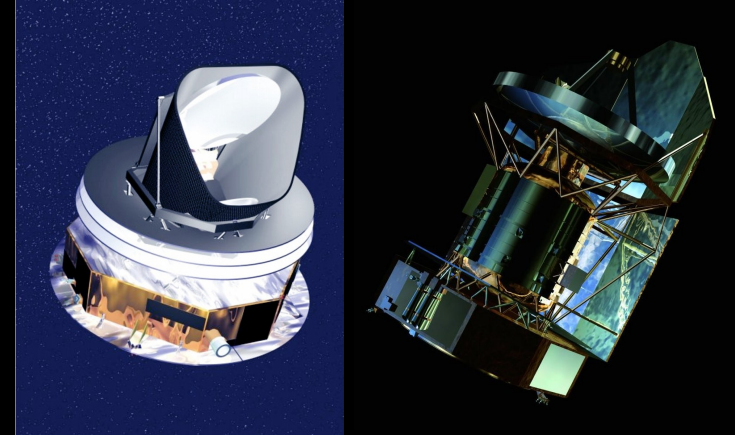


# Cold Cores

on Planck<sup>1,2</sup> and Herschel<sup>1</sup>

*on behalf of the Planck collaboration*



**M. Juvela, I. Ristorcelli** (coord.)

## **Planck**

Desert, Dupac, Giard, Harju, Harrison, Joncas, Jones, Lagache, Lamarre, Laureijs, Lehtinen, Maffei, Martin, **Marshall**, Malinen, Mattila, **McGehee**, **Montier**, Pajot, Paladini, **Pelkonen**, Tauber, Taylor, Valenziano, Verstraete, Ysard Abergel, Bernard, Boulanger, Cambresy, Davies, Dickinson, Fischera, Macias-Perez, Meny, Miville-Deschenes, Nartallo, Pagani, Puget, Reach

## **Herschel**

Andre, Kiss, Klaas, **Krause**, Molinari, Motte, Schneider, **Toth**, Ward-Thompson, **Zavagno**

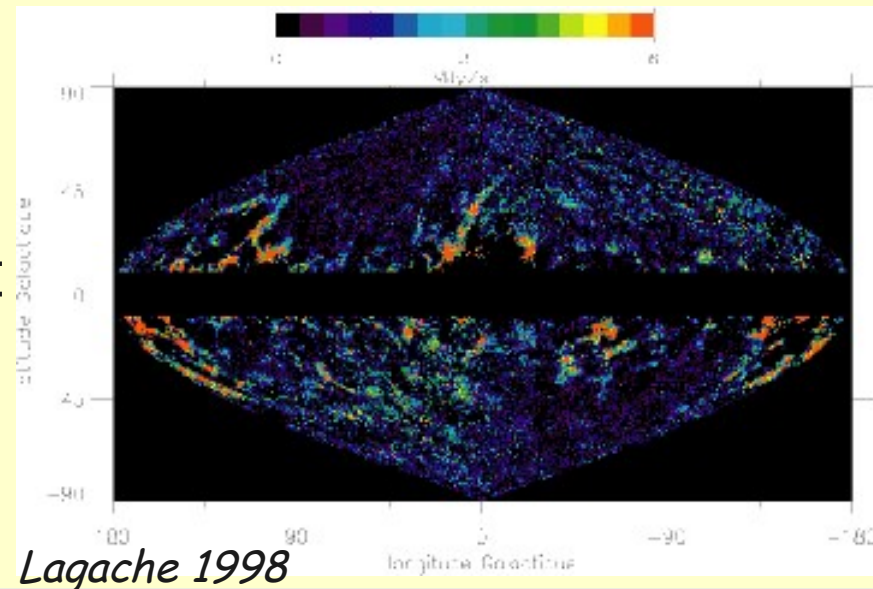
## **External**

Doi, Ueno, Kitamura, Nikeda, Kawamura, Onishi

*With acknowledgement to ESA<sup>1</sup> and the Planck HFI & LFI consortia<sup>2</sup>*

# Scientific Justification

- Cold dust: a **tracer** of dense, hidden regions of star-forming clouds
- Access to the **earliest phases of star formation**
  - What generates pre-stellar **cores** and what governs their evolution to protostars and proto-brown dwarfs ?
  - Origin of the global stellar initial mass function (IMF)
- Part of the life **cycle of dust**
  - from diffuse medium  
dense clouds
- as first step, we need to detect **cold and compact Galactic dust clouds**

