

# Locating dust crystals in protoplanetary disks with PACS

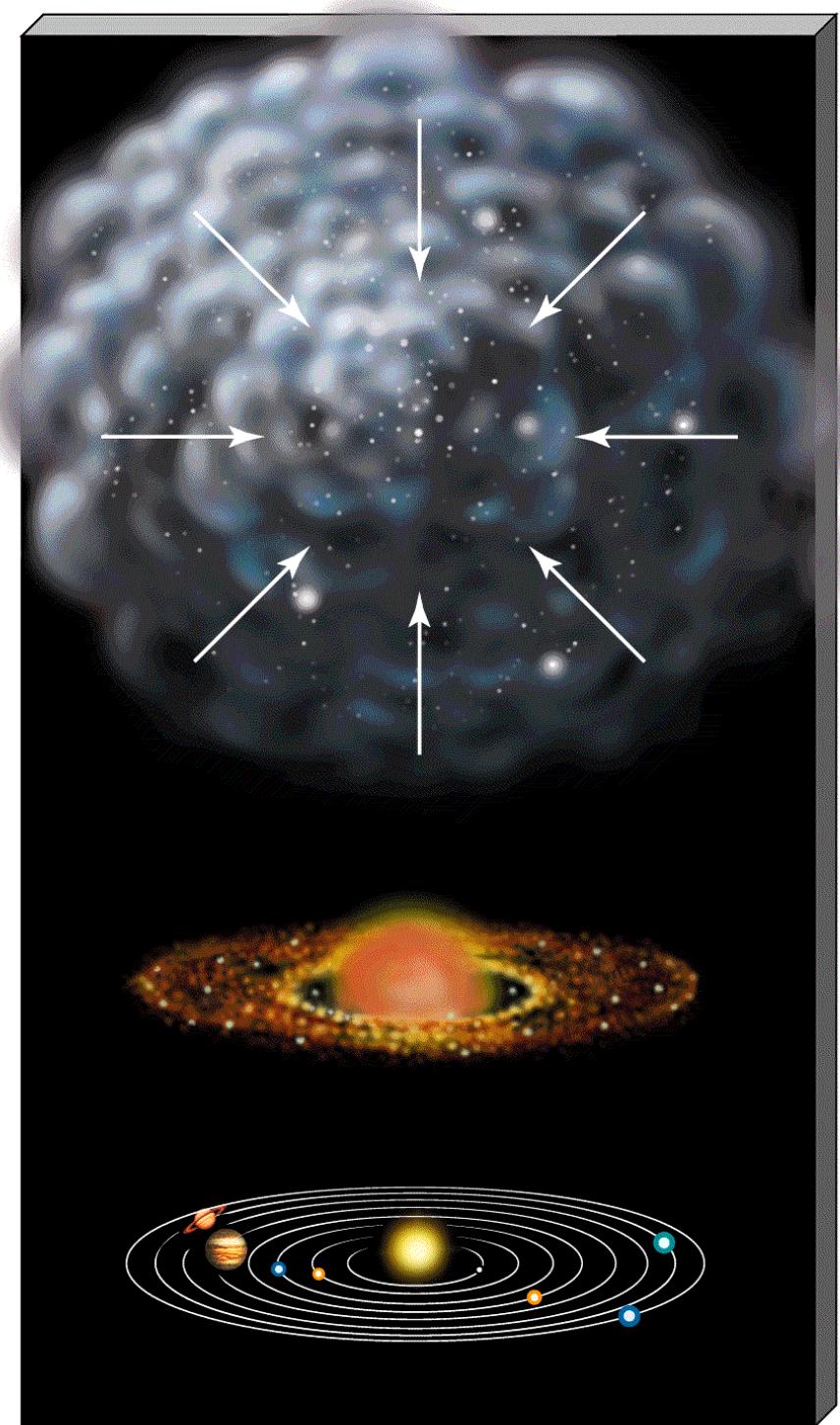
Gijs Mulders

University of Amsterdam

DIGIT: Dust, Ice and Gas In Time



# Mineralogy



# Mineralogy



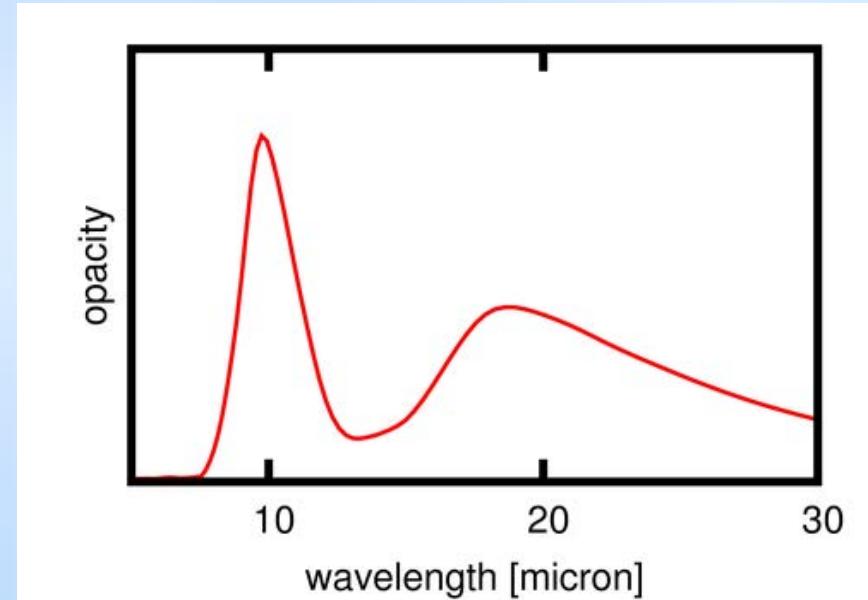
>1000K



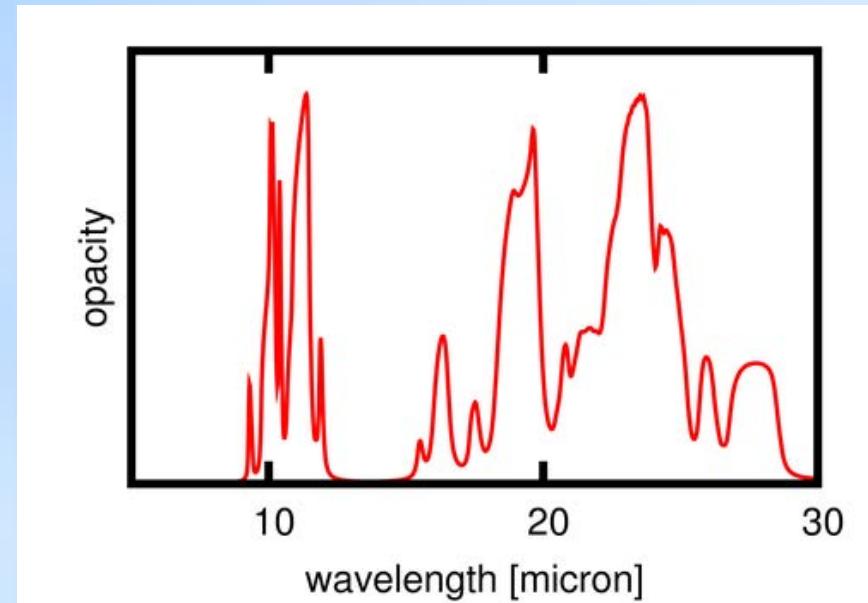
# Mineralogy



>1000K

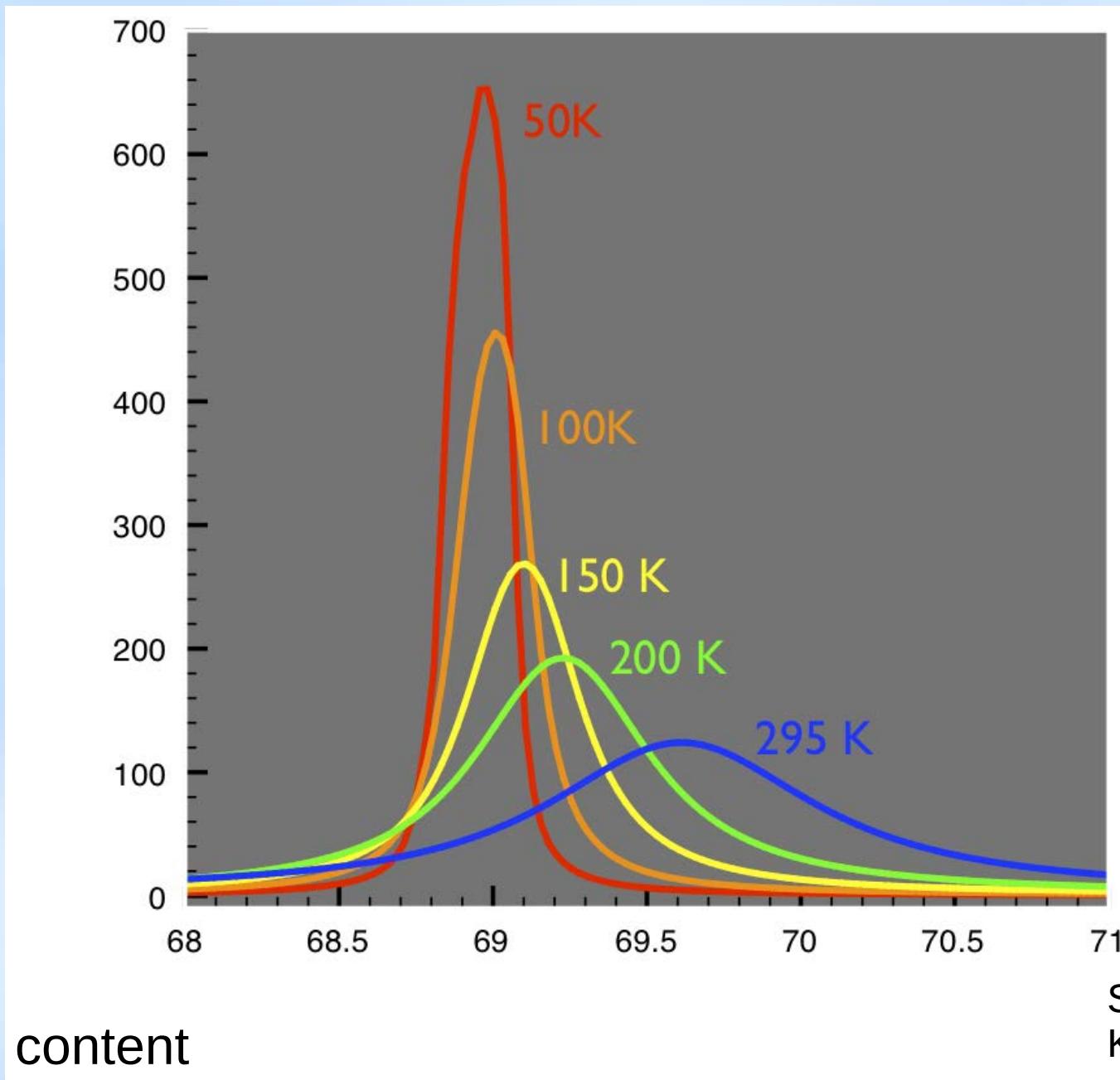


