

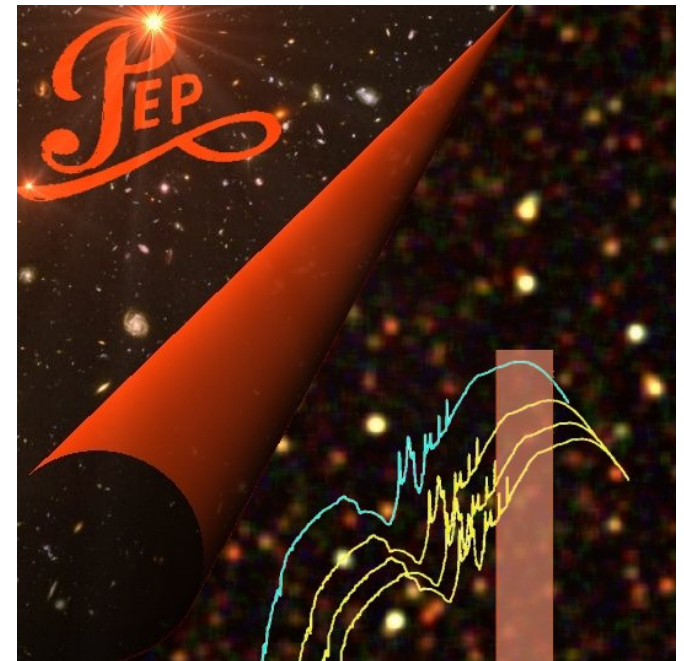
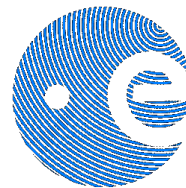
# First results from PACS deep surveys

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Dieter Lutz, for the PACS Evolutionary Probe (PEP) team

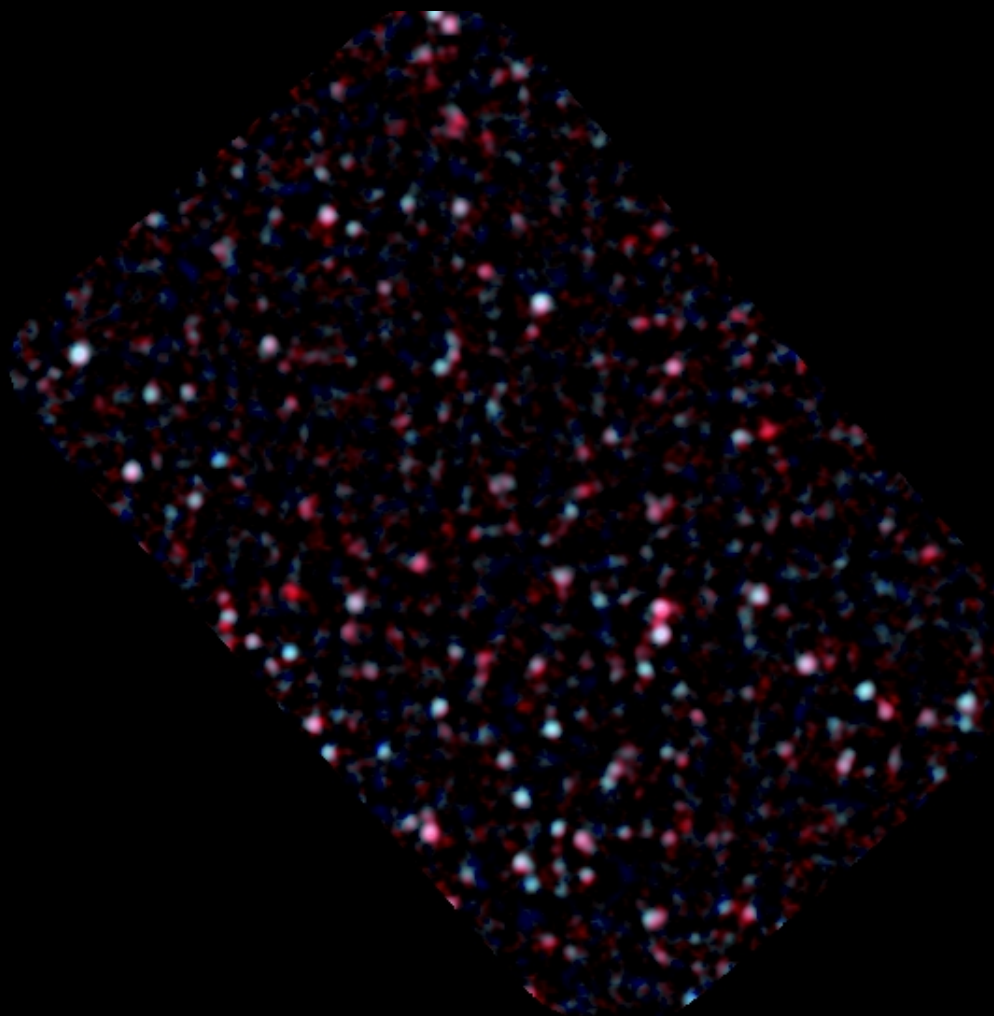
Herschel First Results Symposium - ESLAB 2010

May 4, 2010

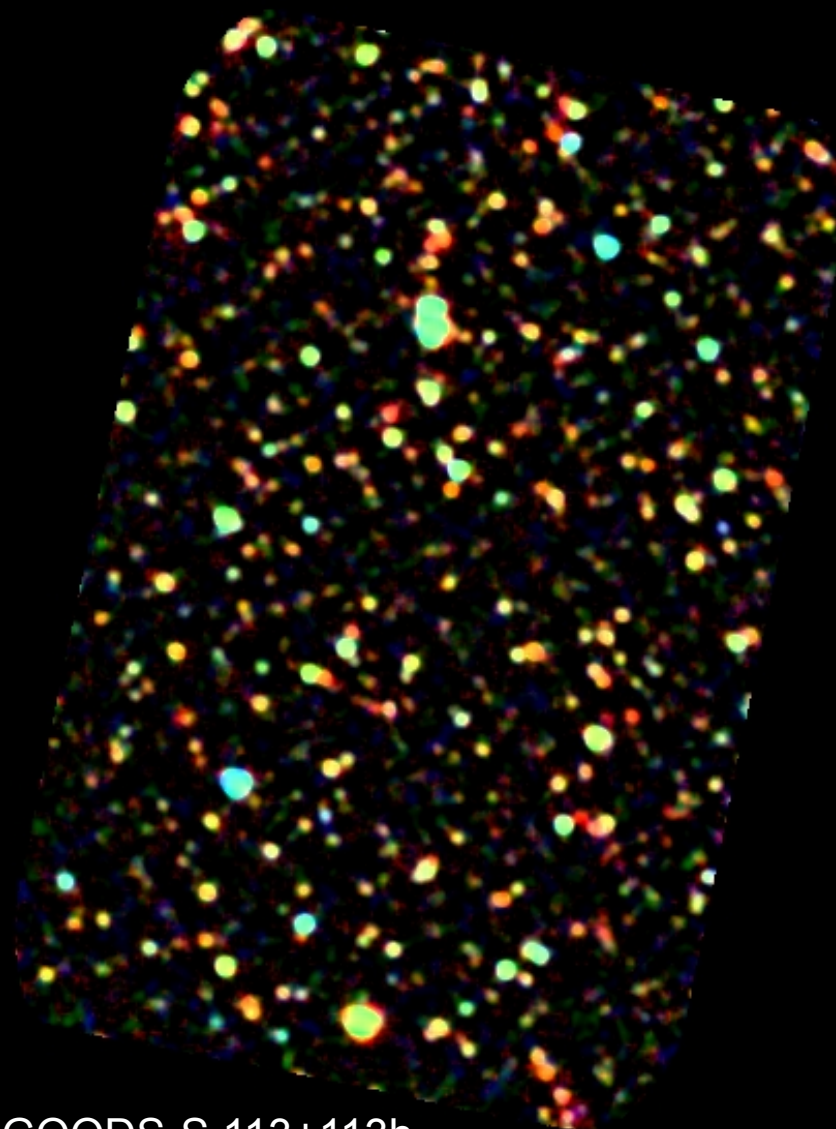


# The deepest Herschel-PACS blank fields taken to date

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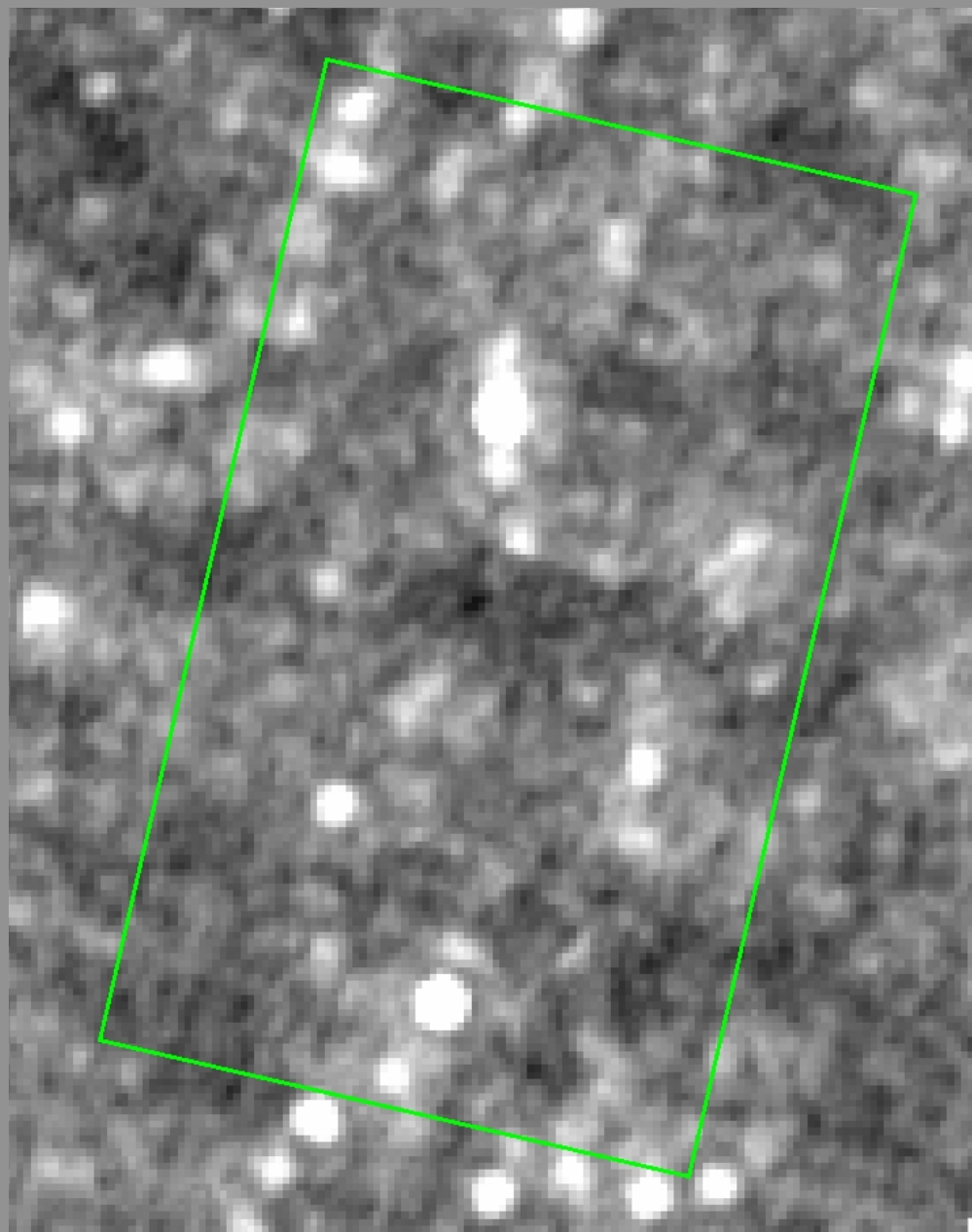
PEP GOODS-N 30h  
100+160 $\mu$ m during  
Science demonstration phase  
~300 sources