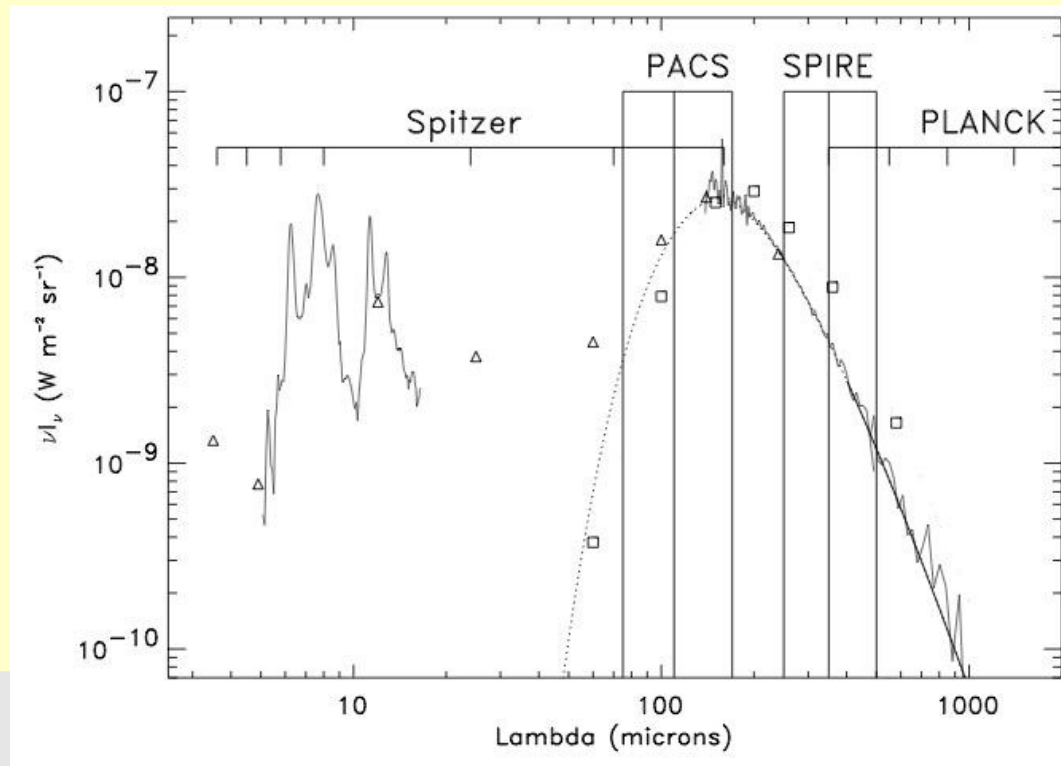
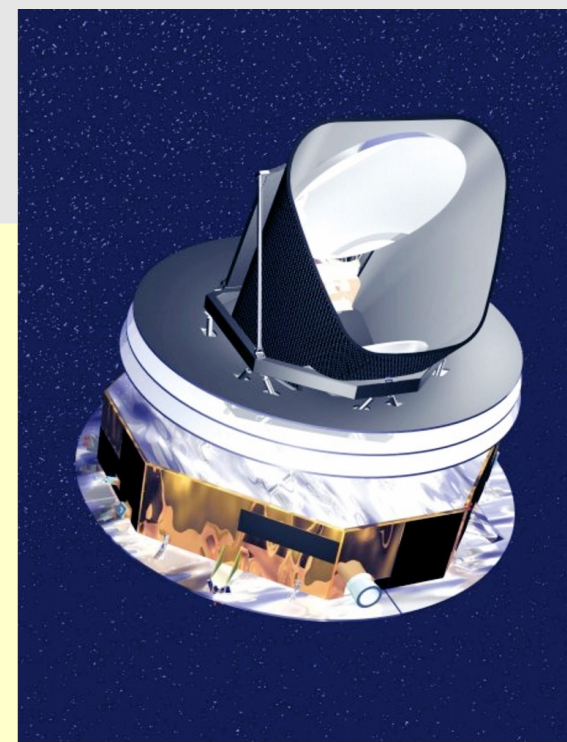


*Lagache 1998*

# Planck

**The Planck satellite is mapping the sky at 9 frequencies in the sub-mm/mm range**

- main goal is sensitive mapping of CMB fluctuations
- wide wavelength coverage allows good separation of foreground components
  - 350, 550, 850, 1380, 2100, ..., 10000  $\mu\text{m}$
  - 5.0, 5.0, 5.0, 5., 7.1, ..., 33 arcmin



# Cold Cores: The Planck project

PLANCK provides the first **unbiased** cold core catalog

- It will include a large list of 'normal' cloud cores
  - Pre-stellar dust cores in molecular clouds (class -I), isolated globules, dark clouds in the Galactic plane, etc.
  - a large number of new objects at high Galactic latitudes
  - many so far undetected objects (very low T)
  - possibility for the discovery of new classes of objects
    - e.g., Jupiter mass clouds?

We expect **>>1000** detections

## Work within **the Planck project**:

- Analysis of Planck frequency maps
- Extract cold cores (paper to be submitted shortly, *Montier et al., 2010*)
- Cross correlate the findings with ancillary data
- Construct **C3PO** = Cold Core Catalog of Planck Objects

## Select targets for the **Herschel follow-up**

- more detailed studies of a representative set of cores
  - higher spatial resolution, inclusion of shorter wavelengths
- the goal: a coherent observational database representing the entire cold core population of the Galaxy, complementing the GT projects that target currently actively star forming regions

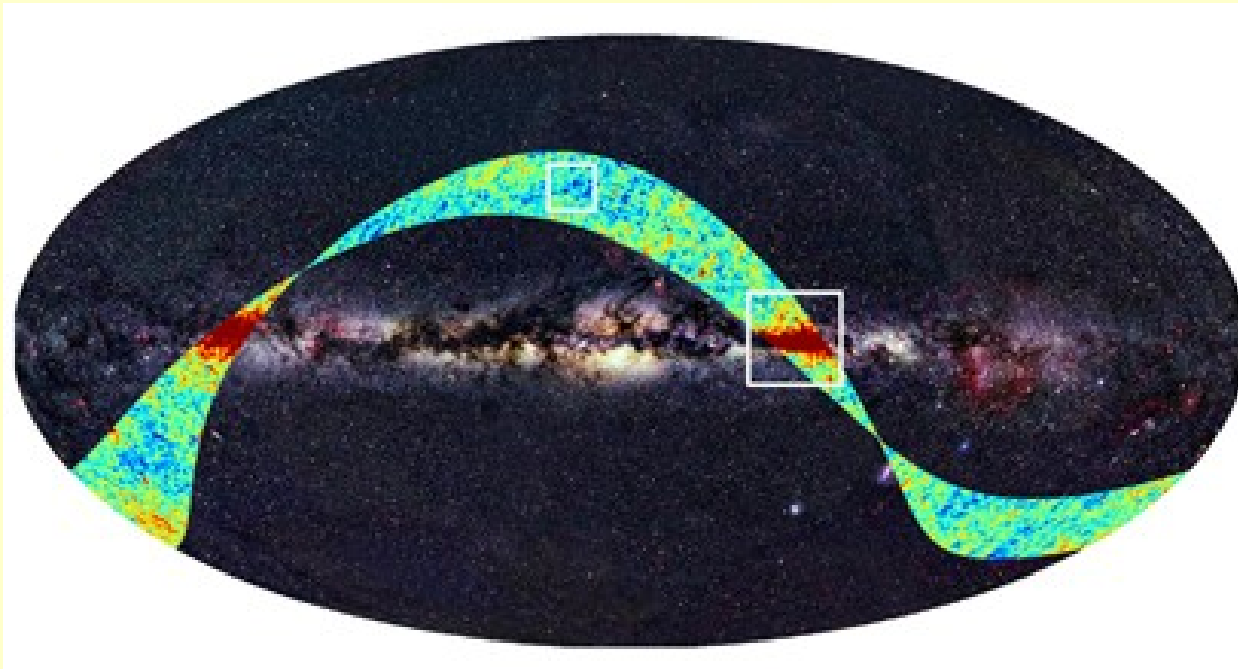
Observations could start in late 2010! ...