wavelength [micron]



wavelength [micron]

# Forsterite: 69 $\mu\text{m}$ feature



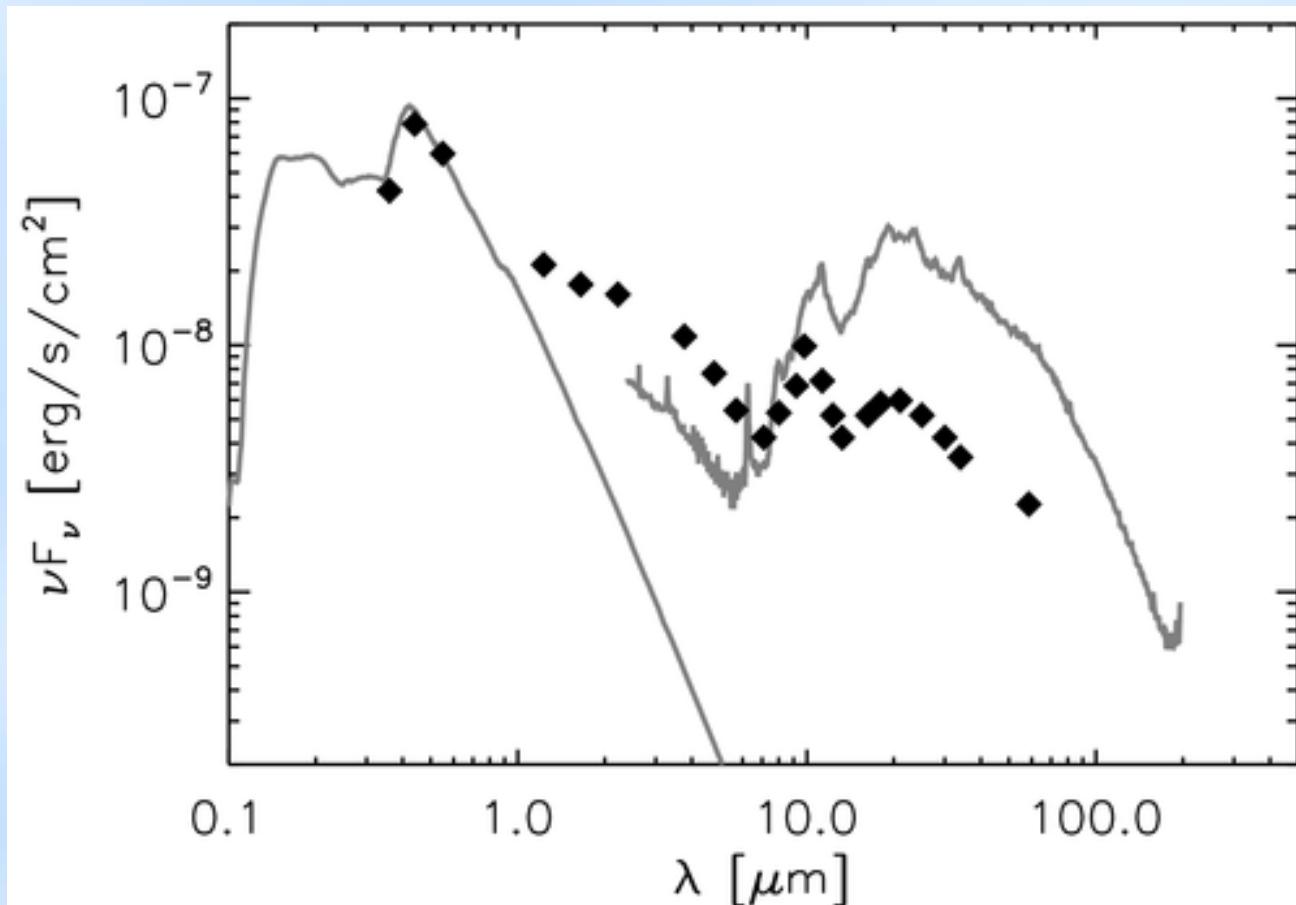
Caveat: iron content

Suto et al. 2006  
Koike et al. 2003



- 69  $\mu\text{m}$ : 6 detections  
(Sturm et al. in prep)
- HD 100546  
(Mulders et al. 2011)
- HD 169142  
IRS 48  
(Maaskant et al. in prep)

