Far-IR Photometric Monitoring of Protostars in Orion Molecular Clouds

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Protostellar Variability

- Variability is a defining characteristic of Young Stellar Objects (Joy 1945)
- Optical / Near-IR: accretion shocks hot spots (e.g. Carpenter et al 2001)
- *Mid-IR*: Inner disk structure (warped disks, variable scale height)
 - YSOVAR (Morales-Calderón et al. 2009), LRLL 31 (Flaherty et al. 2011), etc.
- *Far-IR (>50µm)*: Only sparse and rare time series yet observed...
 - KAO: Harvey et al. (1998)
 - ISOPHOT: Juhász et al. (2007), Kóspál et al. (2007)





70 µm flux good proxy to protostellar total luminosity (Dunham et al. 2008)

Far-IR photometric monitoring

Herschel AORs overlay



IRAC 4.5 µm + Class I/0 From YSOVAR

- Herschel PACS/Photometer 70 + 160 µm (10h obs. time)
- 18 epochs through 4 visibility windows (first 6 epochs are observed & analyzed)
 Sampling ~week ~months ~1.5 years
- Large maps (35'x35') preferred to pointed observations Covering OMC-1 to 3 which contain 100+ Spitzer-identified Class I/0 protostars
- High redundancy of scan directions
 Secular spacecraft rotation + orthogonal consecutive scans



PACS 160 µm PACS 70 µm Spitzer 8/24 µm

6 epochs combined with Scanamorphos (Roussel et al., sub.)



Source Photometry

- Search for point sources in the 70 µm 6-epoch combined map
 - Automated source-finding algorithms gave unsatisfactory results
 - ➔ Visual inspection
- Photometry measured on individual single-epoch maps
 - Small apertures encircling the PSF core (4'/8' radius @ 70/160 μm)
 - Aperture corrections applied (based on measured PACS PSF)
- Results:
 - 43 point-like sources at 70 μm with flux 0.4 500 Jy
 - 34 have counterparts at 160 µm
 - 13 are point like with fluxes 15 300 Jy
 - 21 are resolved/extended sources
- Most sources are Class I/0 protostars
 - (Morales-Calderón et al. 2009, following presentation by T. Megeath)

Building Reliable Light Curves

Photometric errors on single-epoch maps

- Standard deviation in background annulus inappropriate
- Flux uncertainty estimated from varying aperture radii
 - → ~5% uncertainty at 70 µm (up to 20% in crowded/ structured regions), ~20+% uncertainty at 160 µm

Relative uncertainties:

- Same approach with varying aperture radii
- Light curve deemed unreliable if any single-epoch flux varies by > 5% for various aperture radii
- 17 sources have reliable light curves
 - Variability detection: peak-to-peak variations > 10%







- **20%** variation @ 70 μm over 6 weeks
- Phased light curve @ 160 µm
- High S/N detection (relatively isolated source with smooth background)







- **20%** variation @ 70 µm over 6 weeks
- Phased light curve @ 160 µm
- High S/N detection (relatively isolated source with smooth background)
- Mid-IR variability, 30% (YSOVAR)
- Accretion rate ~ 3.10^{-4} M_{\odot}/yr





- 20+% variation @ 70 μm
- Same trend at 160 µm, but noisy photometry





8 protostars out of 17 show 10+% variability at 70 µm

Inconclusive variability detections...



Inconclusive variability detections...



Future Work...

- SED modeling of variable sources
- Extended emission, PDR (data are public, volunteers?)
- Improved photometry with PSF fitting
- 12 epochs remain to be observed/analyzed
- PACS spectrometer monitoring just started on selected sources





More Herschel projects to monitor YSOs:

GT2_zbalog_2, OT2_jforbric_3, OT2_nbillot_2, OT2_pabraham_4, OT2_rvisser_1, OT2_zbalog_3

Data reduction

- HIPE Standard pipeline up to calibrated data cubes (Level1 frames) including non-linearity corrections
- Single-epoch maps (single-direction scan map) are produced with the standard masked high-pass filtering scheme + projection



Same map (PACS 70), different color scale stretch

- Large spatial scales are affected to some extent by the high-pass filter
- Small spatial scales (point sources) are NOT affected

 First 6 epochs combined with the IDL-based mapmaker Scanamorphos (Roussel et al. 2012) to preserve extended emission...