# Cold Cores: The Herschel project

- OT KP *Galactic Cold Cores* (150.9h)
- observations consist of small ( $\sim 20$  arcmin) PACS and SPIRE maps of carefully pre-selected cores
  - selection using Planck and ancillary data
  - a **cross-section** of Galactic cold cores population
  - selection takes into account **physical properties** of the cores (temperature, mass, density, size), their **location** (high/low latitudes, inner/outer Galaxy), **environment** (clustered vs. isolated sources, magnetic fields), **dust properties** (emissivity index, signs of anomalous microwave emission)
  - special emphasis is put on rare source types (e.g., high latitudes, both ends of the core mass spectrum)
  - 151 hours  $\sim$  100-150 fields

# SDP

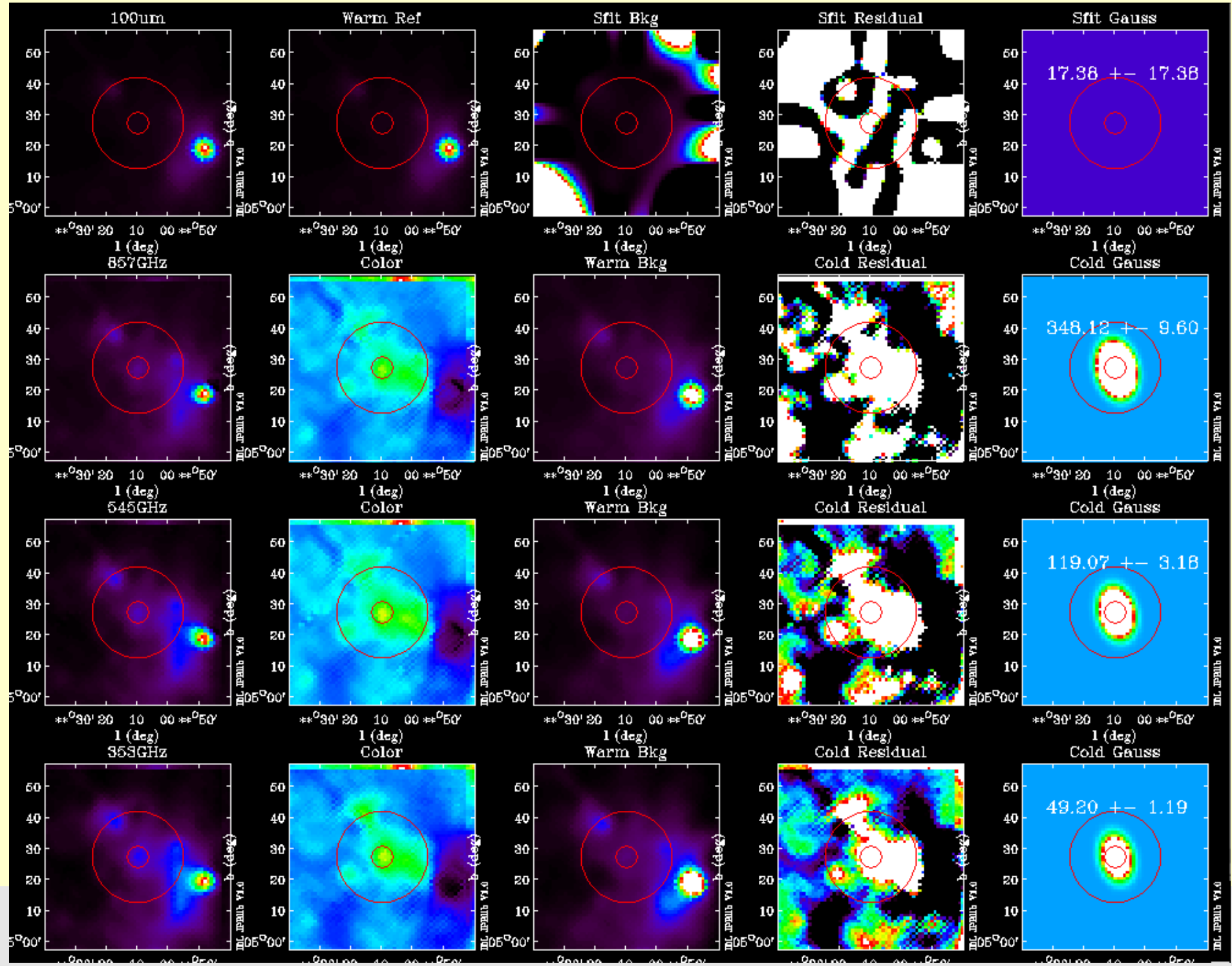
... analysis of **Planck First Light Survey** resulted in the detection of many new sources!



Copyright:  
ESA, LFI & HFI Consortia (Planck),  
Background image:  
Axel Mellinger

this gave the possibility to have targets **already in SDP**

# PCC249 in Planck





# SDP targets

- three secure detections from Planck First Light Survey
- $T_d \sim 13\text{K}$  or below
- visible in SDP and outside GT fields
- test both parallel mode and normal scan mapping with PACS and SPIRE