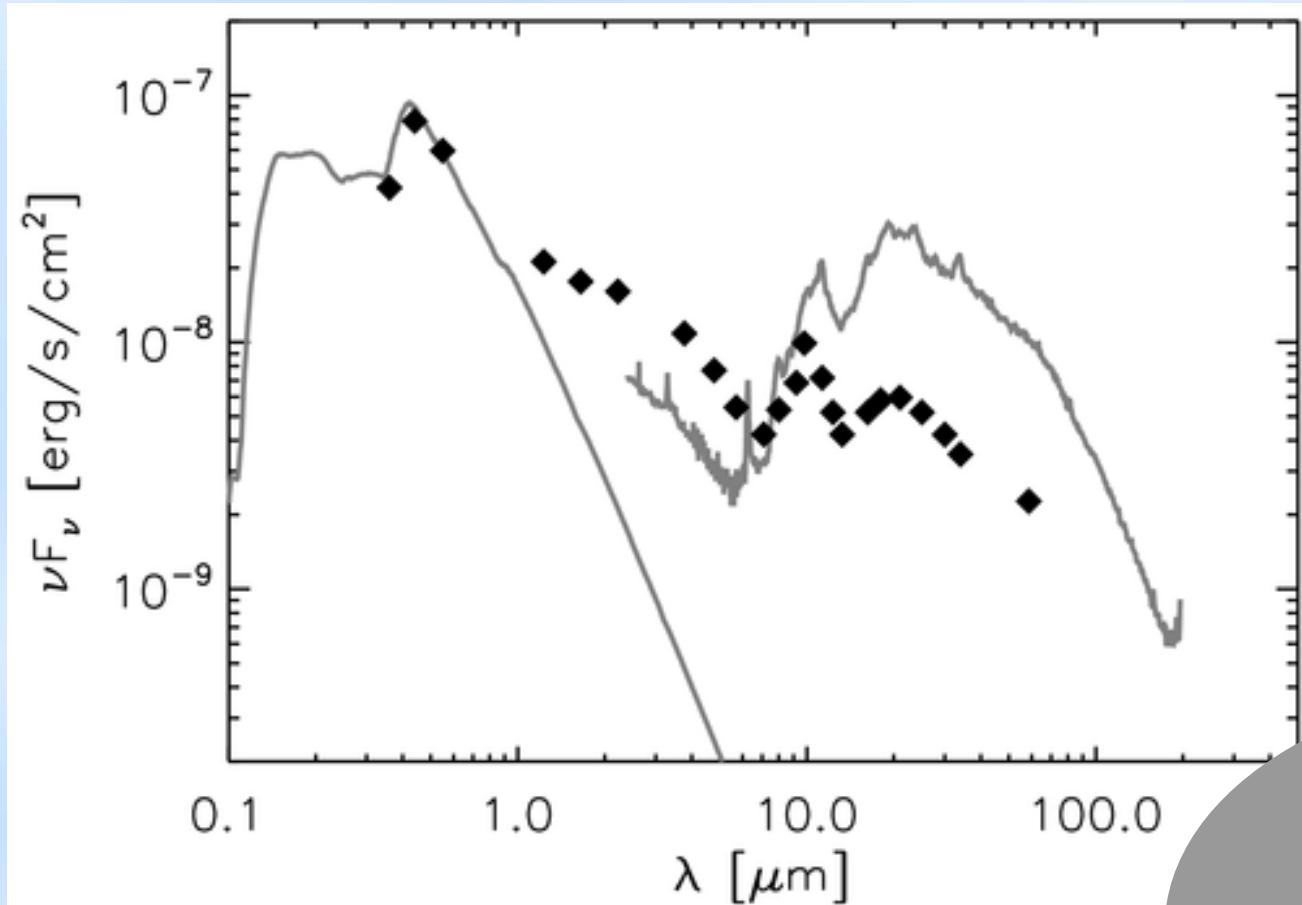
# HD 100546



Malfait et al. 1998

Mulders & Dominik 2012

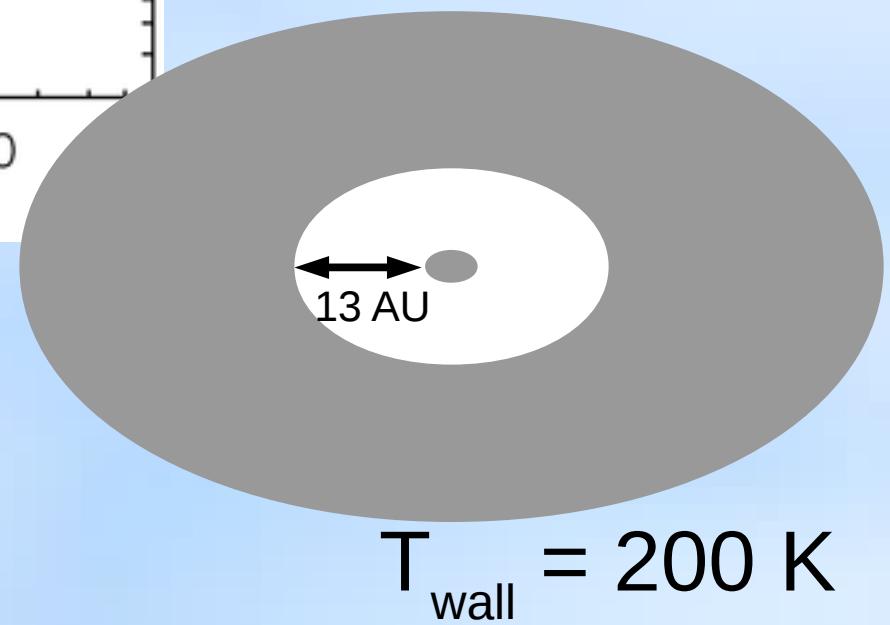
# HD 100546



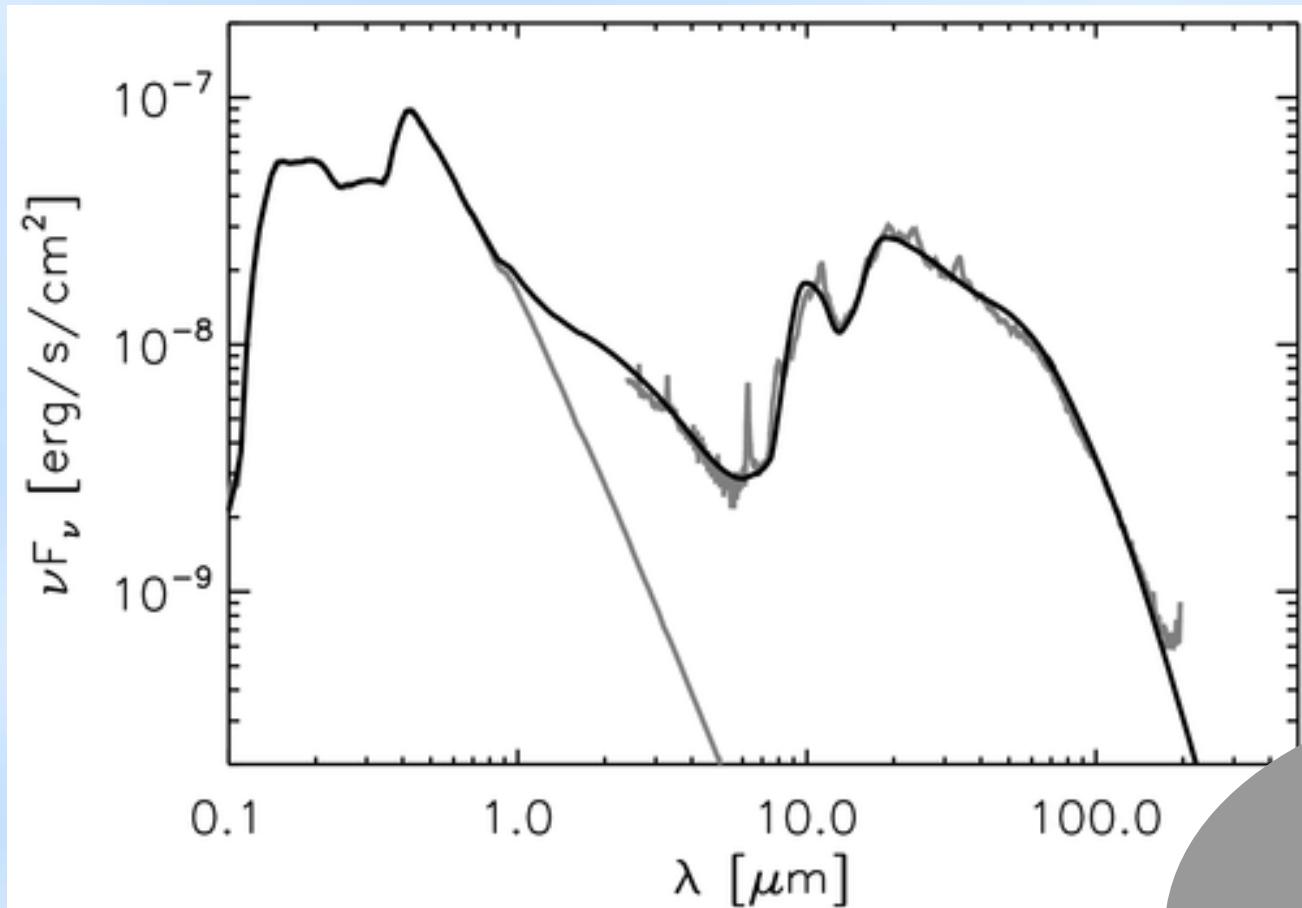
Malfait et al. 1998

Mulders & Dominik 2012

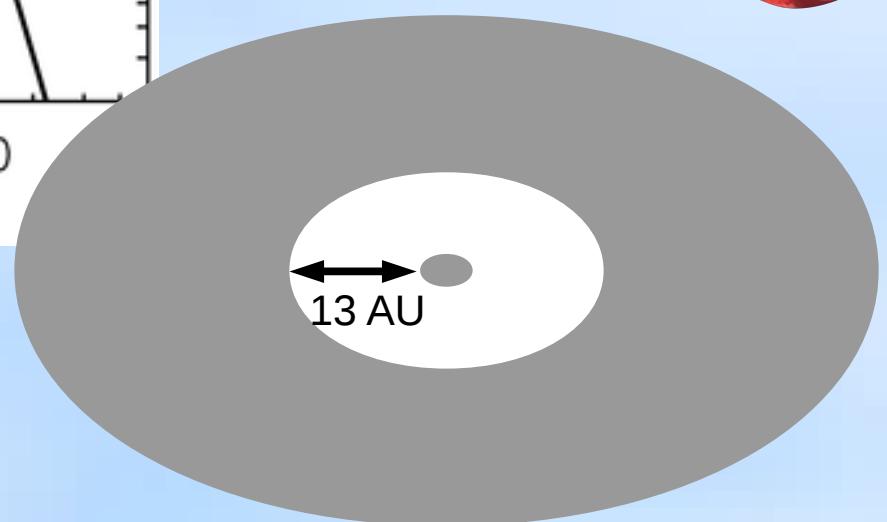
Bouwman et al. 2003  
Grady et al. 2007  
Benisty et al. 2010