PEP GOODS-S 113+113h  
70+100+160 $\mu$ m  
~800 sources

## From MIPS to PACS

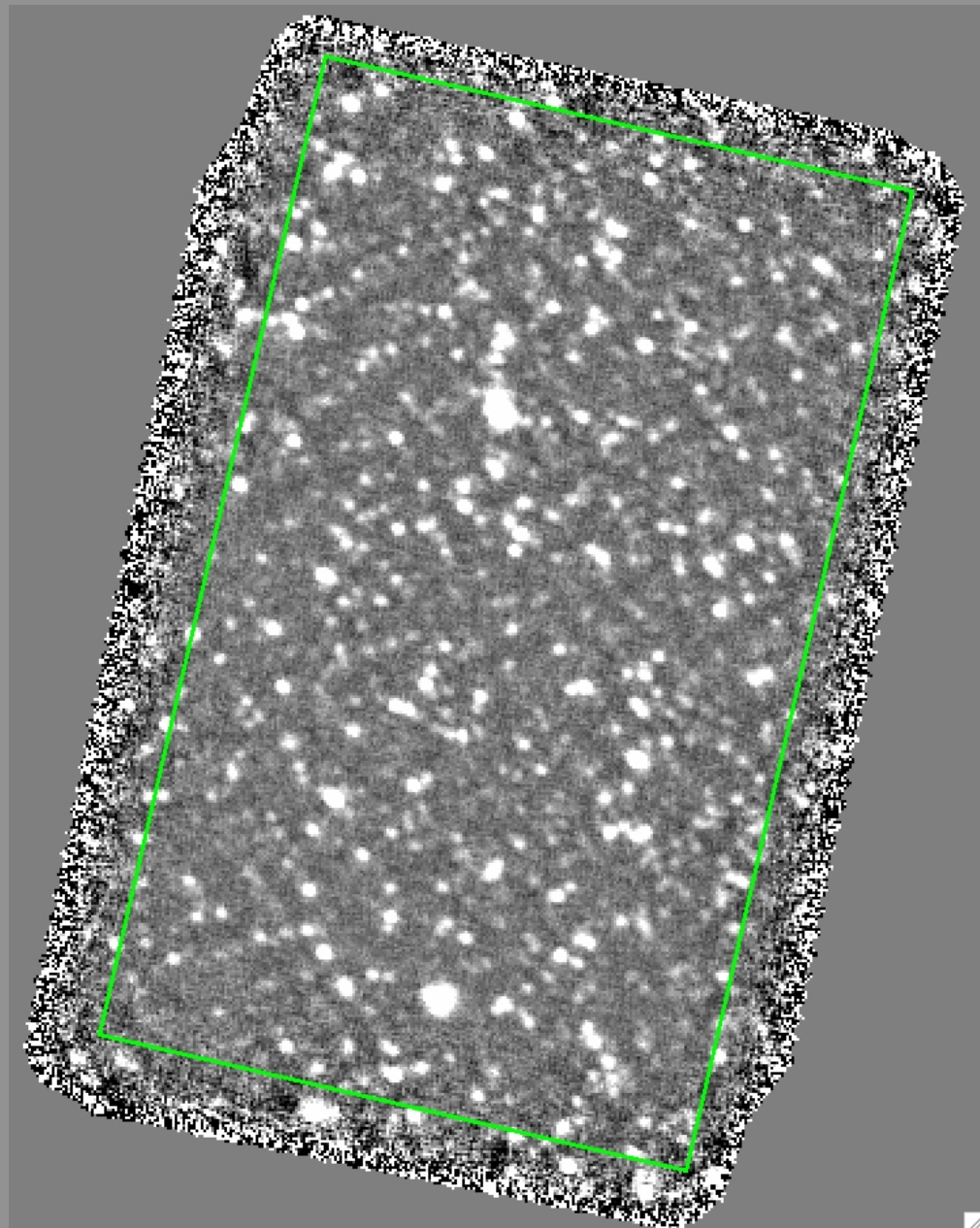
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GOODS-S MIPS 160 $\mu$ m  
FIDEL team