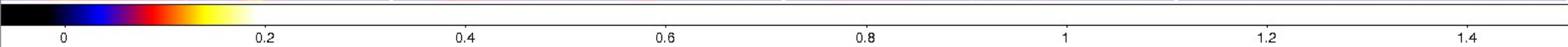
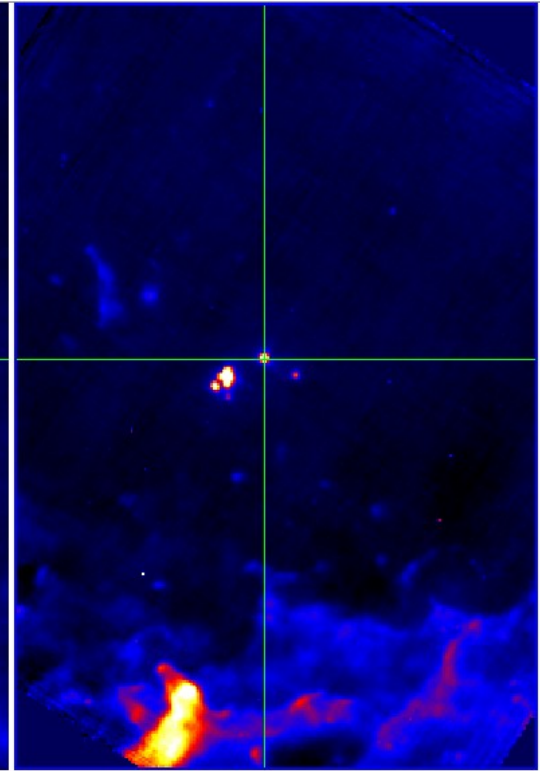
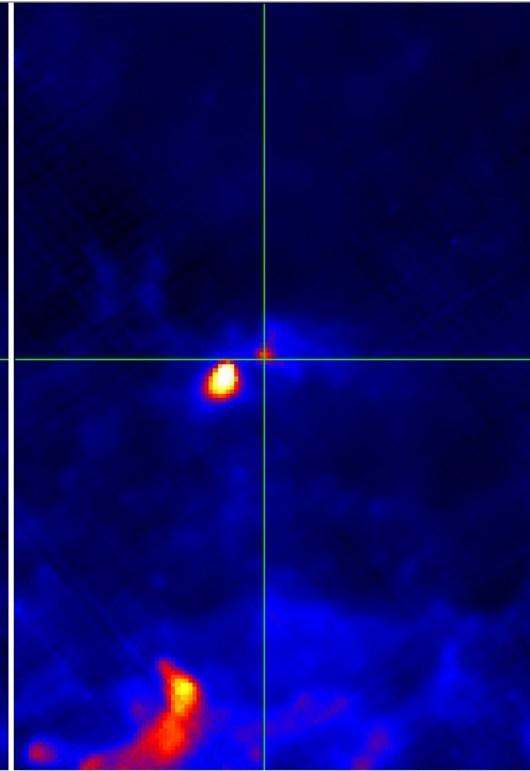
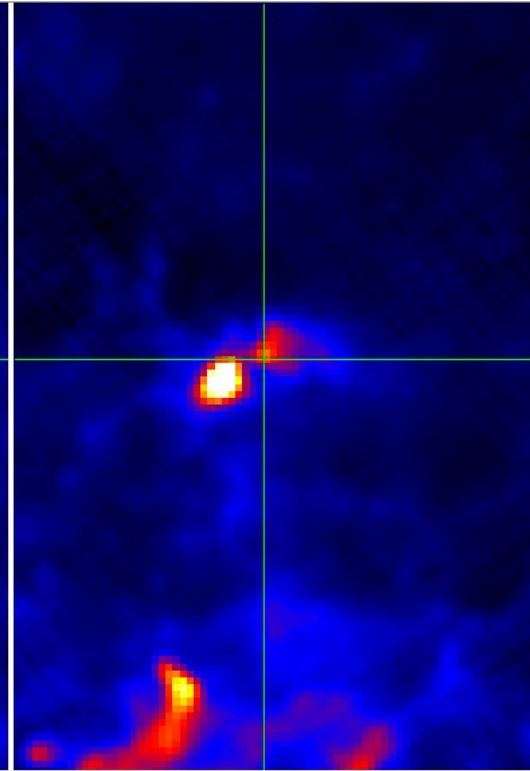
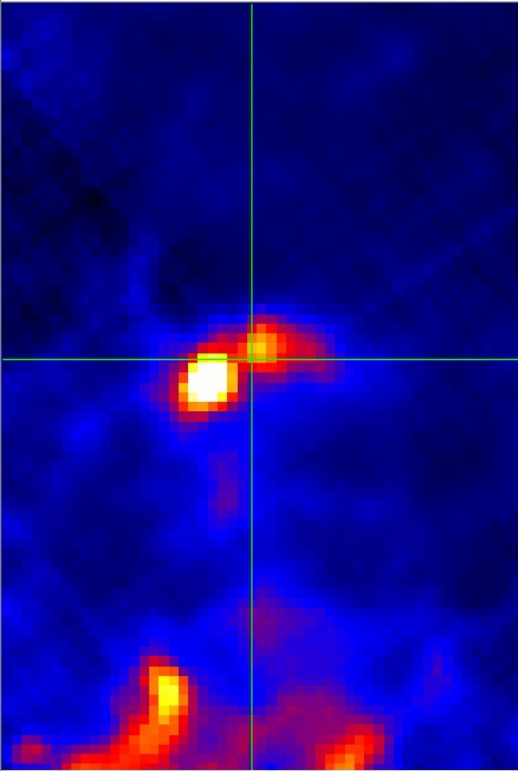
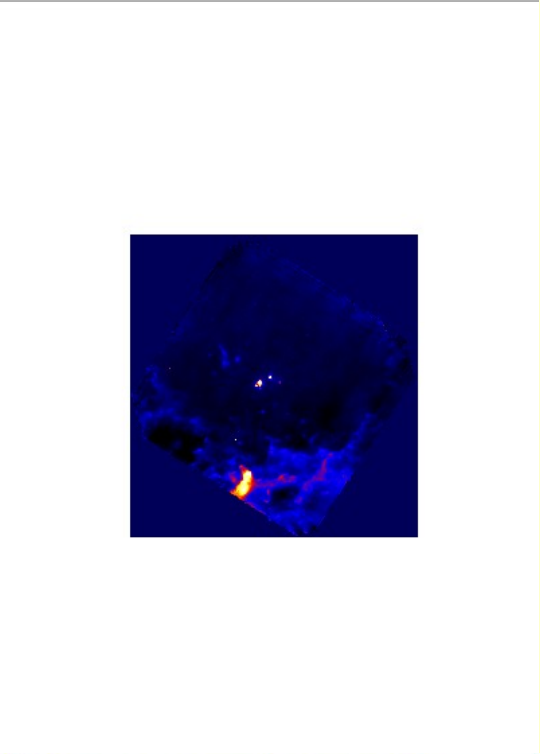
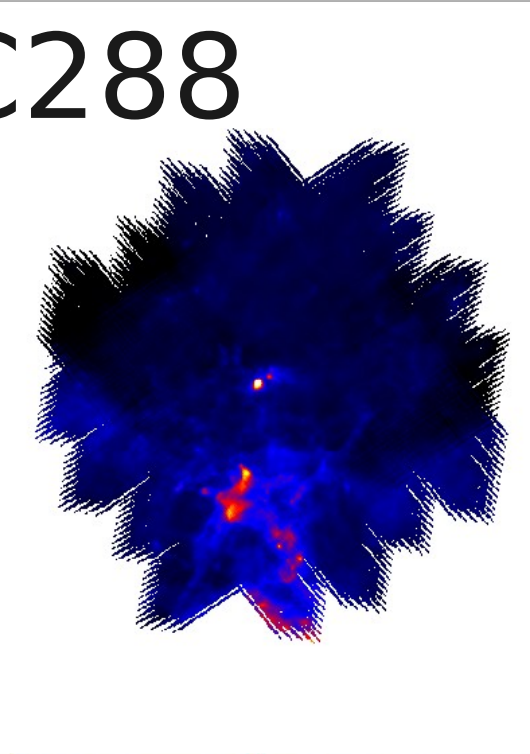
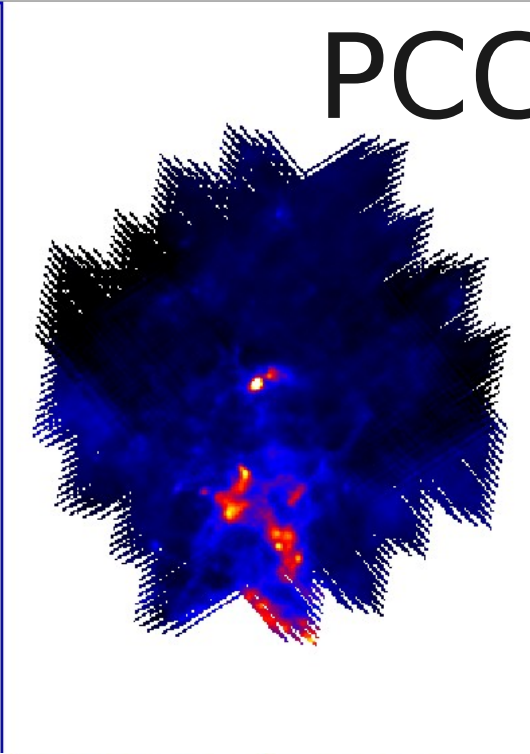
