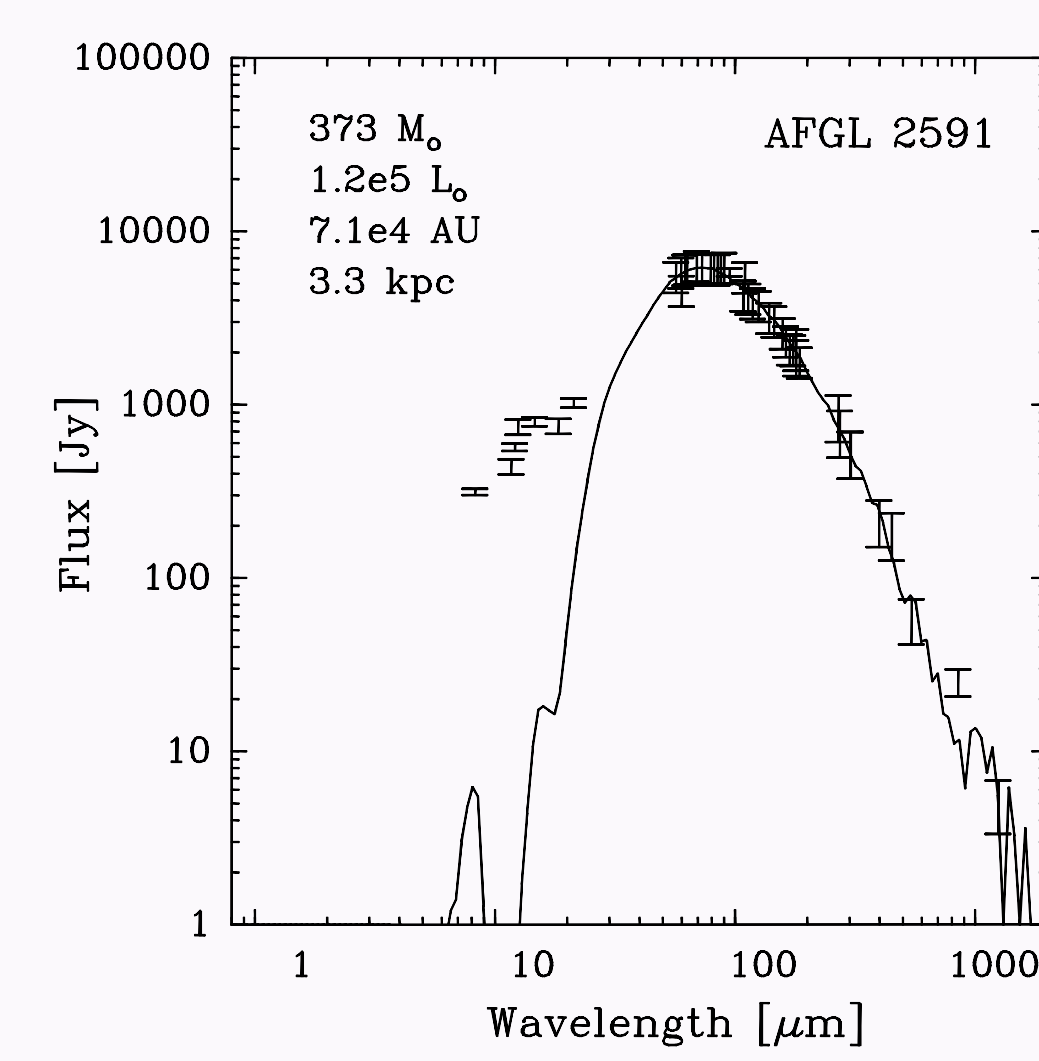
## 2. Line profiles (HIFI & JCMT data)



## 3. Excitation analysis

- ▶ Based on population diagrams method (Goldsmith & Langer 1999), N-bearing molecules represent cold species, with the excitation temperatures below 50 K.
- ▶ In contrast several warm species were found (e.g. CH<sub>3</sub>OH, SO, SO<sub>2</sub>) with temperatures above 100 K.

## 4. Physical model

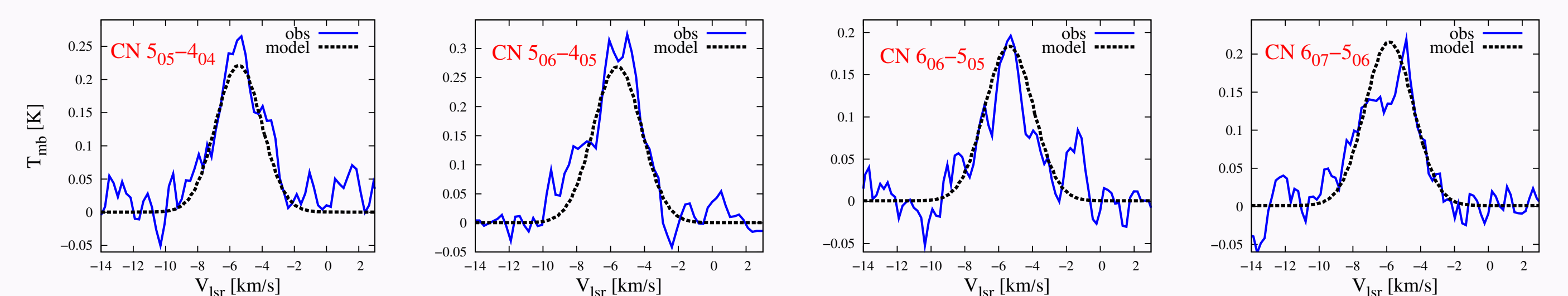


For more details see (Van der Tak et al. 2013).

## 5. Radiative transfer analysis

Abundances were estimated using Ratran code (Hogerheijde & van der Tak 2000).

- ▶ Some molecules are evenly distributed through the envelope:  $[N_2H^+] = 5 \times 10^{-10}$ ,  $[NO] = 2 \times 10^{-8}$  and  $[CN] = 8 \times 10^{-11}$ .



- ▶ HNC, HCN and its isotopologues are more abundant in the inner envelope, when  $T > 230$  K ( $[HCN] = 1 \times 10^{-5}$ ,  $[HNC] = 3 \times 10^{-7}$ ), than in the outer part of the envelope (when  $T < 230$  K,  $[HCN] = 1 \times 10^{-7}$ ,  $[HNC] = 3 \times 10^{-7}$ ). This temperature was predicted by the chemical models in which most of the atomic oxygen is driven into water. As a result atomic C and N abundances are higher, thus HCN abundance is increased as well at  $T > 230$  K.

(For more info about jump model at 230 K, see e.g. Boonman et al. 2001.)

- ▶ NH<sub>3</sub> is concentrated in the inner part of the envelope, when  $T > 100$  K, i.e., where water ice evaporates, with an abundance of  $6 \times 10^{-7}$ .

(For more info about jump model at 100 K, see e.g. Van der Tak et al. 2006).

## 6. Conclusions

The envelope of AFGL 2591 consists of 3 chemical zones, separated by the onset of water ice evaporation ( $T \sim 100$  K) and high temperature gas phase oxygen chemistry ( $T \sim 230$  K).

## Contact

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