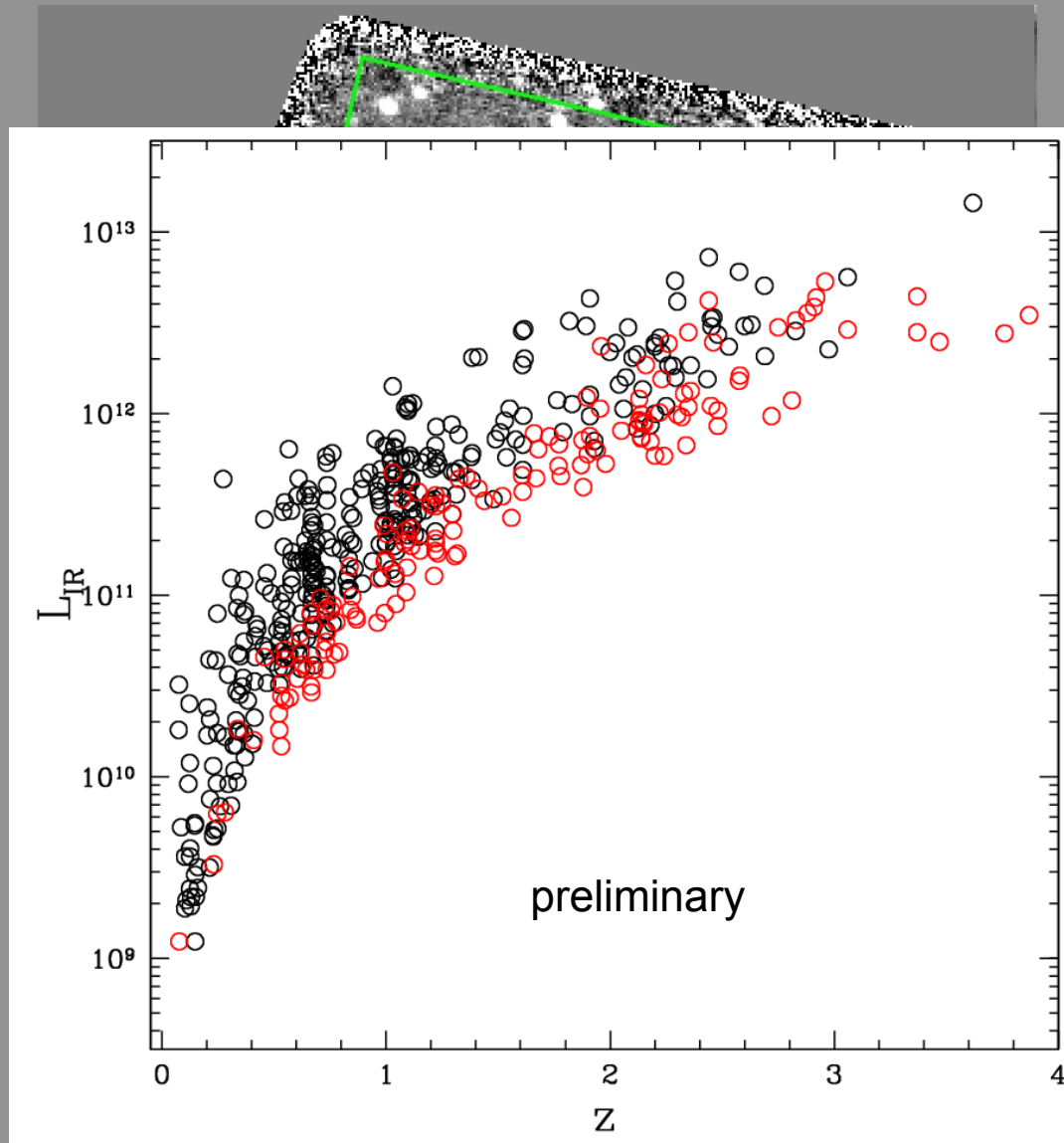
## From MIPS to PACS

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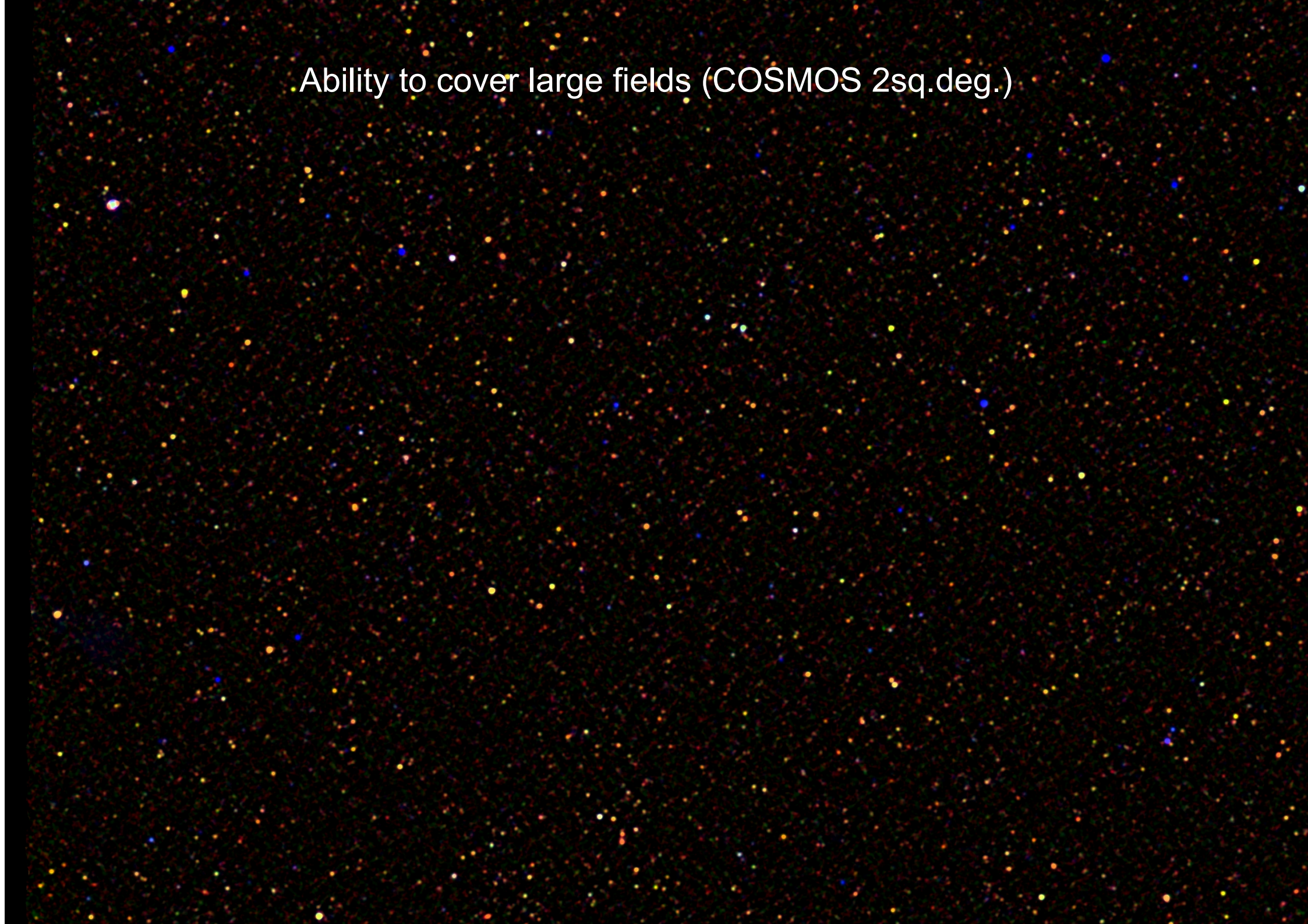
GOODS-S PACS 160 $\mu$ m  
PEP team

# From MIPS to PACS



GOODS-S PACS 160 $\mu$ m  
PEPE team

Ability to cover large fields (COSMOS 2sq.deg.)



# PEP fields

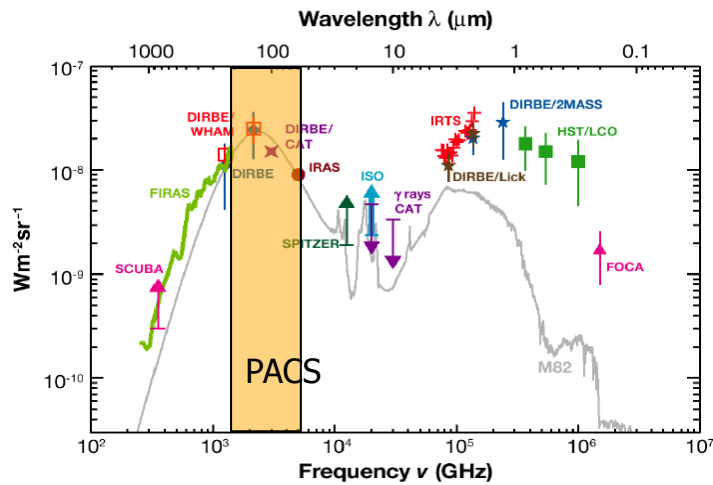
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- PEP is the major Herschel 100/160 $\mu$ m extragalactic survey of key multiwavelength fields

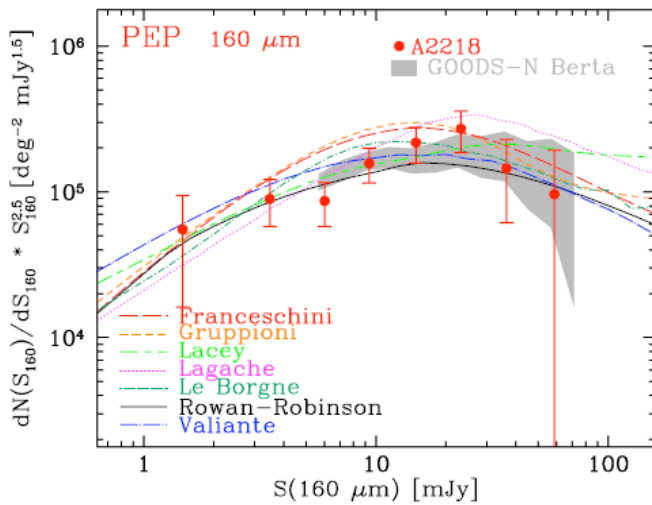
Field	Area	Total Exp. [hours]
COSMOS	85'x85'	213
Lockman Hole	24'x24'	35
E-CDFS	30'x30'	35
Groth Strip	67'x10'	35
GOODS-S	10'x15'	113 113
GOODS-N	10'x15'	30

- +10 lensing galaxy clusters
- Coordinated with Hermes for SPIRE coverage
- Hermes and Atlas extend to wider+shallower PACS coverage
- GOODS-Herschel will go deeper on (parts of) GOODS fields
- Herschel lensing survey substantially extends the number of lensing clusters

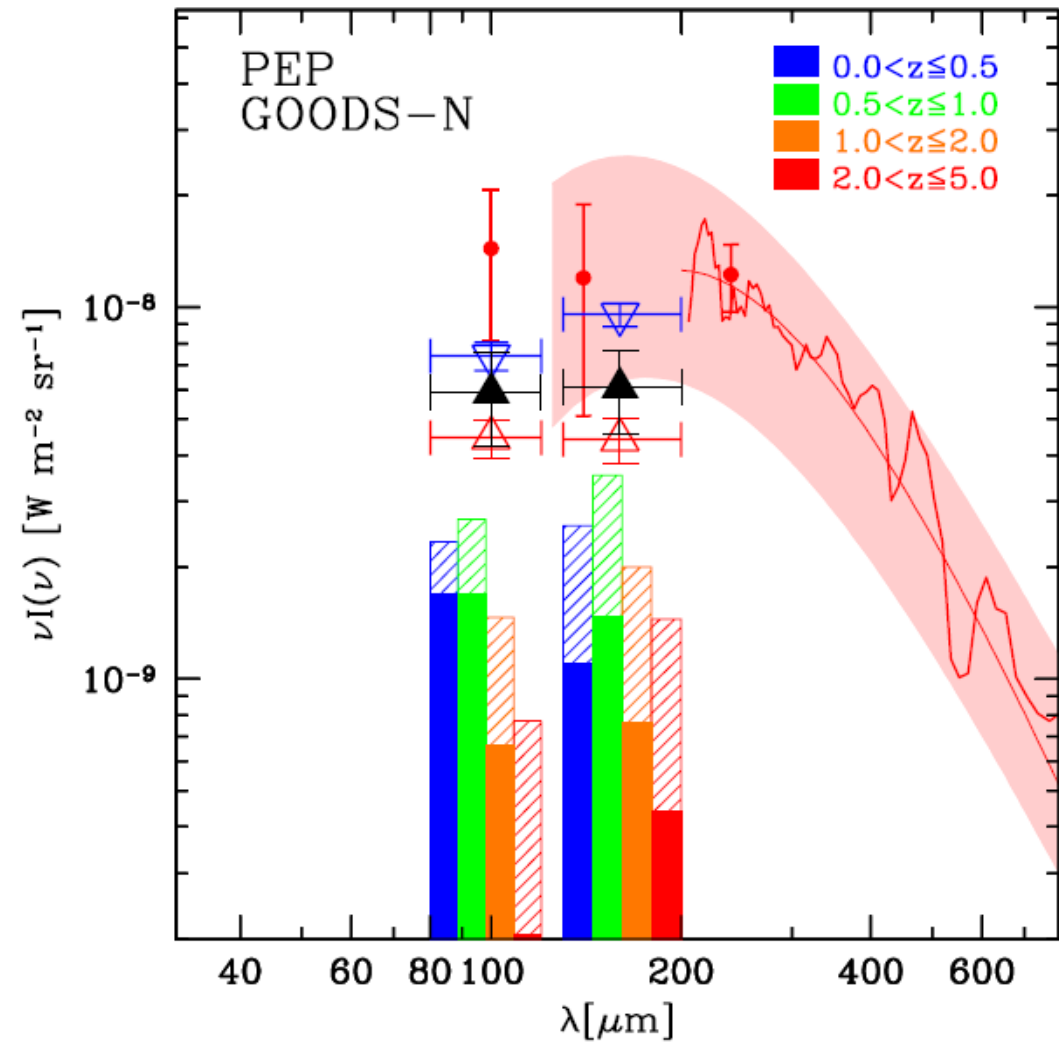
# Resolving the Cosmic Infrared Background with PACS



Lagache et al. 2005 ARAA



Altieri et al. 2010 Abell 2218 lensed counts

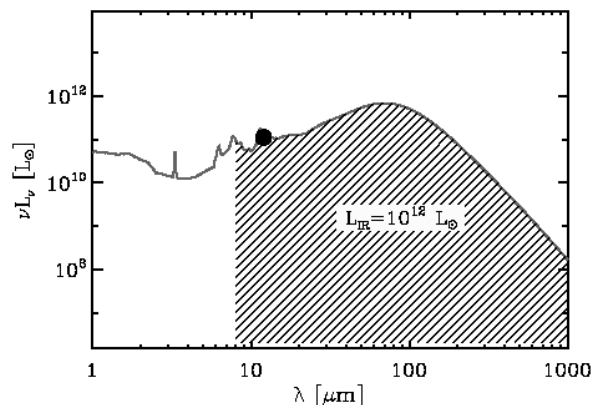


Berta et al. 2010 (talk on Thursday)