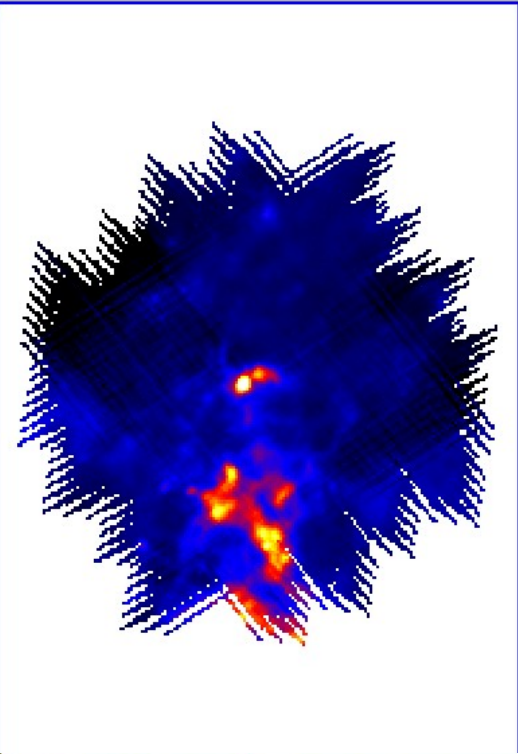
Target ID	Position (Galactic / J2000)	IRAS 100 $\mu\text{m}$ <sup>1</sup> (MJy/sr)	$T_{\text{dust}}$ (K)
PCC249	107.20+5.52 22 21 17.6 +63 42 25	150	12.2
PCC550	300.86-9.00 12 25 16.5 -71 46 03	25	12.6
PCC288	109.80+2.70 22 53 31.3 +62 31 44	120	11.3