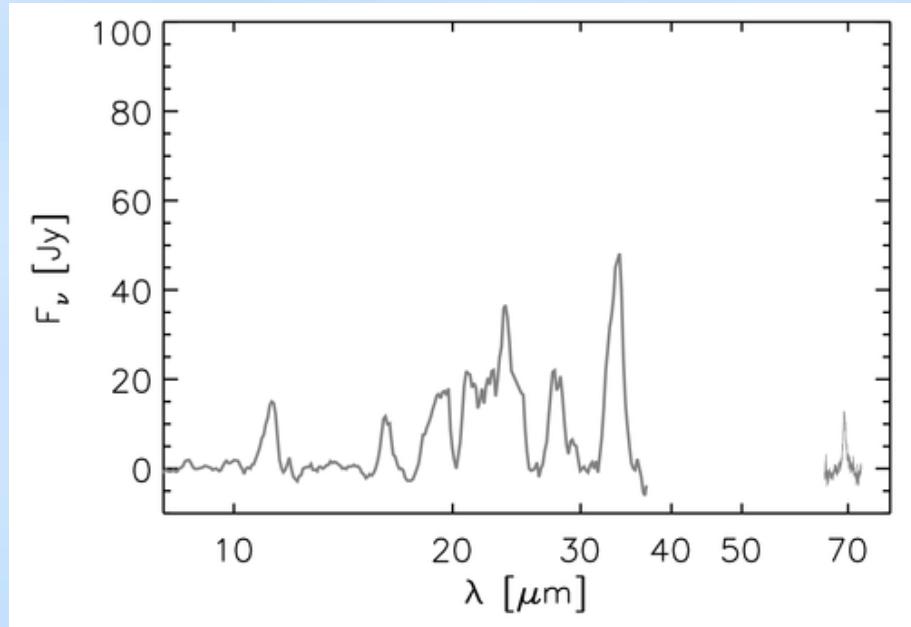
# HD 100546



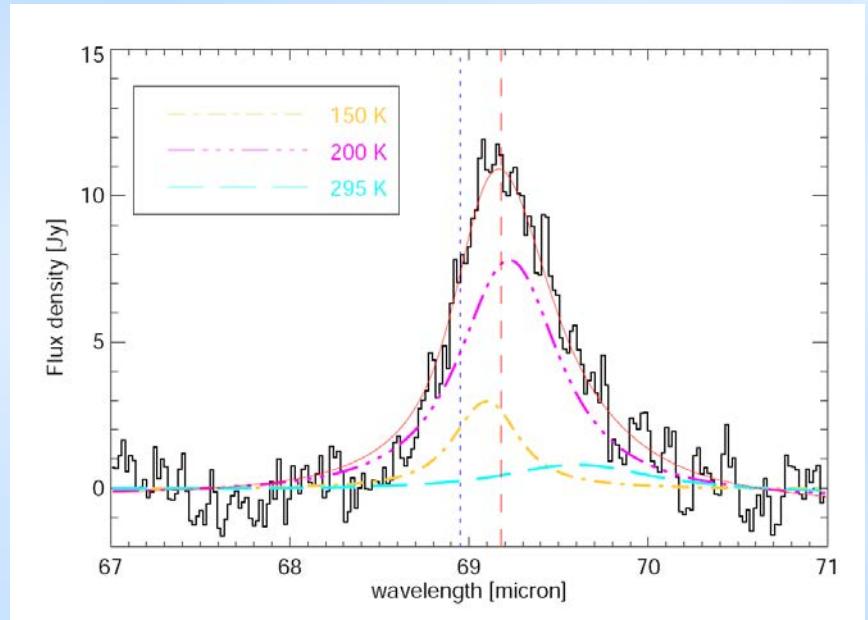
Mulders et al. 2011



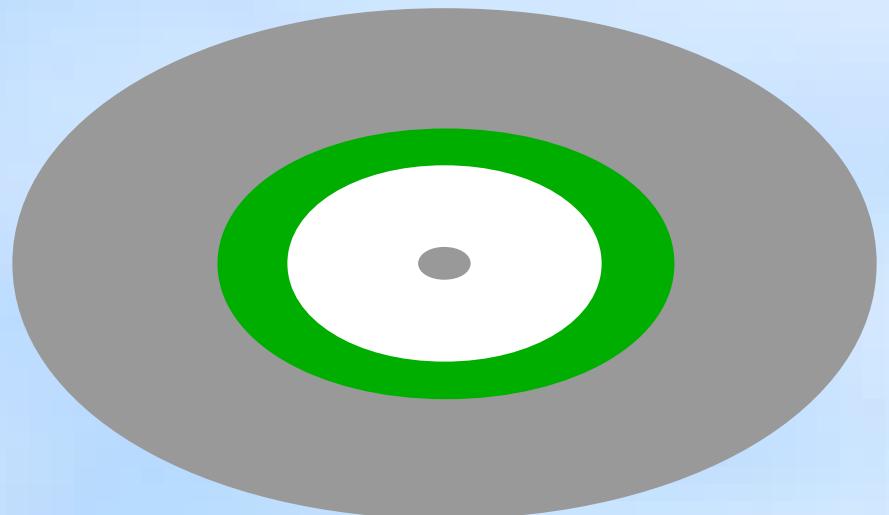
$$T_{\text{wall}} = 200 \text{ K}$$

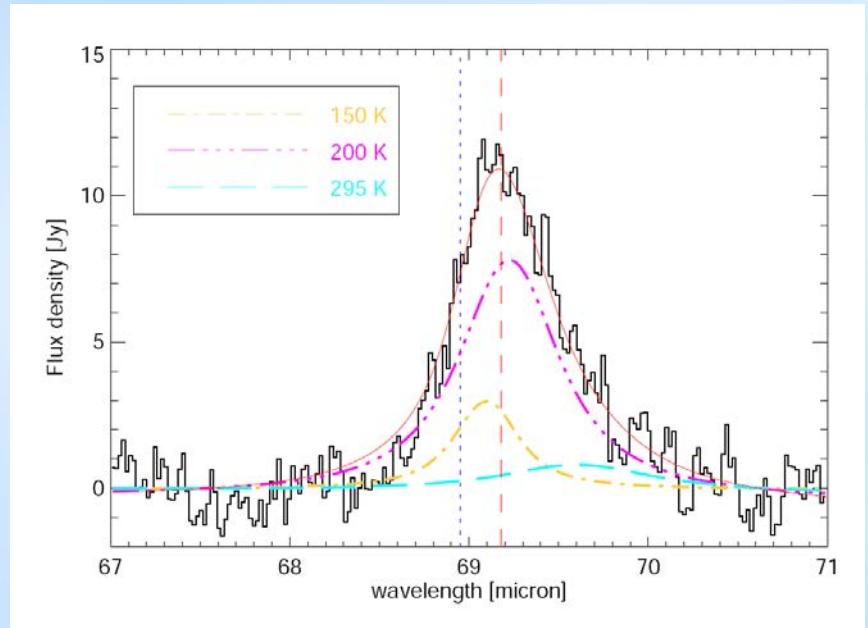
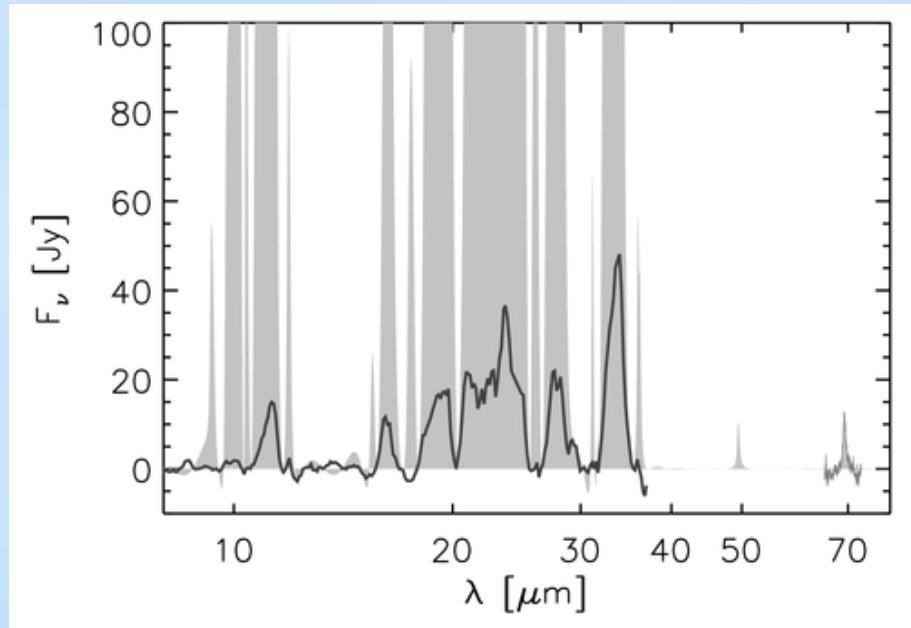