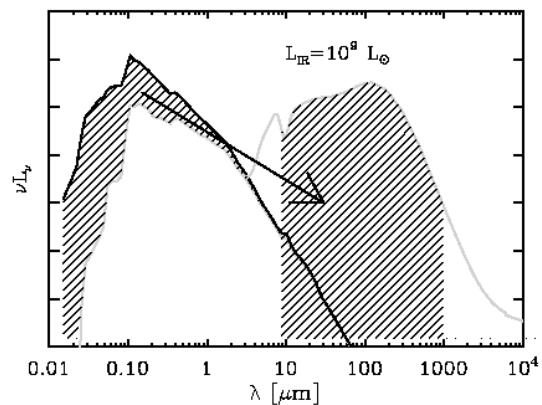


## The need for far-IR calorimetric star formation rates

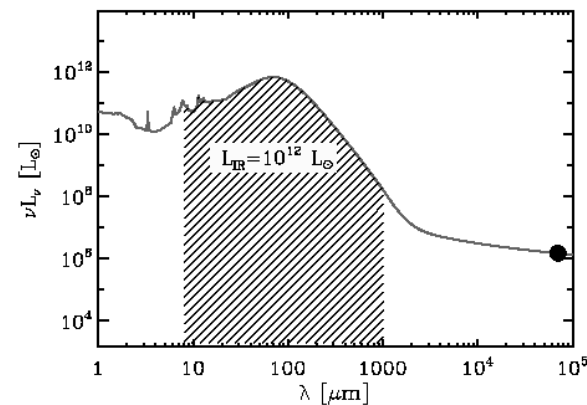
- Our community has been relying almost exclusively on extrapolation from the optical and mid-infrared as the avenue towards studying galaxy evolution and star formation rates
- We know this extrapolation is pretty good
- **But how good?**



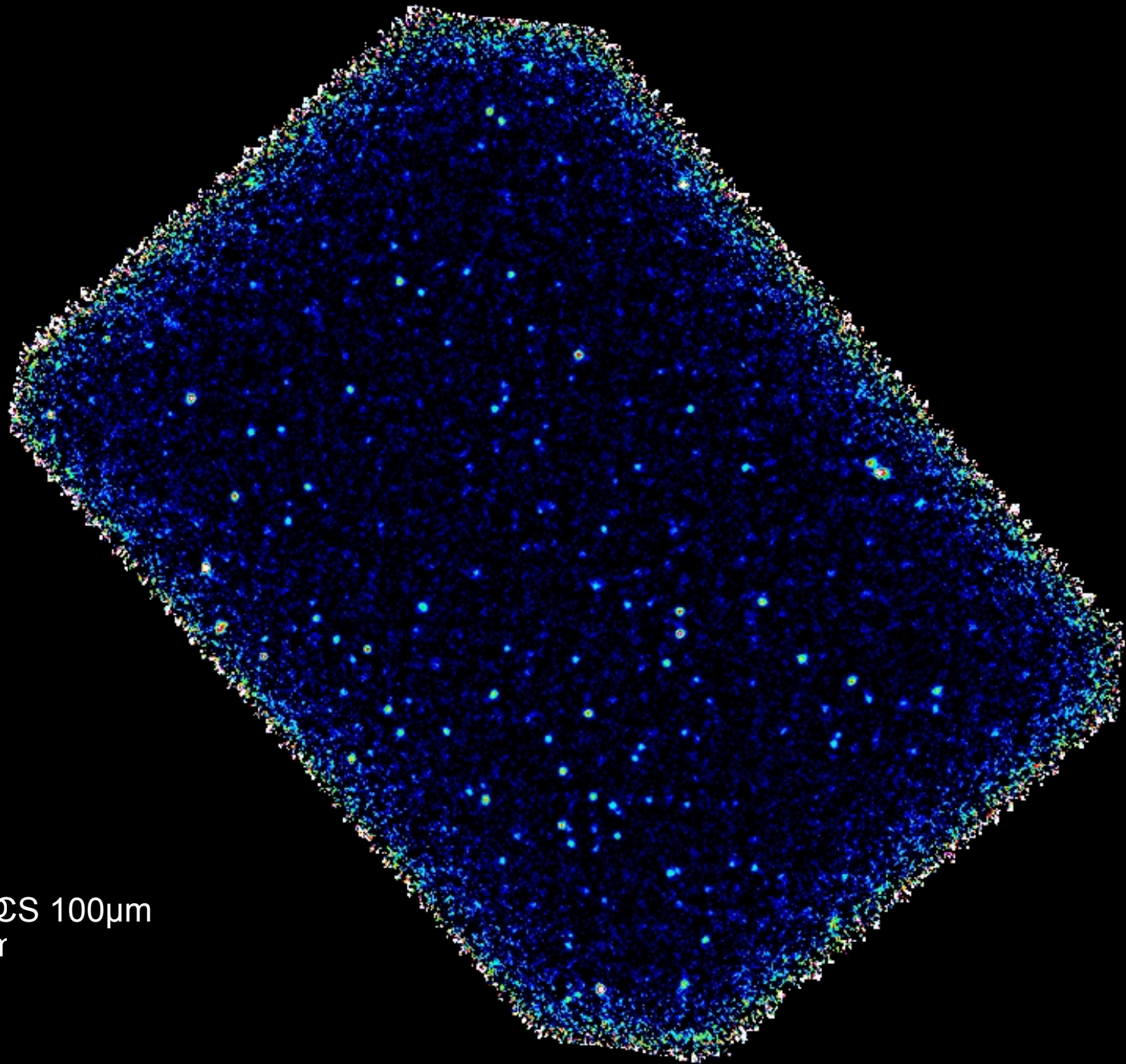
From 24 $\mu\text{m}$



From rest frame UV



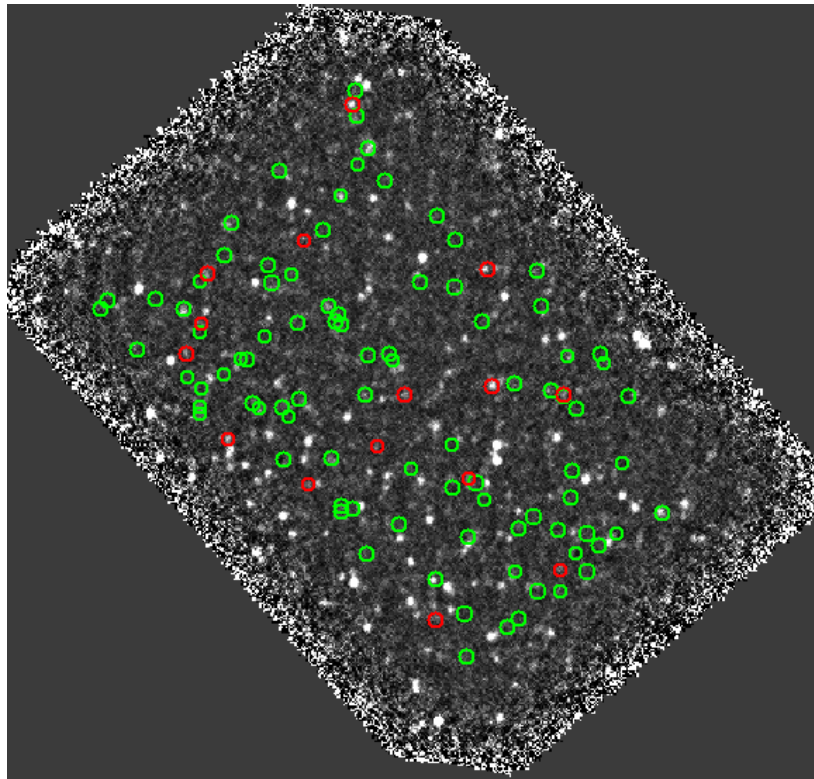
From submm/radio



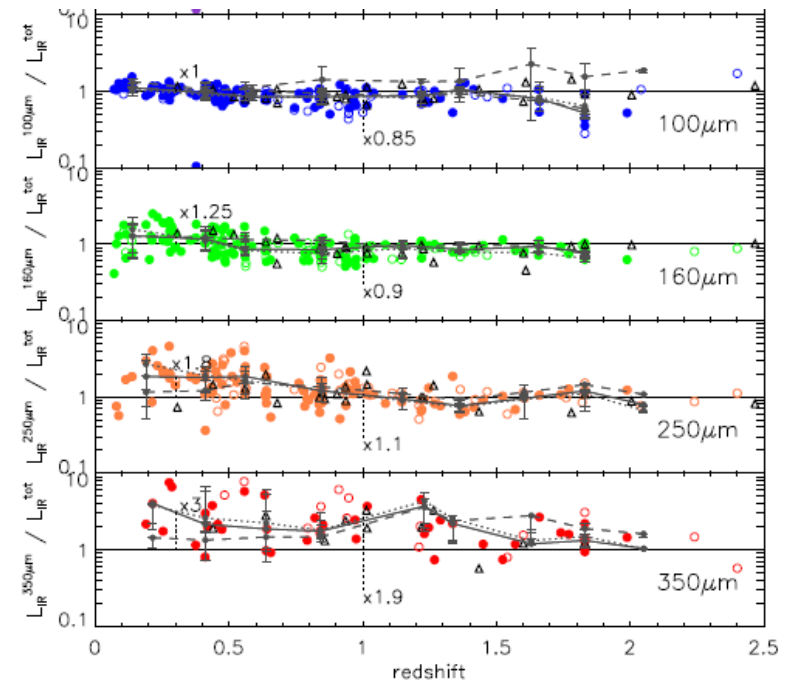
ExpAACS 100µm  
100µm

# The star formation rates of typical $z \sim 2$ star forming galaxies

- BzK star-forming galaxies in GOODS-N,  $K_{AB} < 22$ ,  $z = 1.5 - 2.5$
- Far-infrared luminosity from 160 $\mu\text{m}$  flux, redshift, Chary & Elbaz 2001 SED