<sup>1</sup>Level of local background (IRAS gives only upper limits of source fluxes)

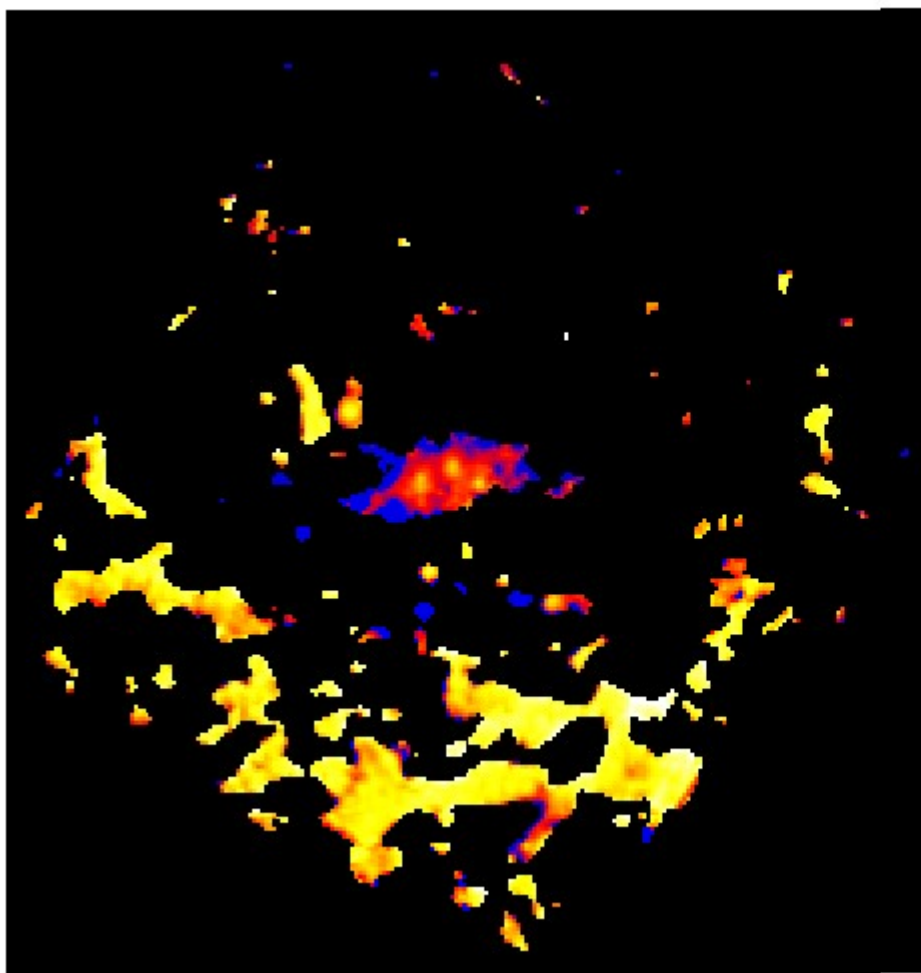
# SDP observations

- PCC288
  - PACS: three AORs, total ~0.5 hours; 15 arcmin square, homogeneous coverage
  - SPIRE: 20 arcmin, orthogonal scans
- PCC249
  - parallel mode, orthogonal scans,  $\sim 30 \times 30'$ ; ~2 hours
- PCC550
  - PACS: three AORs (0+45+135 deg), 15 arcmin; ~1 hour