Feature strength



Sturm et al. 2010  
69  $\mu\text{m}$  shape



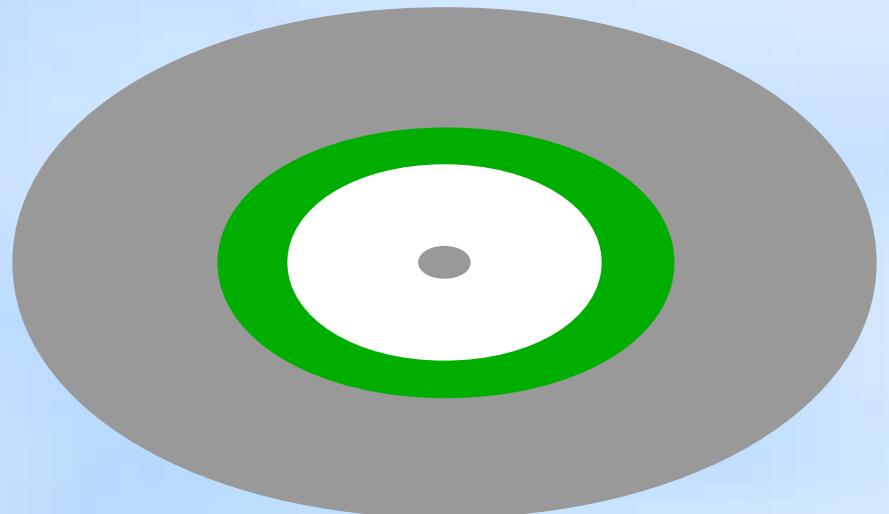


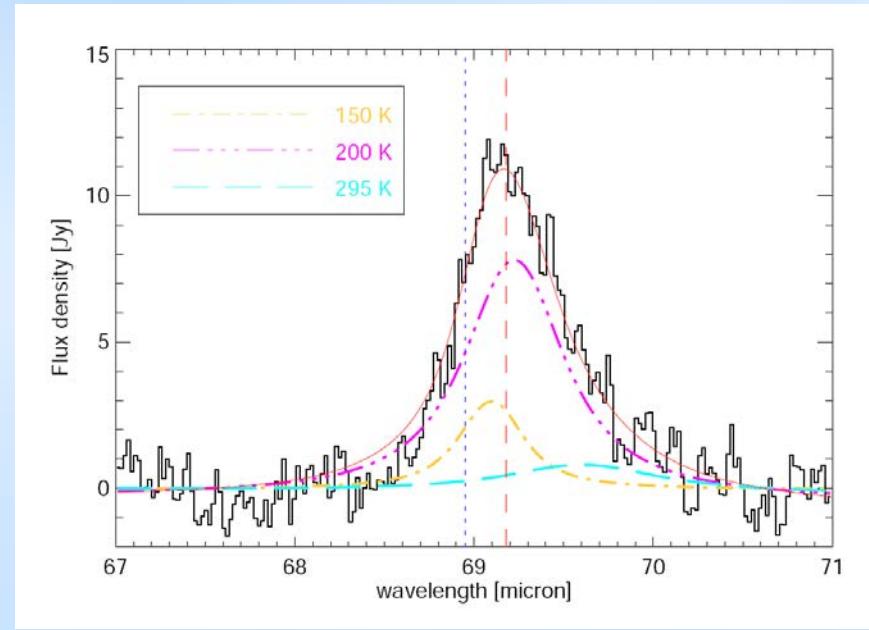
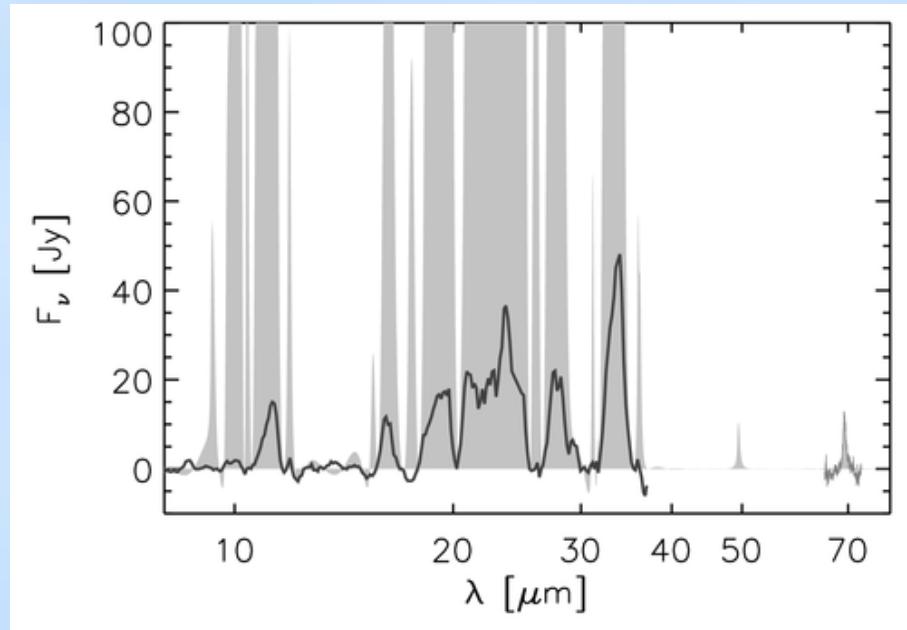
Sturm et al. 2010

## Feature strength

69  $\mu\text{m}$  shape

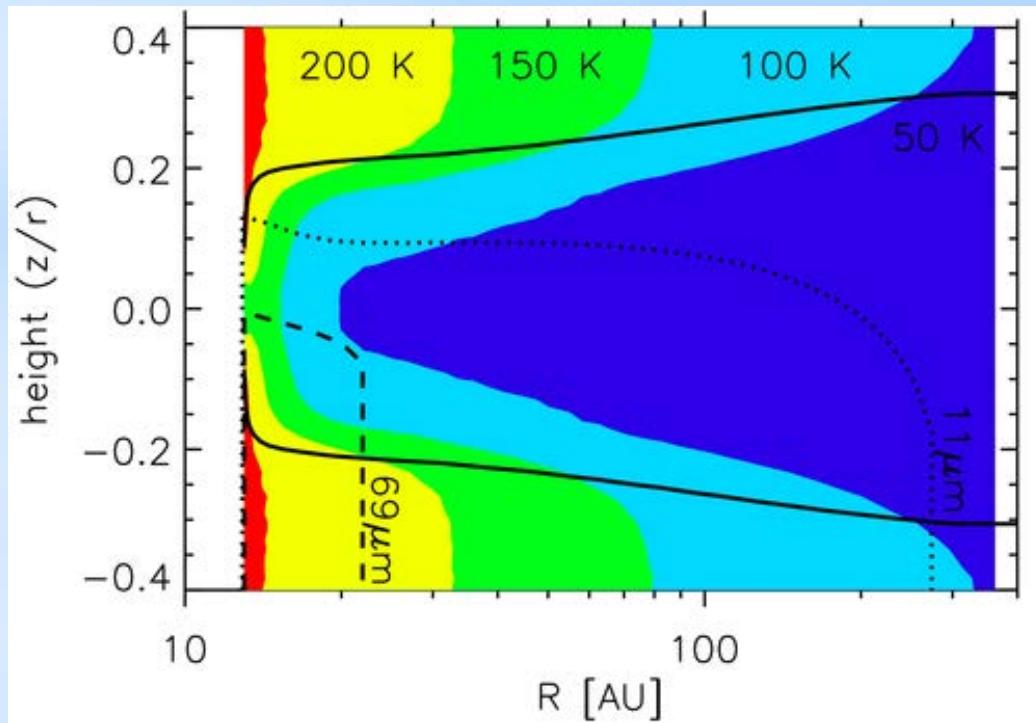
- Iron (few %)
- Optical depth effects



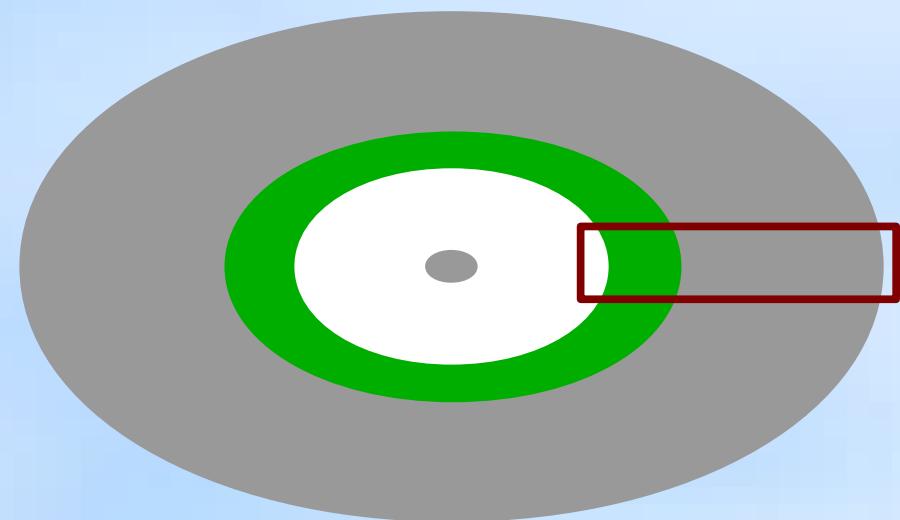


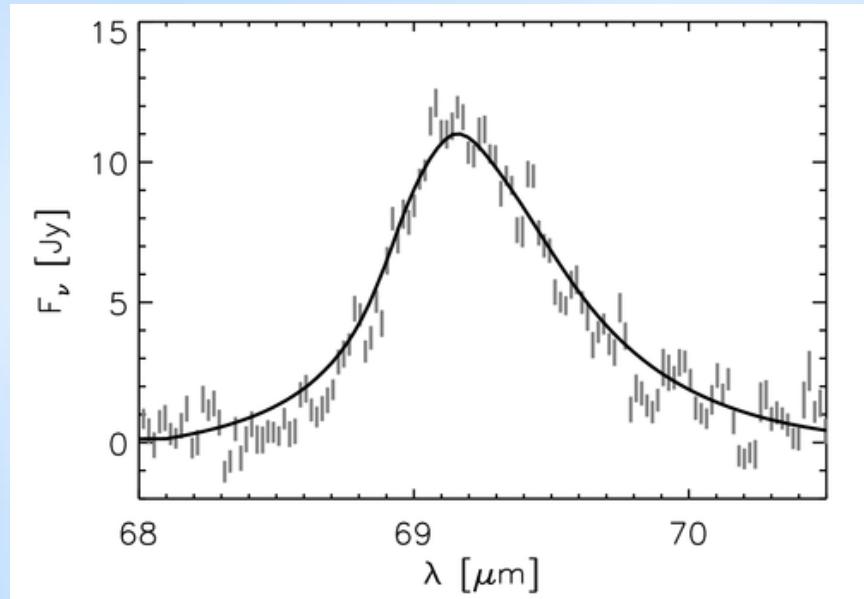
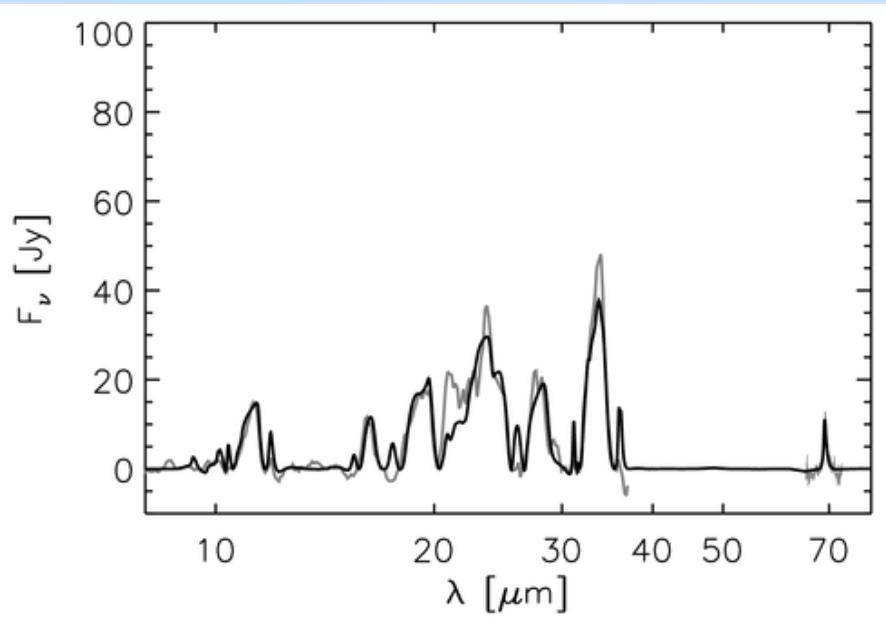
Sturm et al. 2010

