Exporting HIFI spectra to CLASS

- HiClass exports HIFI spectra from HCSS to CLASS.
- Used a lot recently since we collected interesting data
 - => user feedback.
- HIClass creates a CLASS readable FITS file.
- HiClass is offered as an application for HIFI SpectrumDatasets.
- Some limitations:
 - CLASS expects regular frequency grid
 - Data need to be resampled, stitched and space craft velocity corrected before export.







Standing Wave Removal

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Level $1 \rightarrow 2$ Processing **Standing Wave Removal: the Problem**

Standing waves from optical and electronic components seen (as expected) in HIFI spectra. Pipeline removes those by using appropriate chopping against sky or load, nodding, or frequency switch. Standing wave residuals might still be seen in Level 2 data. Wave-type is HIFI-band dependent:

•Beamsplitter bands 1, 2, and 5 show sine waves

•Diplexer bands 3 and 4 show sine waves with amplitude increasing to IF band edges

•HEB bands 6 and 7 waves are not sine waves. Requires special treatment, not discussed here.



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Level 1 → 2 Processing Standing Wave Removal: Principles ove residual standing waves (SW) with as little

Steps to remove *residual* standing waves (SW), <u>with as little</u> <u>user-interaction as possible:</u>

1.02 •Separate SW from: ∄ 1.00⁺ other baseline fluctuations... – emission/absorption lines 0.98Ē 0.02 •Fit N sine waves with different 0.00 periods, amplitudes, phases to -0.01 597 598 599 wavelength baseline-subtracted, line-masked spectrum.

•Subtract SW fit from original input data. Caveat: should be division, but often not possible as 'continuum' level uncertain. Not big effect as SW amplitude <few%

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Level 1 → 2 Processing Standing Wave Removal: Principles

- Baseline and mask determined using iterative filtering, smoothing, sigmaclipping algorithm. Needs to know 'typical' longest standing wave period, to separate waves from other baseline structure.
- N sine waves A*sin(x)+B*cos(x)=C*sin(x+phase) fitted to baselinesubtracted, line-masked spectra iteratively:
- A) Determine period, amplitude, phase of wave 1 at minimum chi^2
- B) Determine period wave 2 at minimum chi^2 in sine-wave subtracted data
- C) Solve for amplitude and phase *combined waves* **1+2** using 'LU' matrix decomposition in original input data
- D) Subtract waves 1+2 from original input data
- E) Repeat steps B, C, D for waves 1+2+3, ..., 1+2+3+...+N





Level 1 → 2 Processing Standing Wave Removal: in HIPE

•FitHifiFringe: HIFI standing wave removal tool. Available in HIPE 1.2 and 2.0, but still being improved based on ongoing PV experience. Makes use of core tasks:

- FitFringe: fits sine waves to SpectralSegments (basically, any wavelength [in micron!], flux spectrum). Not instrument specific, i.e. could be adapted for use with PACS and SPIRE spectra.
- SmoothBaseline: used by FitFringe, but can be used in standalone mode. Returns baseline AND line mask. NOT a polynomial baseline fit, but rather a smoothed, clipped version of the spectrum. Use with care!



Level 1 → 2 Processing Standing Wave Removal: in HIPE

Most relevant FitHifiFringe input:

- WBS or HRS SpectrumDataset
- nfringes: number of sine waves to fit
- Plot: 0=not, 1=most important plots
- Averscan: determine SW from average of all scans, and subtract that from all. Might be useful if waves are weak.
- Doglue: determine SW on combined WBS subbands
- Midcycle: typical SW expected in data [MHz] Longer period structures are assumed to be baseline or sky features.
- Cycle: longest period SW to search for [MHz]

Output: sine wave(s)-subtracted SpectrumDataset and list of sine wave parameters fitted

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Level 1 → 2 Processing Standing Wave Removal: in HIPE

FitHifiFringe shows fitted periods, amplitudes, phases in HIPE console and by default produces 2 plots:

•Chi² as function of period. Minima found are indicated with red vertical lines



•Result plot: input data, sine wave fit, baseline, mask, sinewave subtracted spectrum





Level 1 → 2 Processing Standing Wave Removal: Future

- •FitHifiFringe will keep evolving as more PV data becomes avaivable and the 'problem' becomes better constrained
- •HIPE 3.0: manual masking possible and ObsContext in/output
- Suggestions welcome
- •Removing band 6 and 7 waves a different problem, which cannot be solved by fitting sine waves. Alternative methods are being worked on.



Million Adjusting the HIFI Pipeline

- The most appropriate level to make changes in the HIFI pipeline is processing from Level 1 to Level 2. At this level is may be necessary to:
 - Resample onto user specified grid.
 - Stitch subbands together
 - Correct for standing waves
 - Remake spectral cubes
 - Run a spectral survey through the deconvolution tool
 - Export to CLASS
 - Etc...
- The pipeline allows to introduce these changes my altering the Level2PipelineAlgo.py
 - The file is found in

<hipe directory>/scripts/hifi/pipeline/generic/Level2PipelineAlgo.py



How to setup HifiPipelineTask

- HifiPipelineTask is the jython task which controls the HIFI pipeline.
 - This task is activated by an observation context present in the Hipe session.
 - For the Hipe 2.0 RC3 version use the following:
 - Change to Expert Mode
 - Turn off palStore option
 - Turn off Quality plugin.
 - Type in: Configuration.setProoperty("hcss.ia.pal.store.spgstore","{dummy}")
 - Select run levels (for this example, I set fromLevel to 1.0 and upTo level to 2.0
 - Press "Accept" to rerun the pipeline.





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Altering and applying an updated Algorithm

- Copy the Level2PipelineAlgo.py to your own area and rename the file.
- Edit the algorithm file.
 - In my example I just comment out the doAvg task. And add a line for the pipeline which reflects that I've done this.
- In the leve2Algo input window for the task dialog, add the new file name for the algorithm you have just edited.
- Set fromLevel to 1.0 and upToLevel to 2.0 in the dialog box.
- Run the pipeline by pressing "Accept" in the dialog box



HIPE 2.0 - Level2Pi...Algo.py







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• Running the pipeline again using this algorithm, shows that final spectra are not averaged.