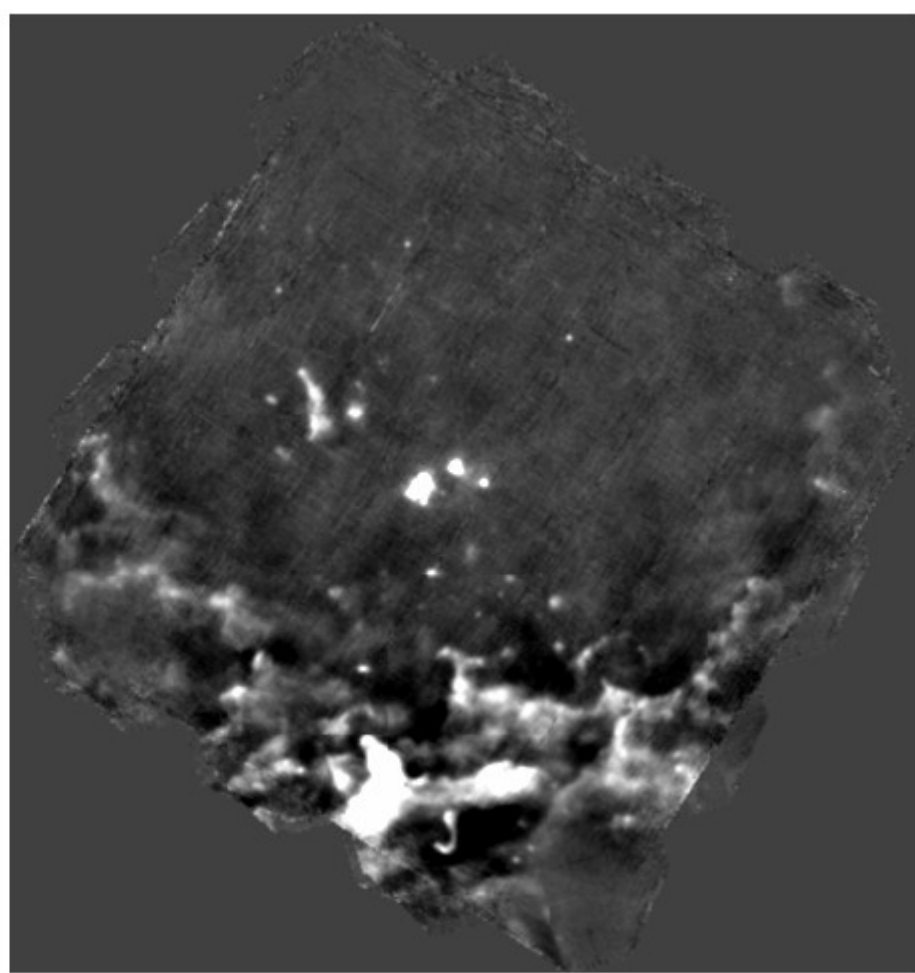
# PCC288



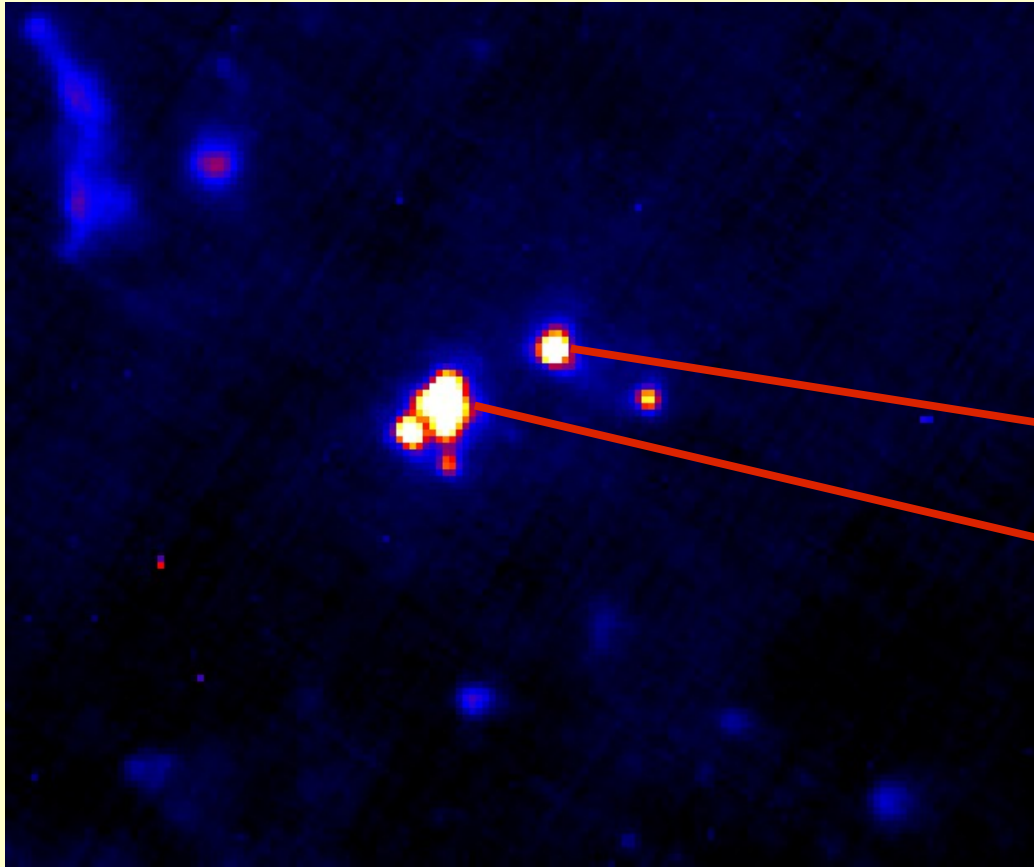
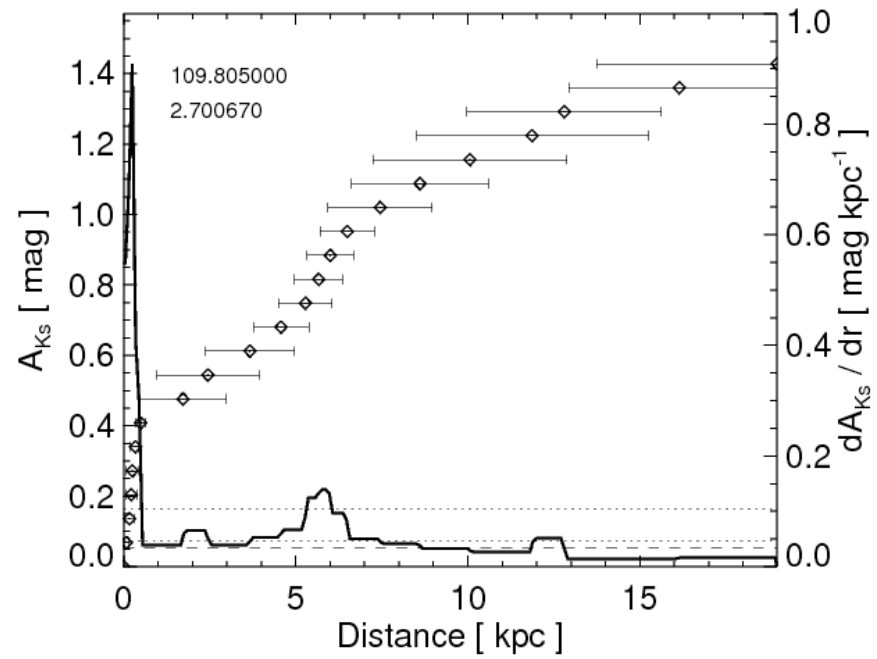
*Relative temperature*