## Feature strength

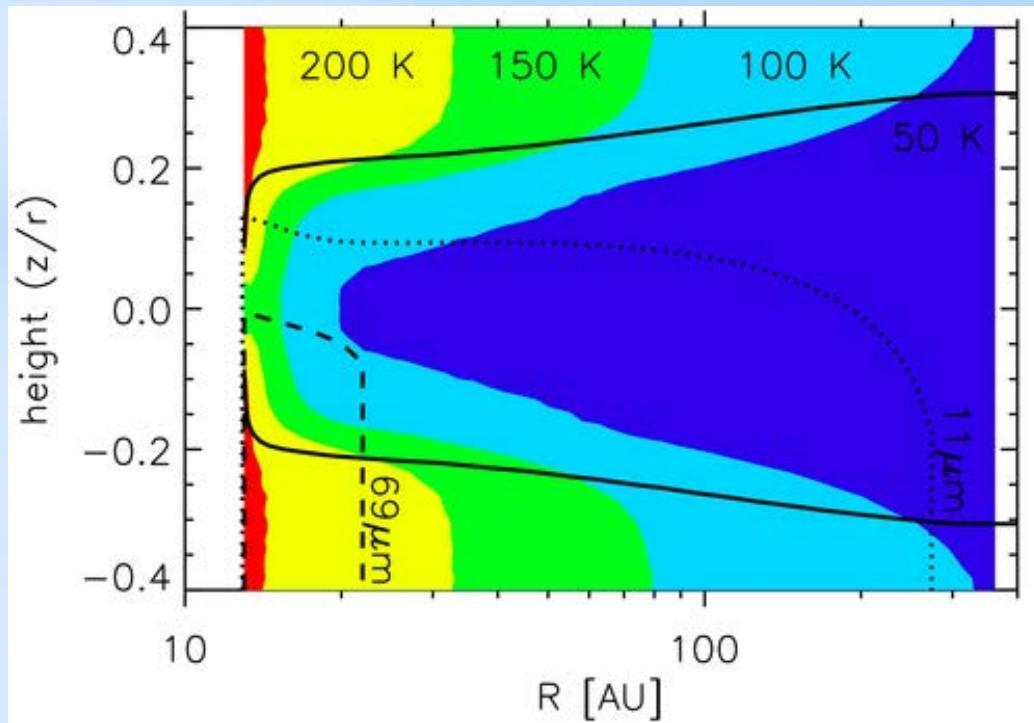


## 69 \mu m shape

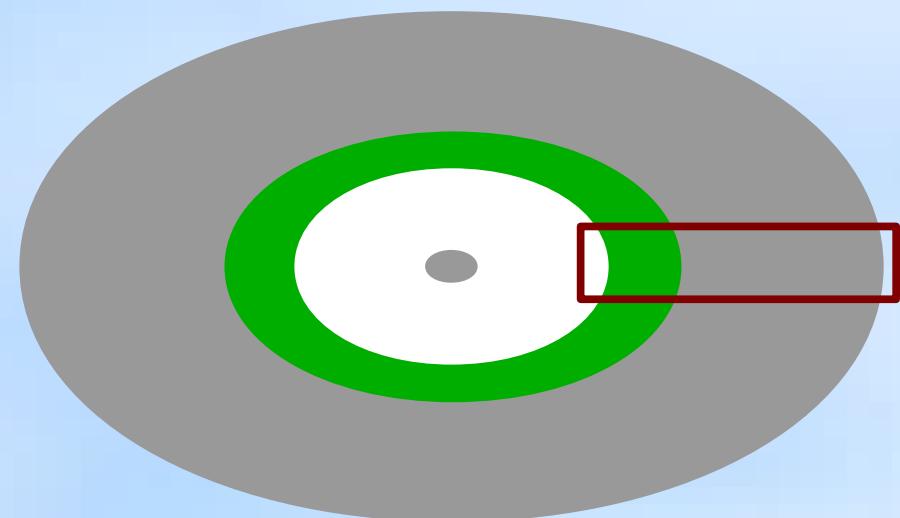




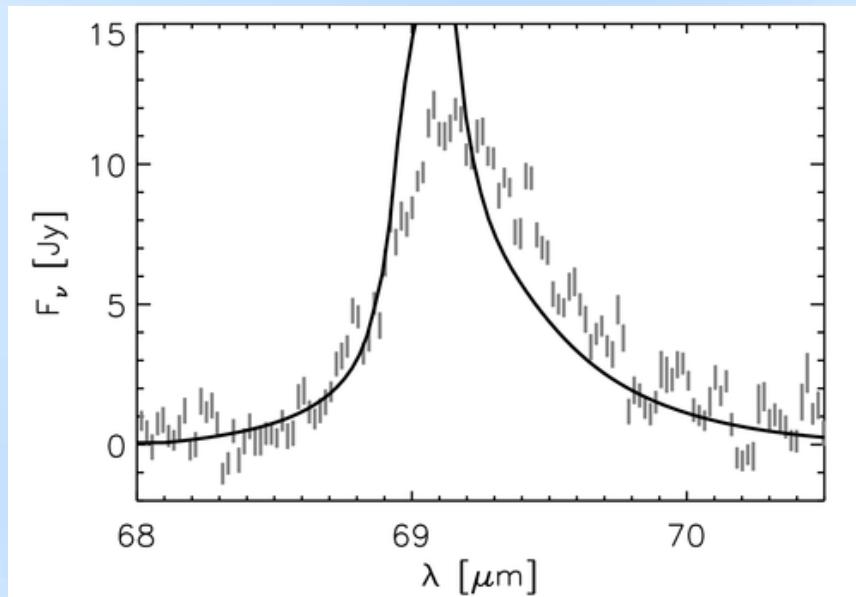
## Feature strength



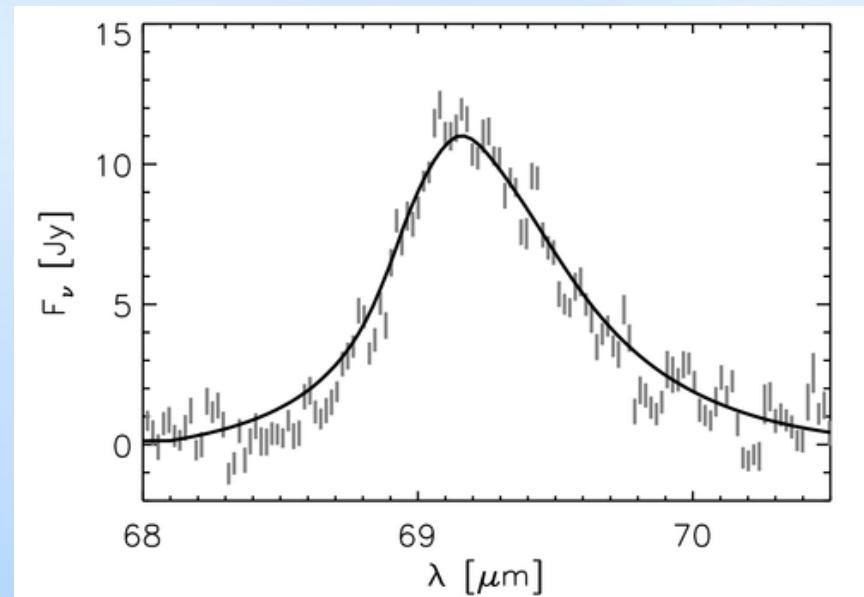
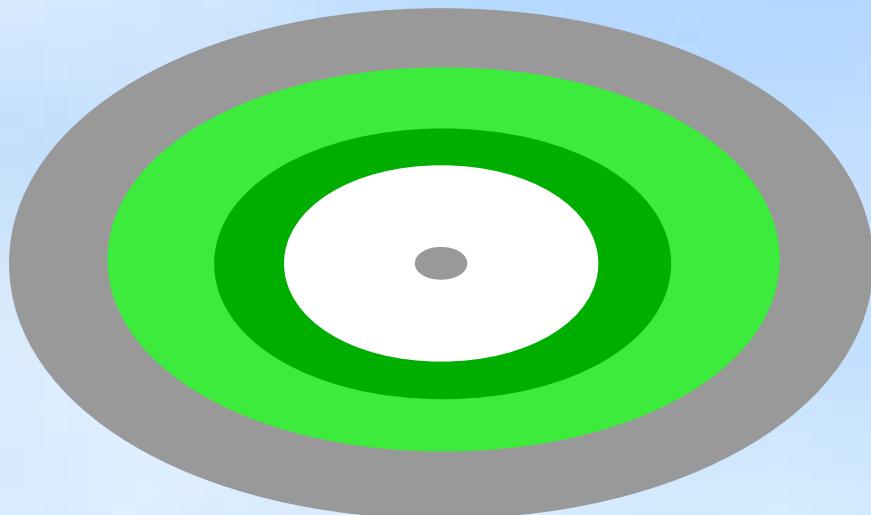
## 69 $\mu\text{m}$ shape