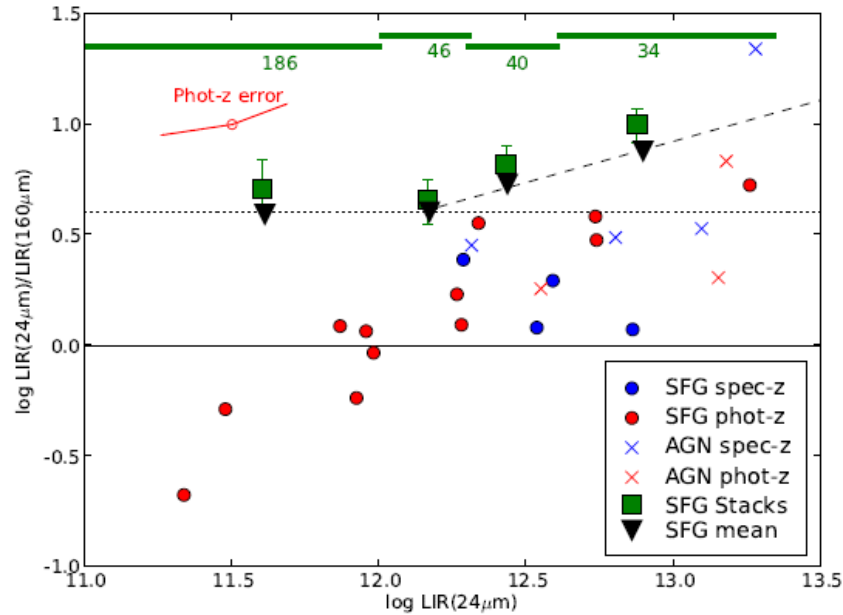


Nordon et al. 2010



Elbaz et al. 2010 (next talk)

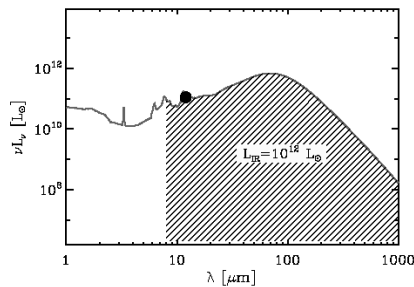
## Z~2: Extrapolation from 24μm overpredicts FIR



- Star formation rates based on 24μm, z, and Chary/Elbaz SED family **overpredict the calorimetric FIR by factors 4-7.5 at this redshift**

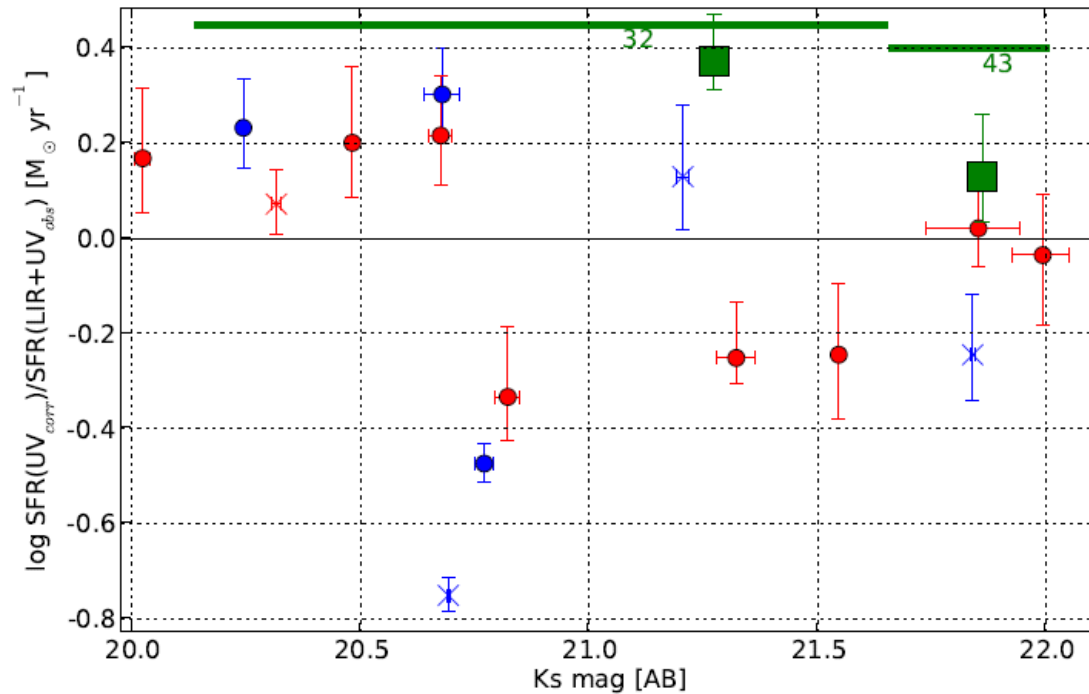
- Similar behaviour of X-ray AGN and other galaxies

- **Obscured AGN and/or changing SED shape/PAH strength?** Setting in of the effect at  $z=1.5$  favours the latter, to be continued...



Nordon et al. 2010 (poster)

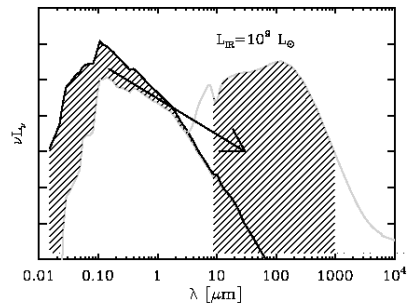
# Z~2: Extrapolation from rest frame UV slightly overpredicts FIR



Extrapolation from rest frame UV (used here: Daddi (2004) BzK recipe and Calzetti (2000) extinction law

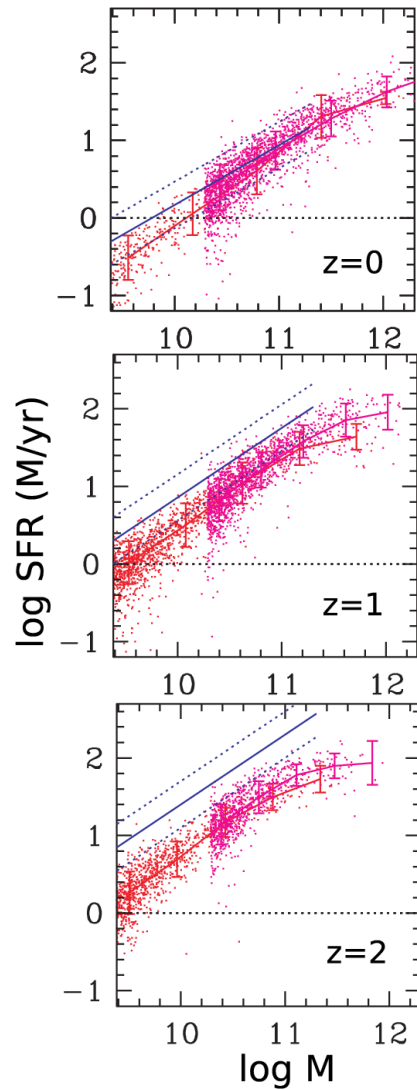
**Overprediction by factor ~2**

Modest modification to extinction law needed?

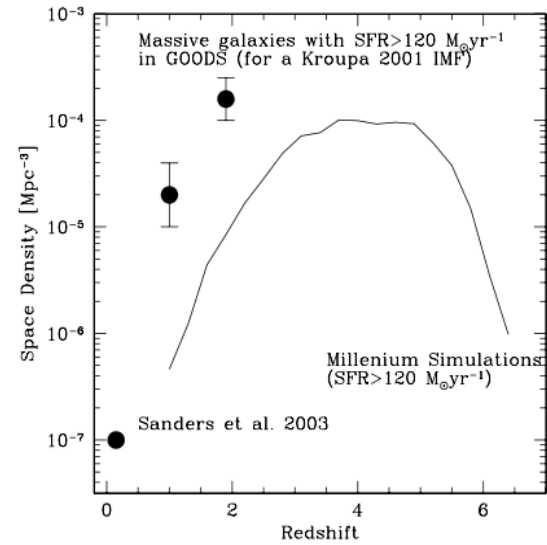


Nordon et al. 2010 (poster)

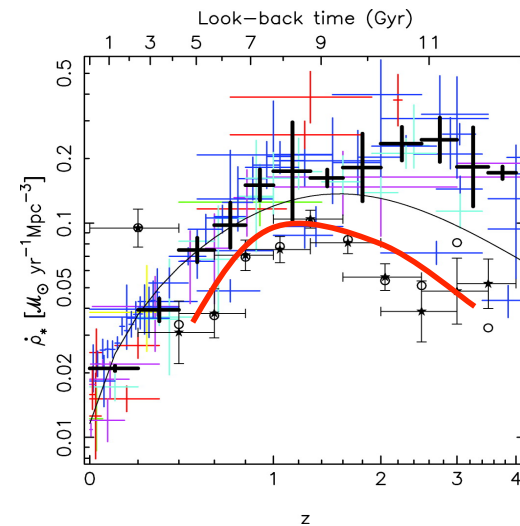
# Towards reconciling observed and theoretical star formation rates



Dave 08



Daddi+ 07

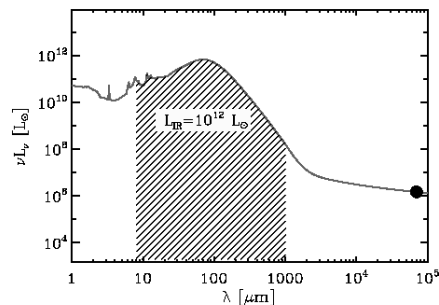
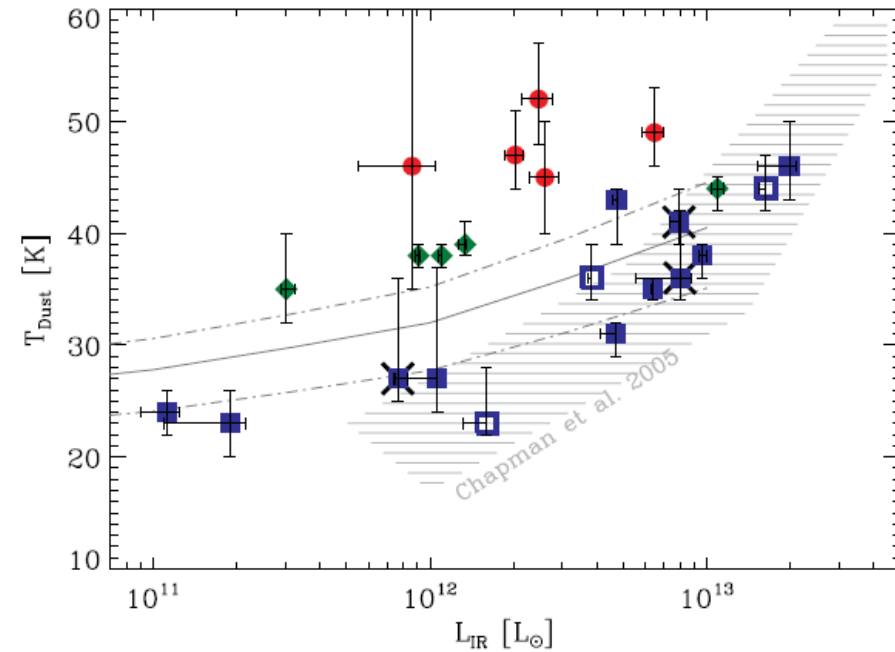