*PACS green*

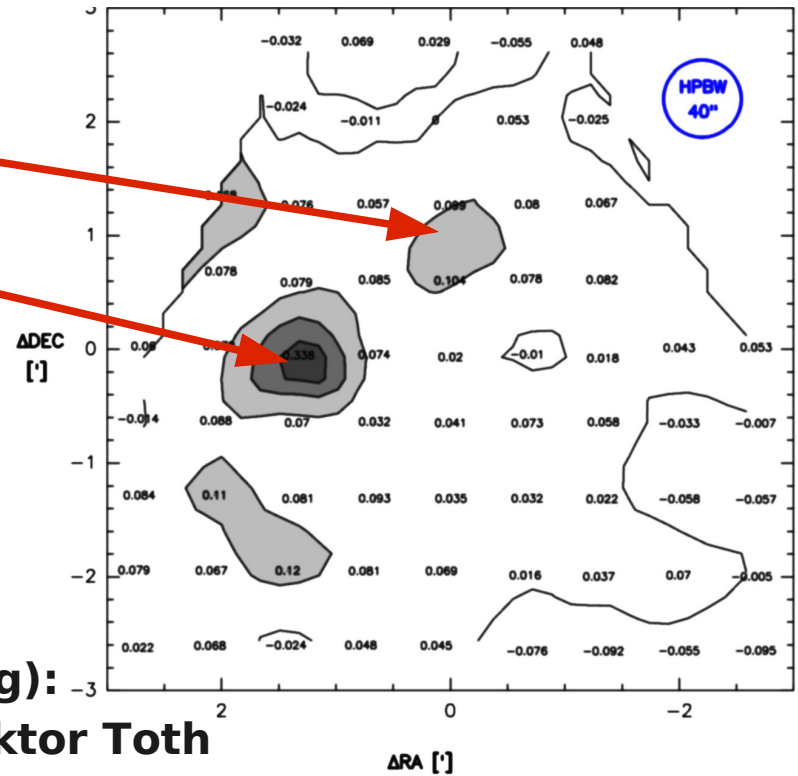


# 3D extinction: Doug Marshall

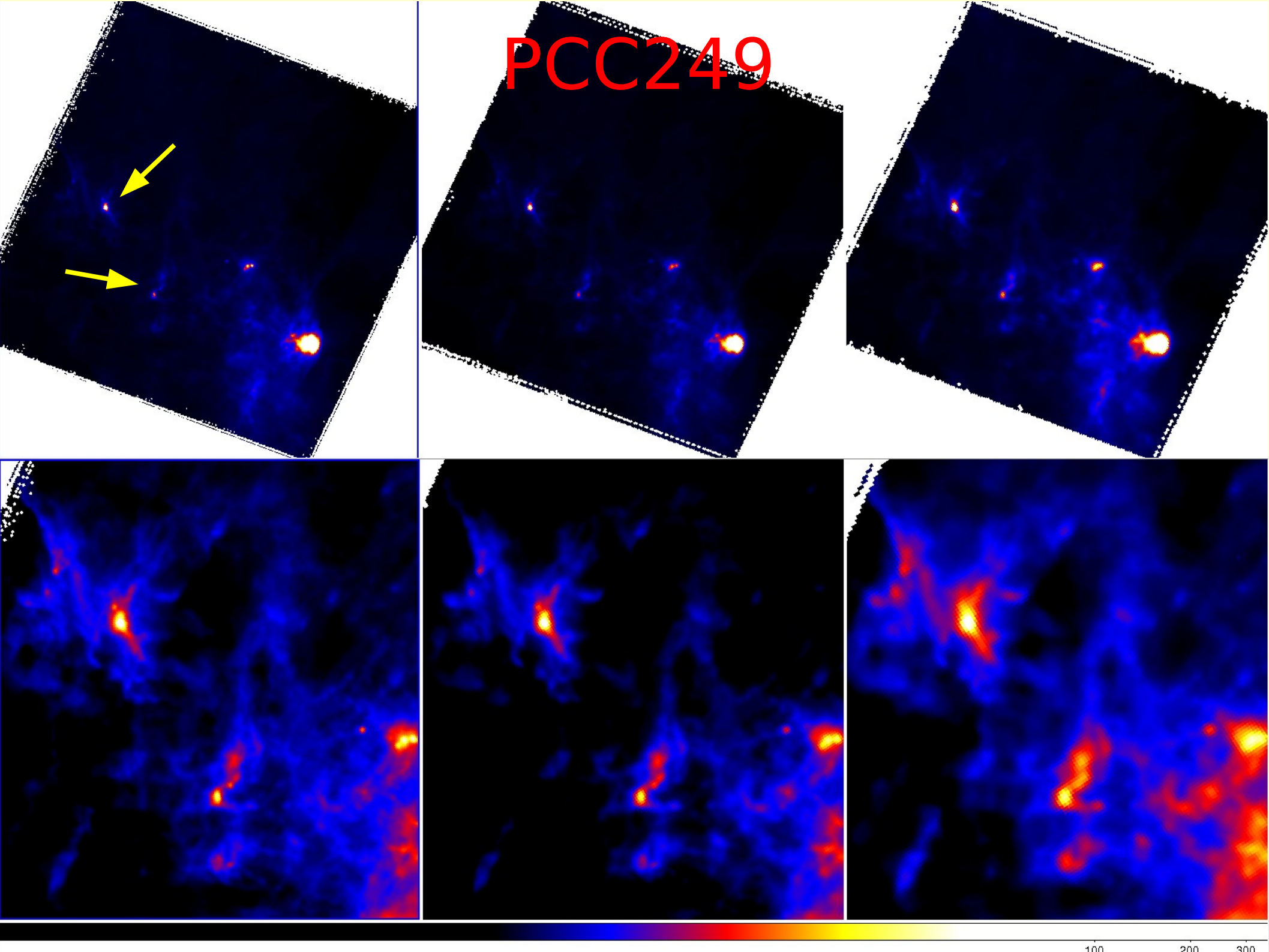


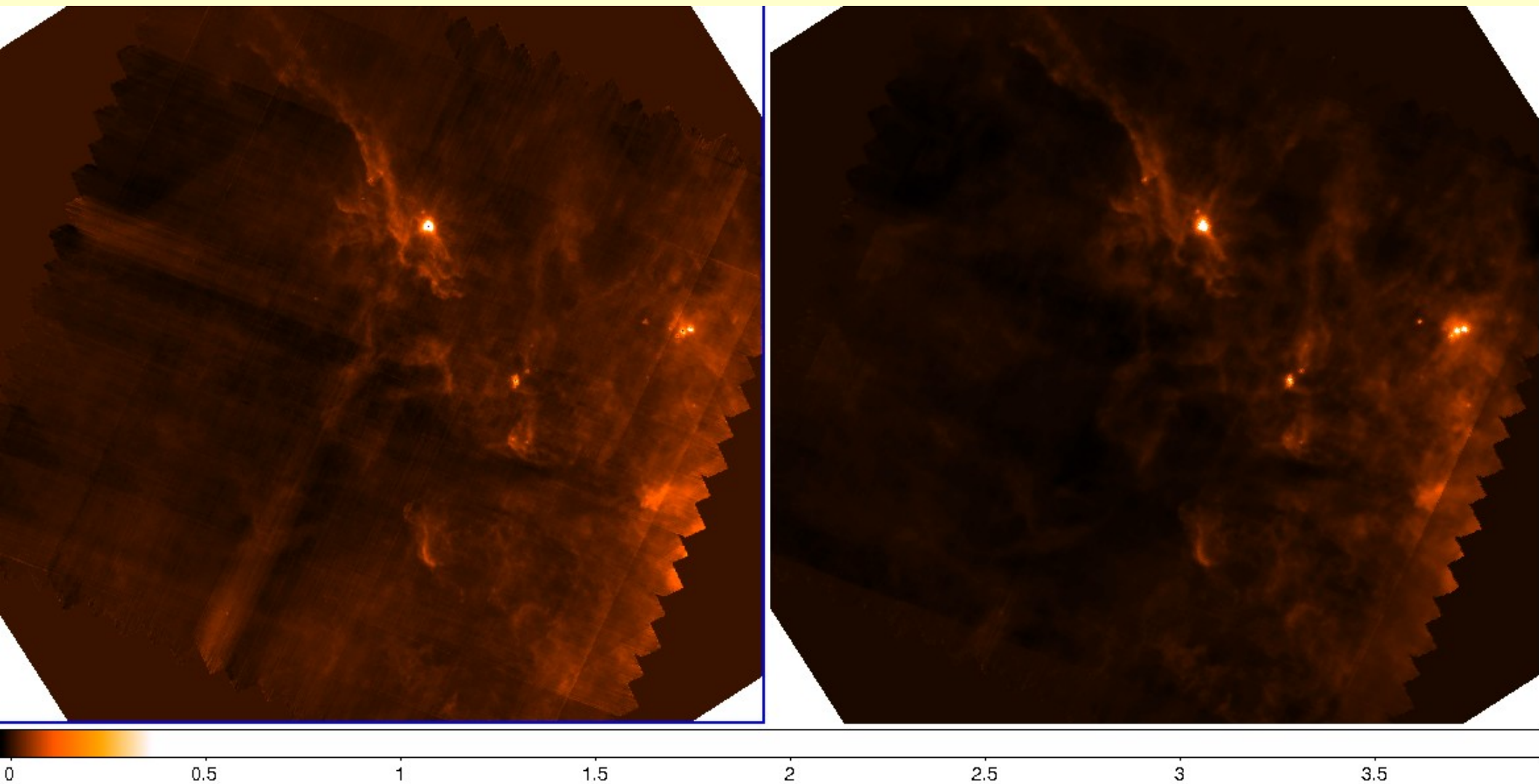
PACS data reduction:  
L. Anderson

NH<sub>3</sub> data (Effelsberg):  
Erika Werebelyi, Viktor Toth



PCC249





## Plan for the future

- follow-up observations and analysis of SDP sources
- publication of a paper in the Herschel special issue
- 2010: analysis of first Planck sky survey
- end of 2010: target list for the actual Herschel survey of Cold Cores