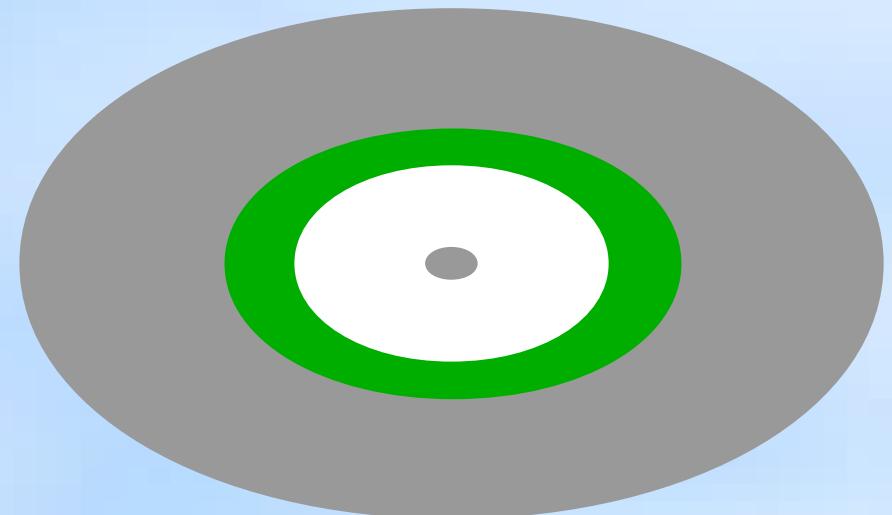
# Forsterite distribution



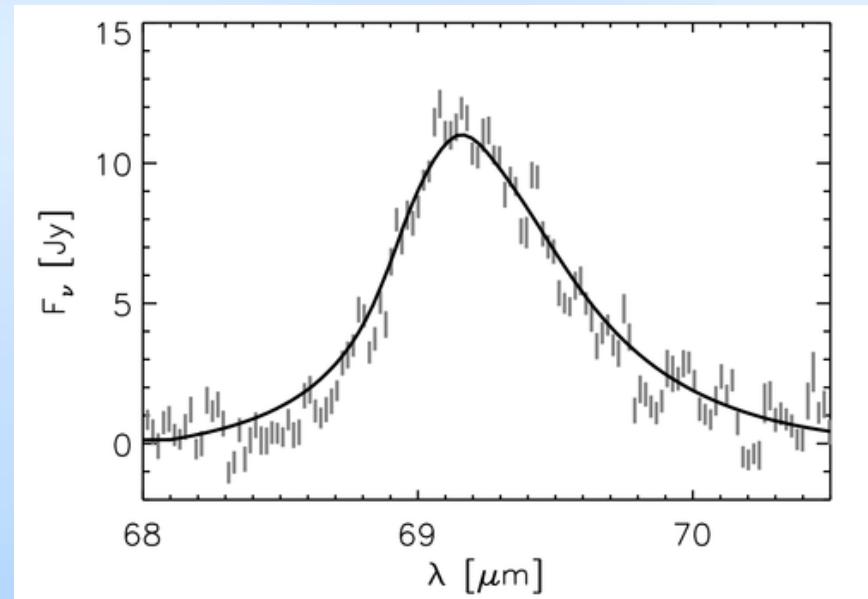
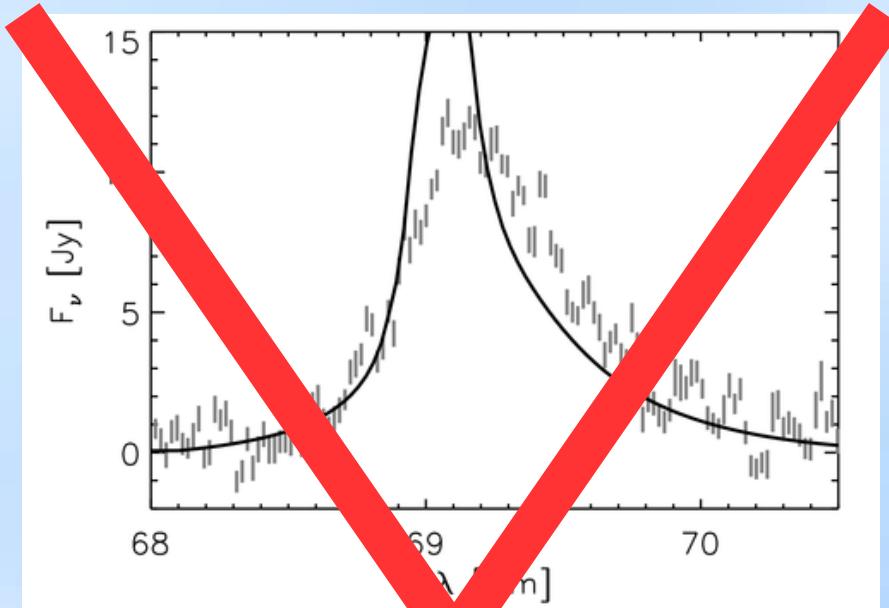
gradient