Perez-Gonzalez+08

# Extrapolation from submm/radio assuming radio/FIR correlation

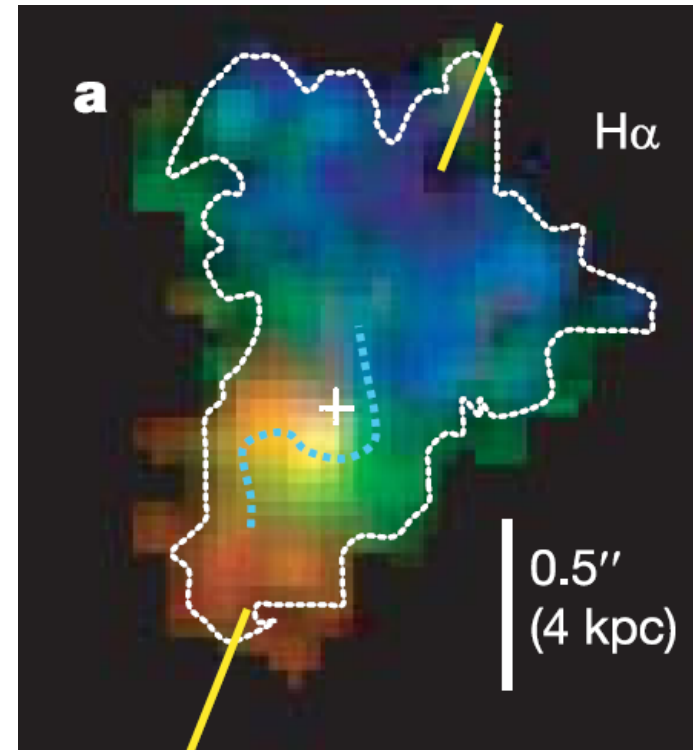
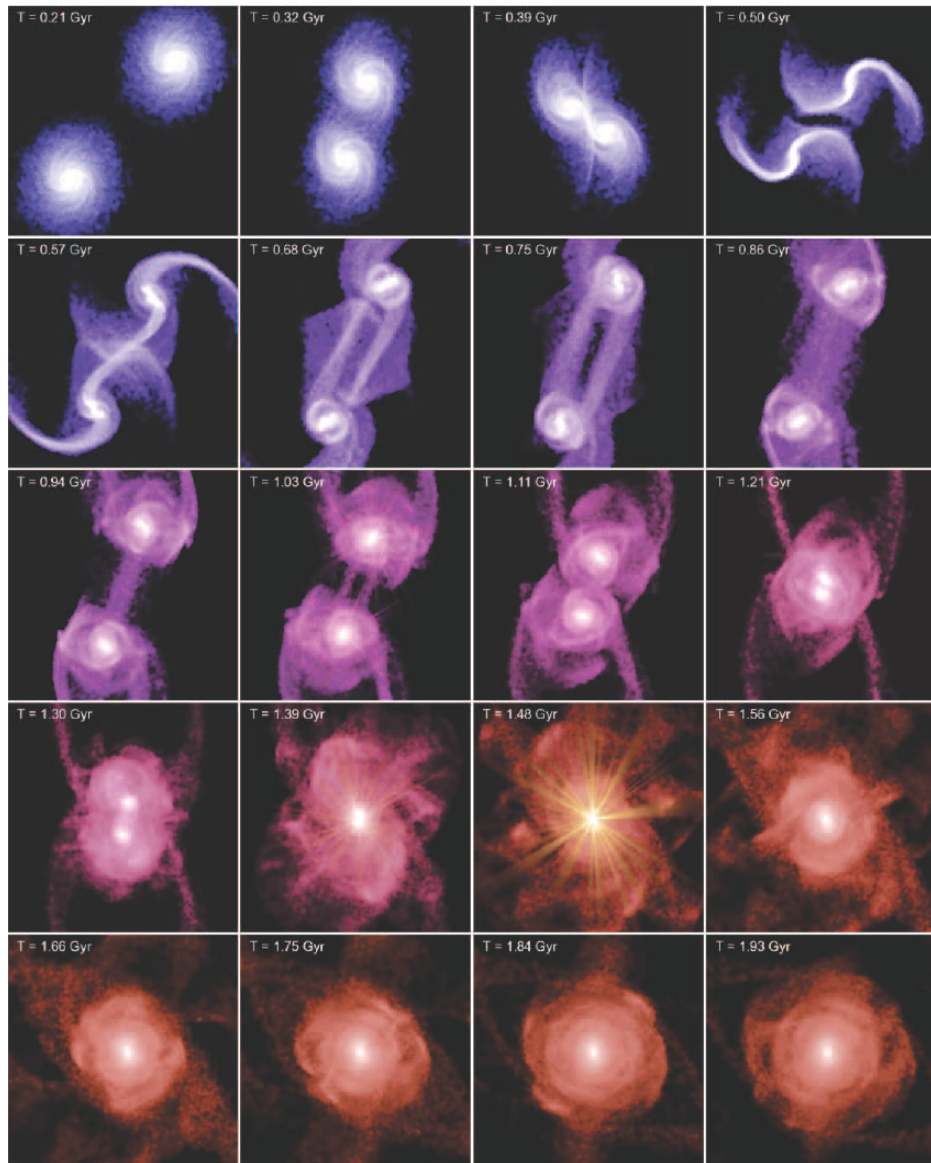
Very high IR luminosities of SMGs as estimated from submm/radio via local radio-FIR relation are substantiated by Herschel data

.. Previous selection effects on dust temperature can now be largely avoided



Magnelli et al. 2010 (talk on Thursday)

# The co-evolution of AGN and star formation

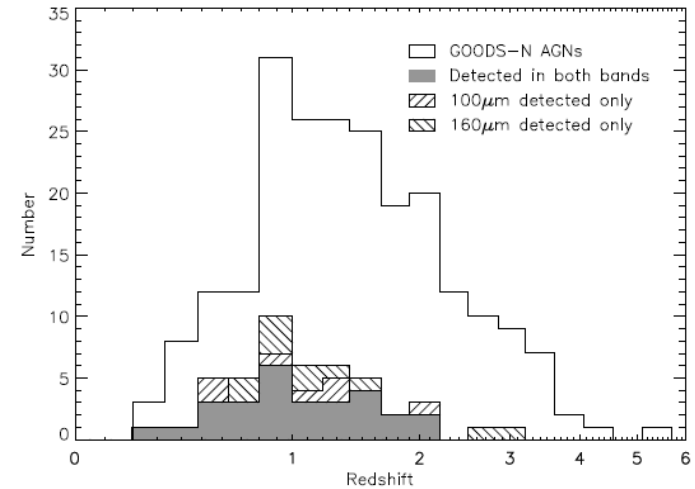
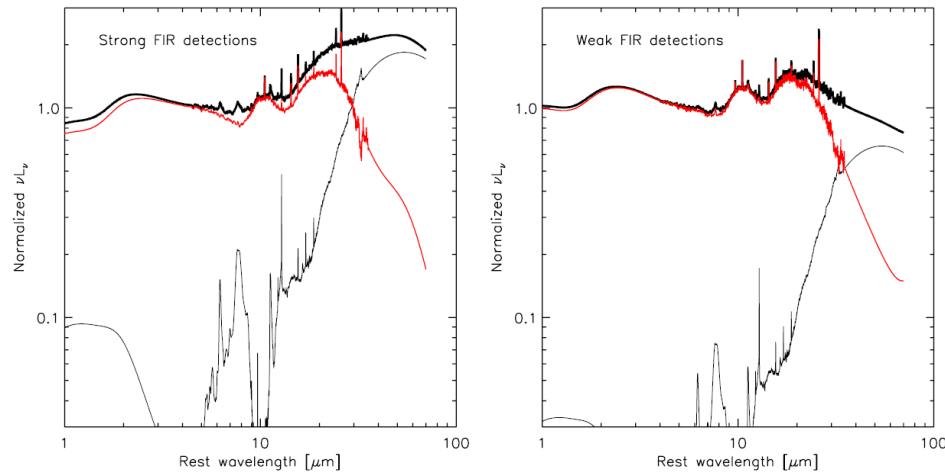


BzK-15504  $z \sim 2.38$  rotating disk with central AGN (Genzel+08)

Models of merging galaxies (Hopkins+06)

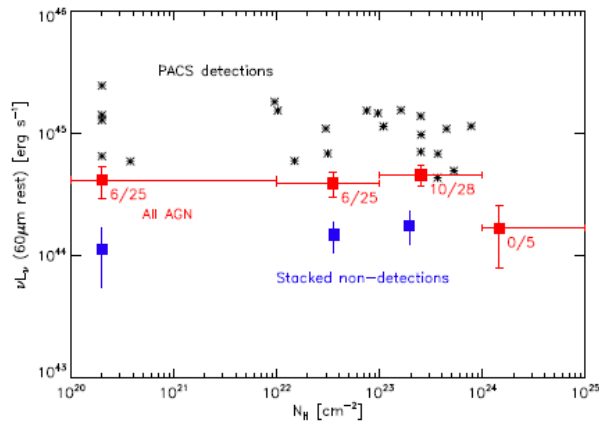


# X-ray selected AGN in GOODS-N: Using FIR to measure star formation



FIR has best contrast between host star formation and intrinsic AGN SED (QSO SEDs from Netzer+07)

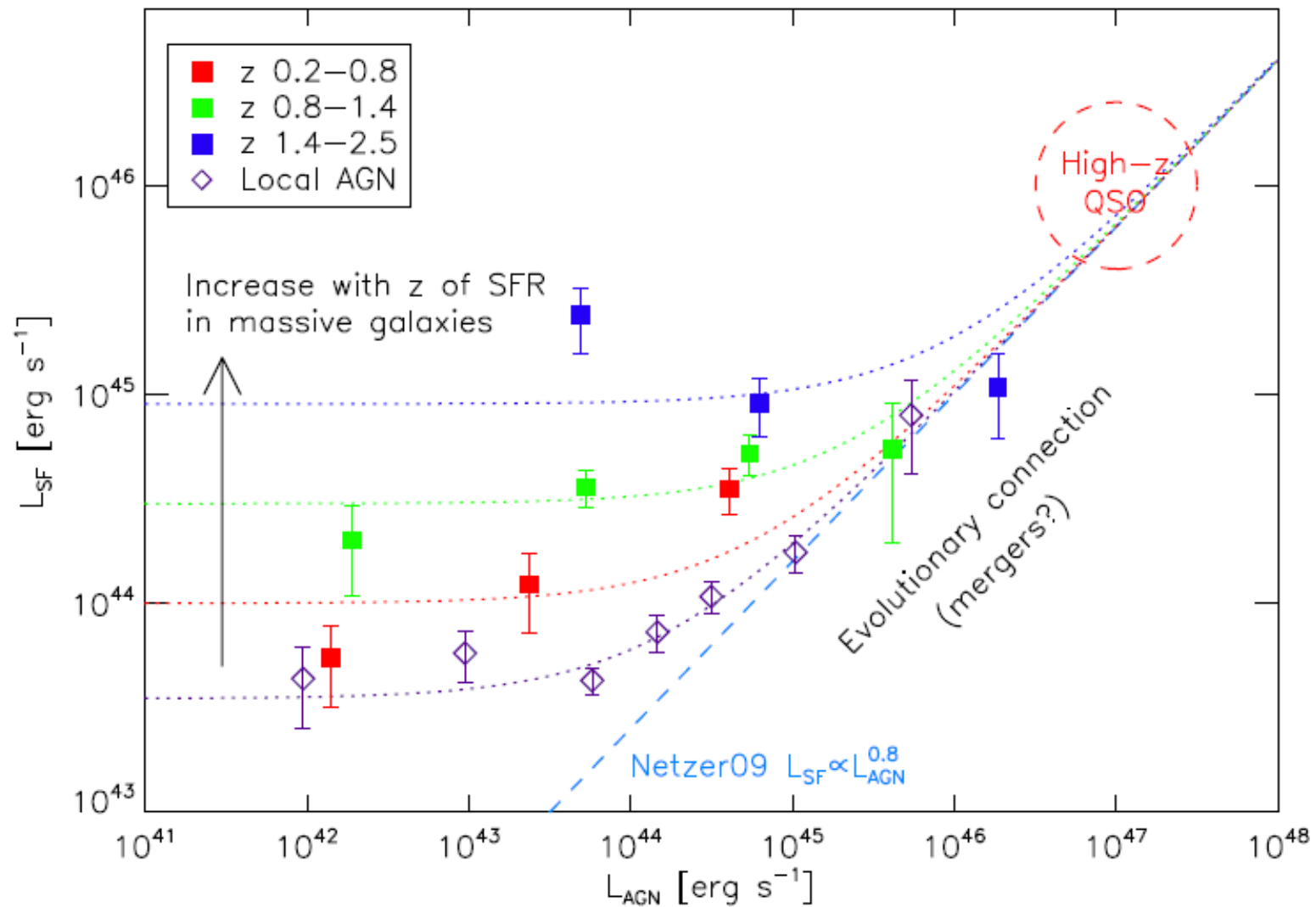
FIR detection rate 21% for X-ray AGN



No trend with obscuration (but only few  $L_X > 10^{44}$  in sample)

Shao et al. 2010 (poster)

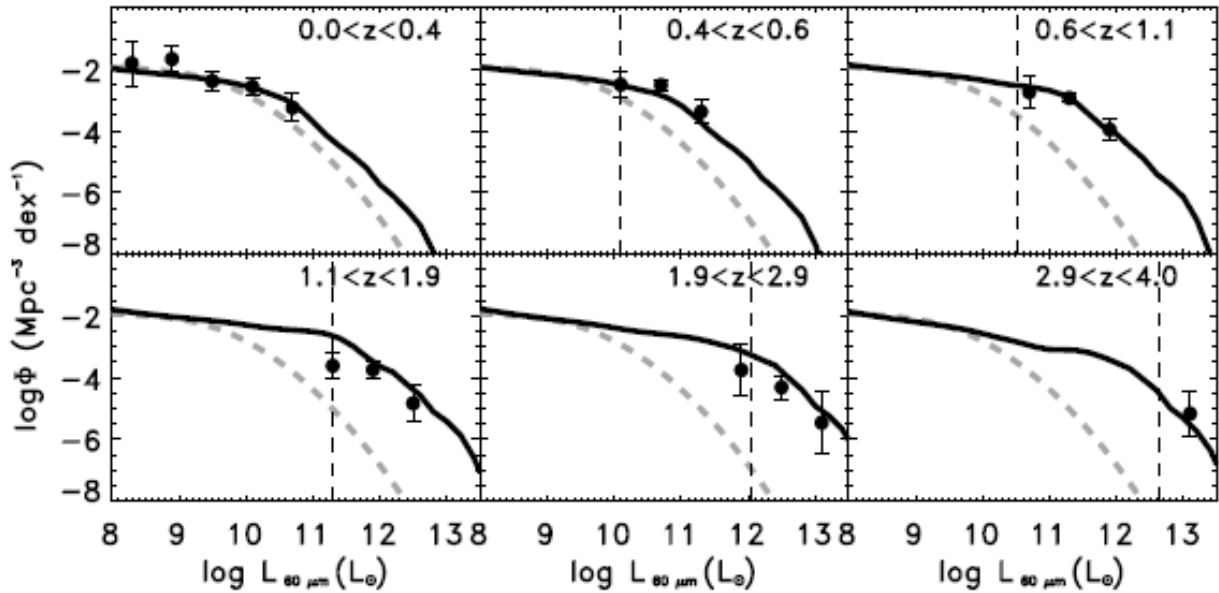
# Two modes of AGN / host coevolution: Merger vs. secular



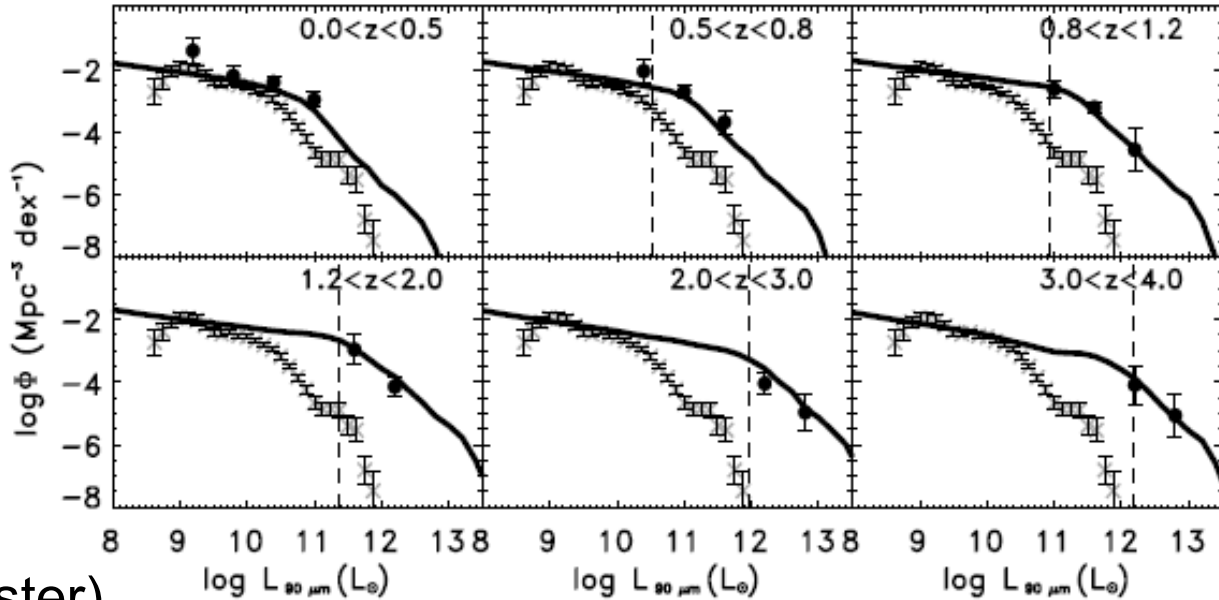
Shao et al. 2010 (poster)