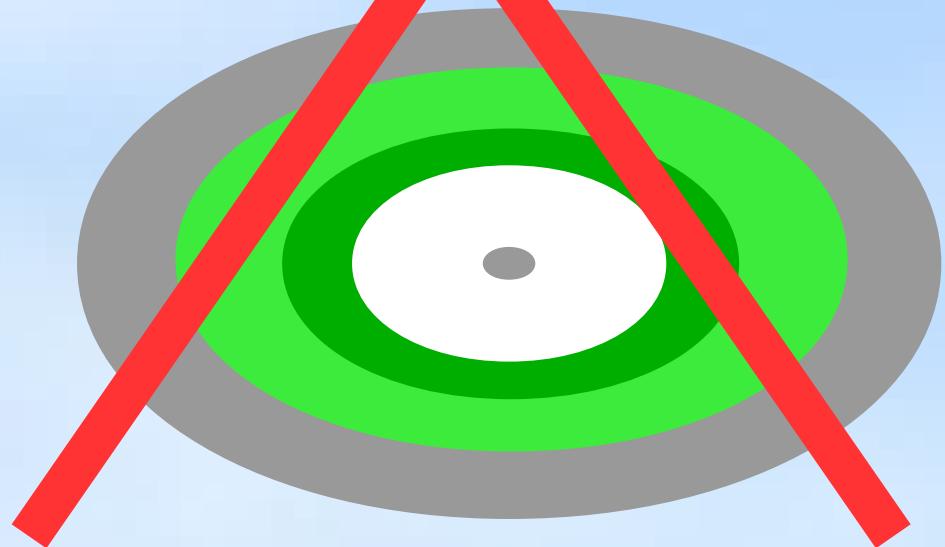
ring



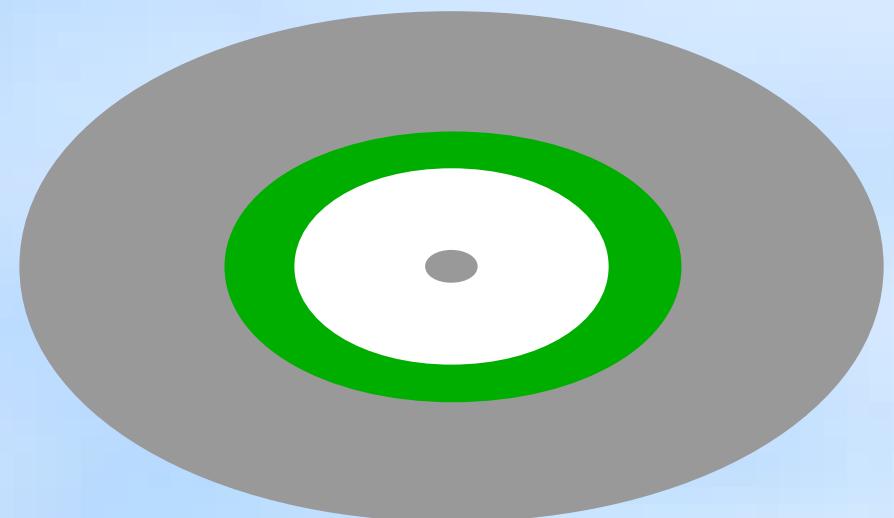
# Forsterite distribution



gradient



40%, 13-20 AU



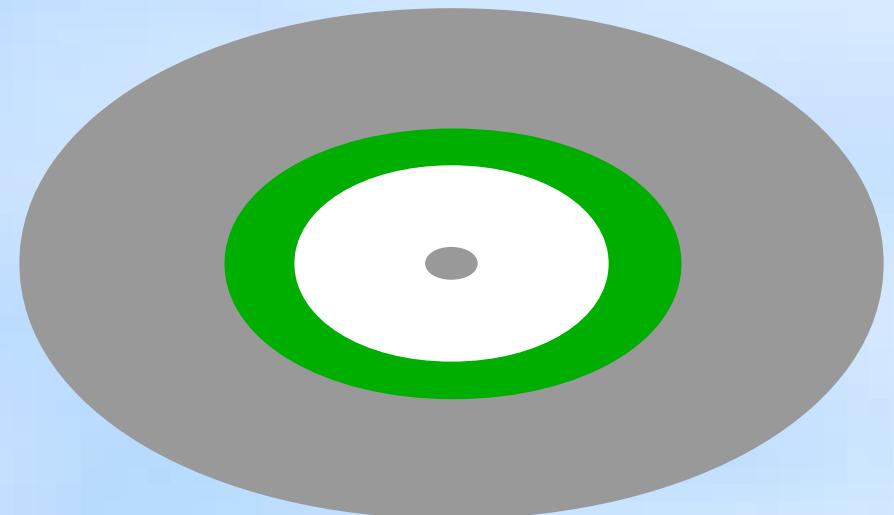
iron <0.3%

# Crystal origin

- Shocks
- Collisions
- Parent-body processing



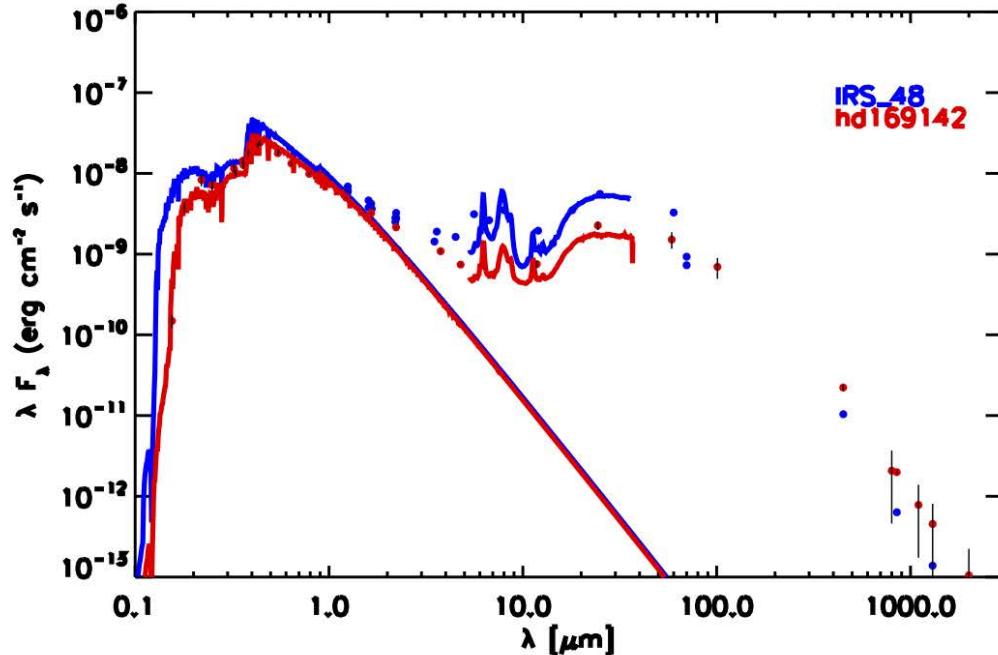
Bouwman et al. 2003



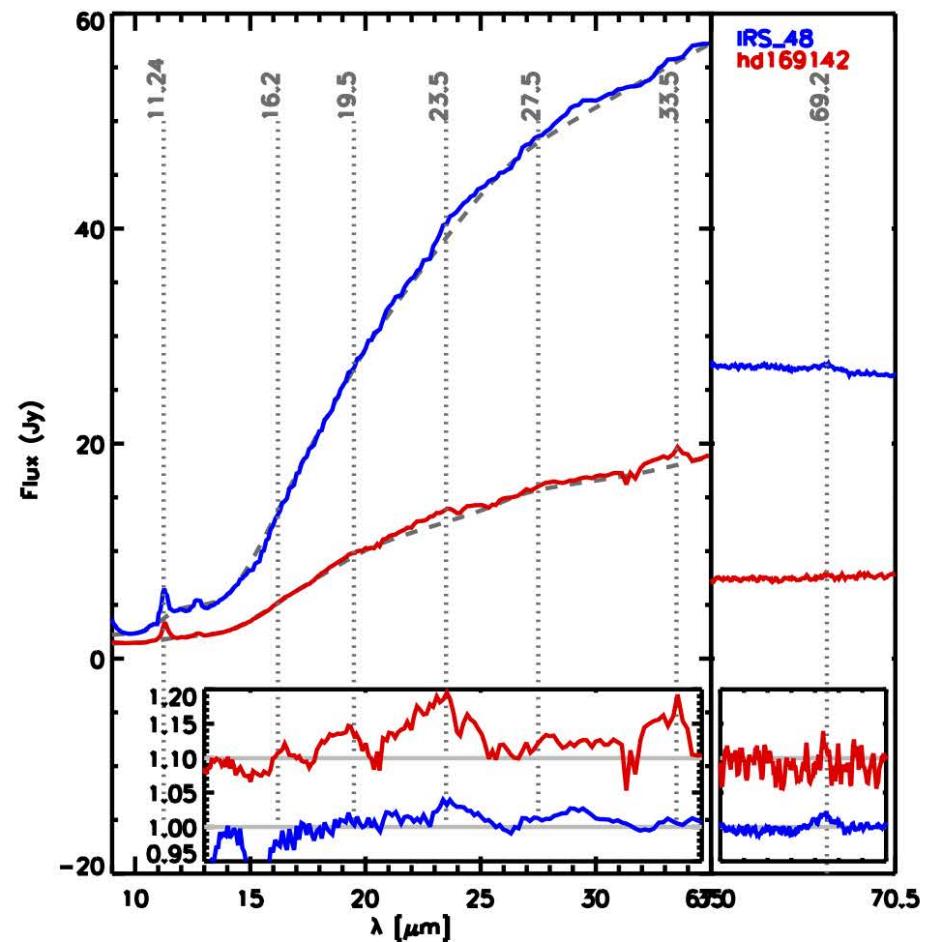
# Comparison of forsterite in:

**HDI169142** gap 23 AU *(Honda et al submitted)*

**IRS 48** gap >50 AU *(Geers et al 2007)*

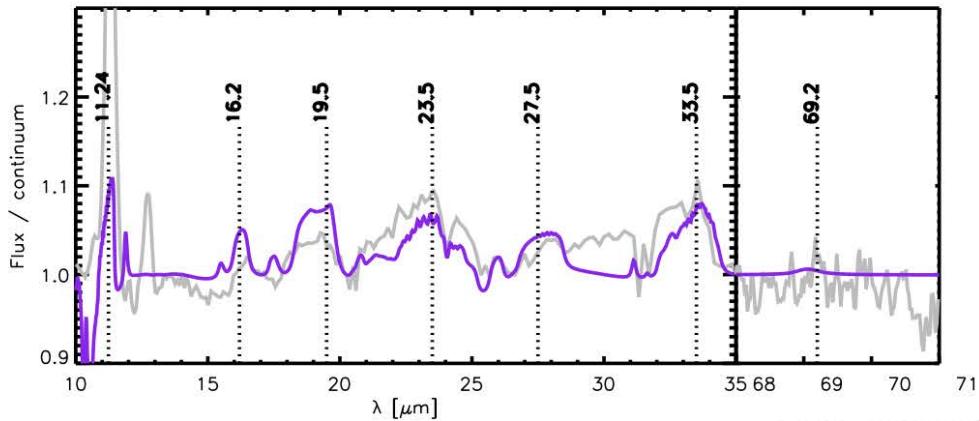


Maaskant et al in prep

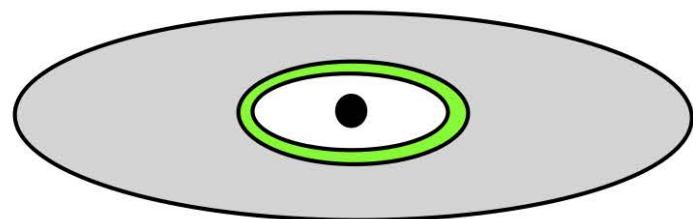


# Radiative transfer solutions...

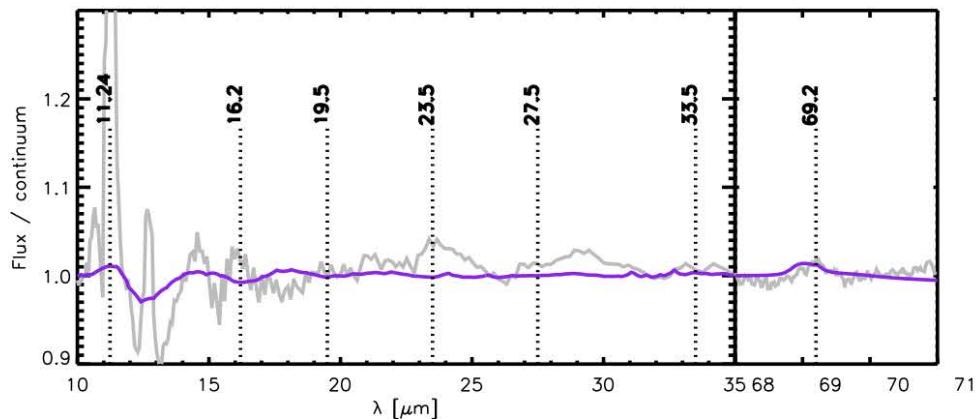
**HD169142**



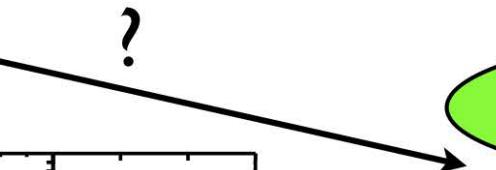
wall



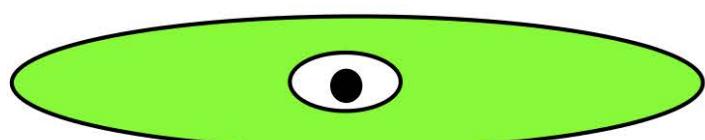
**IRS 48**



outer disk



large >10 micron grains  
whole disk



Maaskant et al in prep

# Locating dust crystals in protoplanetary disks with PACS

- Temperature dependent 69  $\mu\text{m}$  feature
- 2D radiative transfer
- Near disk gap (2x), unrelated (1x)
- Local production  
(Shocks, collisions, **parent-body processing**)
- Iron poor (<0.3%)