# Direct far-infrared luminosity functions

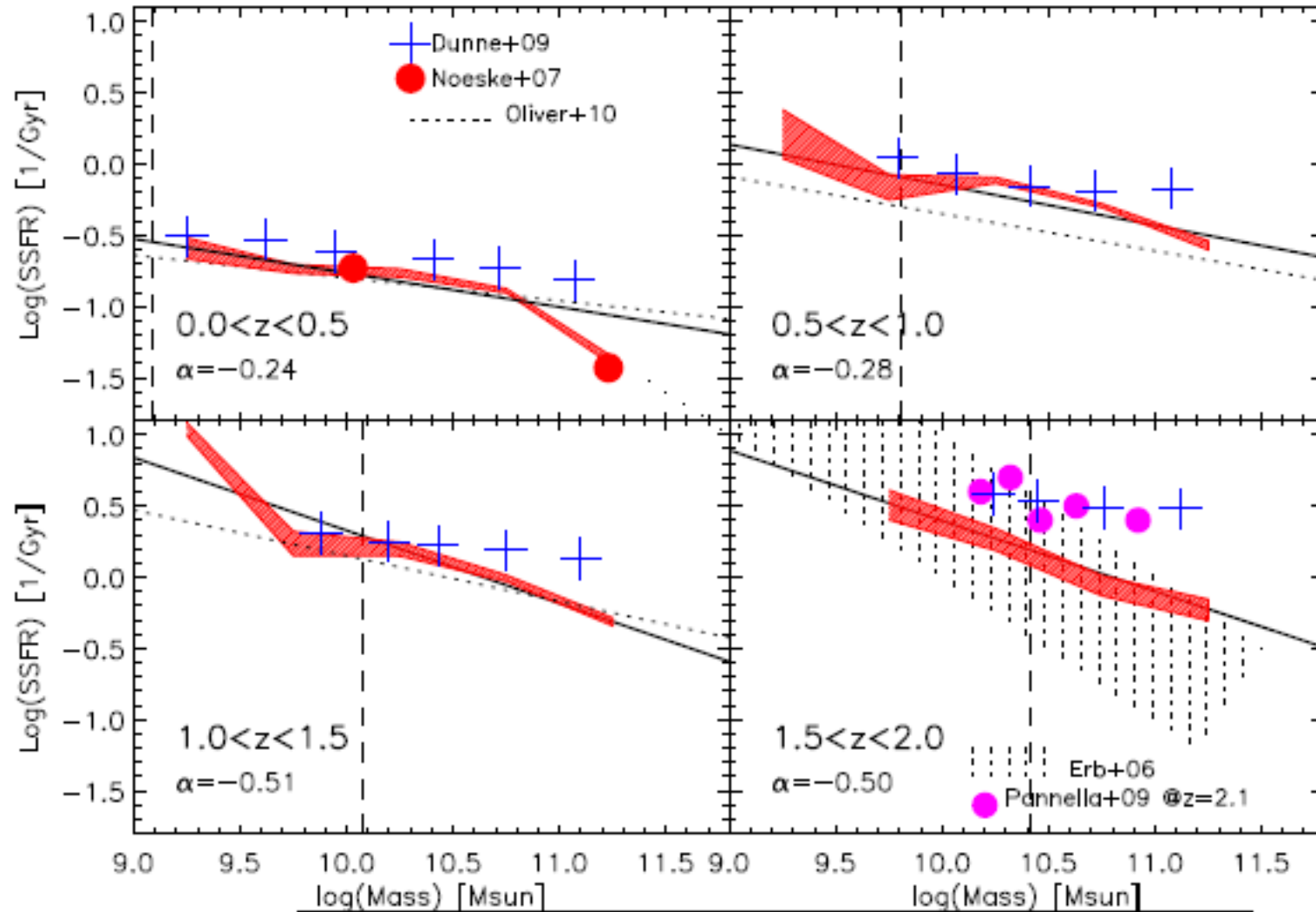
Rest frame  
60 $\mu$ m



Rest frame  
90 $\mu$ m

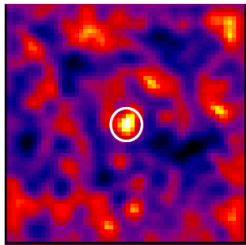
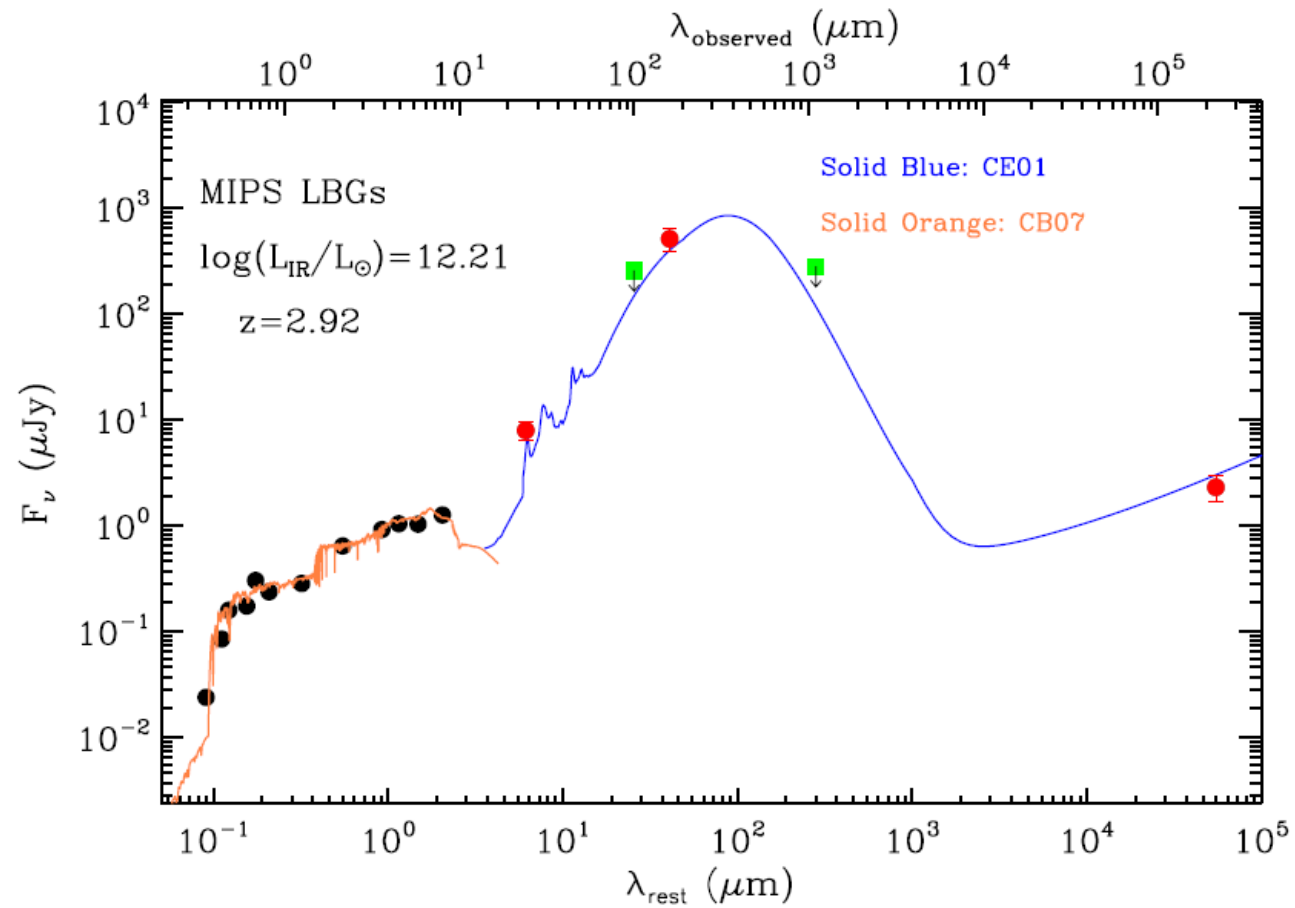


# FIR-based determinations of the specific star formation rate



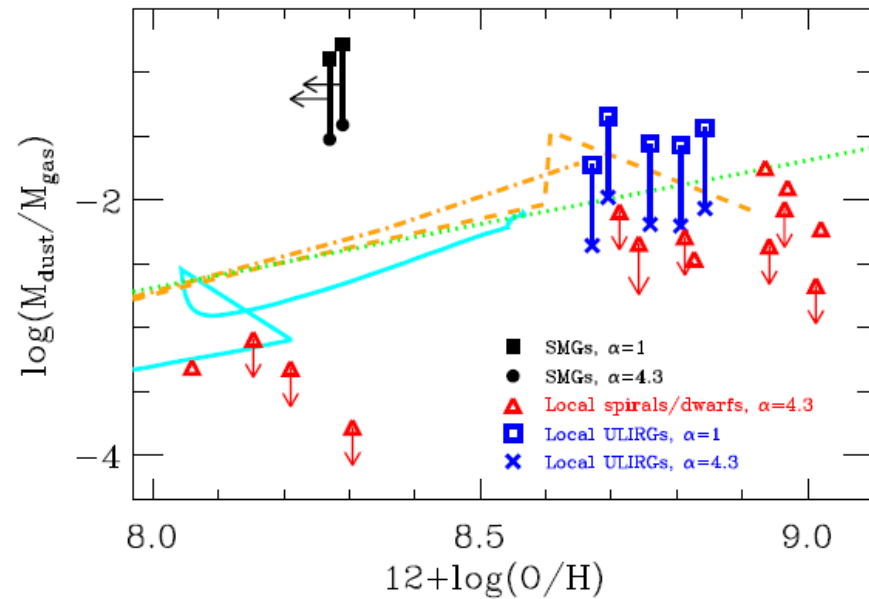
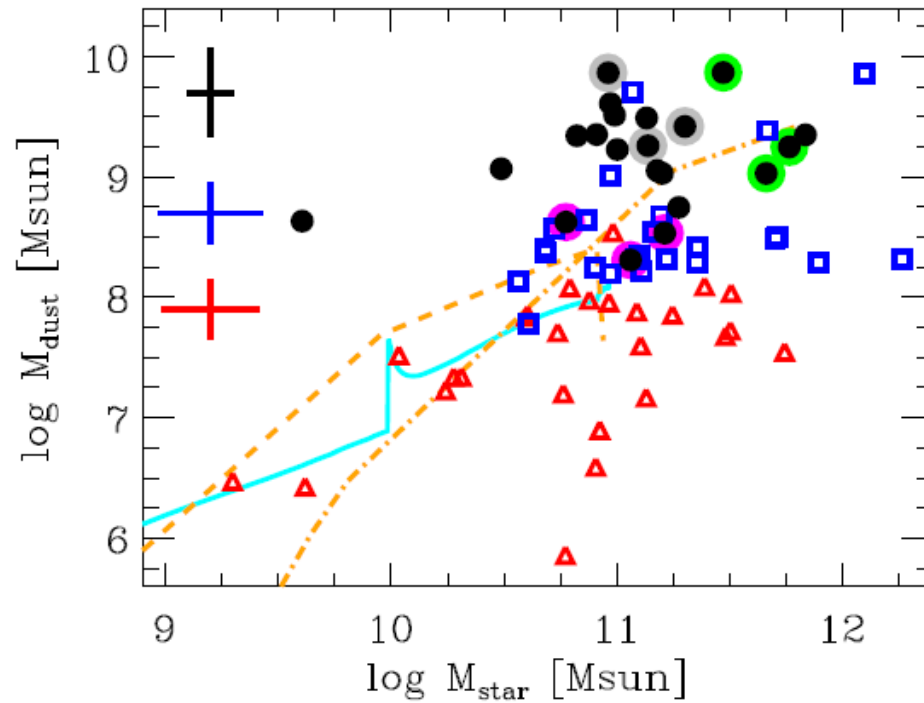
Rodighiero et al. 2010 (talk on Thursday)

# Stacked far-infrared detection / SED constraints for Lyman Break Galaxies



Magdis et al. 2010 (Poster)

# Surprisingly large dust masses of submillimeter galaxies



Santini et al. 2010

... more dust than expected for gas phase metallicity

# Thank you!

Jose Acosta  
Bruno Altieri  
Paola Andreani  
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Stefano Berta  
Angel Bongiovanni  
Damien Le Borgne  
Nicolas Bouche  
Drew Brisbin  
Hector Castaneda  
Antonio Cava  
Jordi Cepa  
Andrea Cimatti  
Emanuele Daddi  
Helmut Dannerbauer  
Helena Dominguez-Sanchez  
David Elbaz  
Emeric Le Floc'h  
Natascha Förster Schreiber  
Reinhard Genzel  
Ignacio Gonzalez  
Gianluigi Granato  
Andrea Grazian  
Carlotta Gruppioni  
Martin Harwit  
Ho-Seong Hwang  
Georgios Magdis  
Benjamin Magnelli  
Roberto Maiolino  
Leo Metcalfe  
Raanan Nordon  
Koryo Okumura  
Ana Perez  
Ismael Perez Fournon  
Albrecht Poglitsch  
Paola Popesso  
Francesca Pozzi  
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Jose Miguel Rodriguez  
Amelie Saintonge  
Fadia Salmi  
Miguel Sanchez  
Paola Santini  
Li Shao  
Eckhard Sturm  
Linda Tacconi  
Ivan Valtchanov  
Michael Wetzstein  
Eckhard Wieprecht

Lockman Hole

