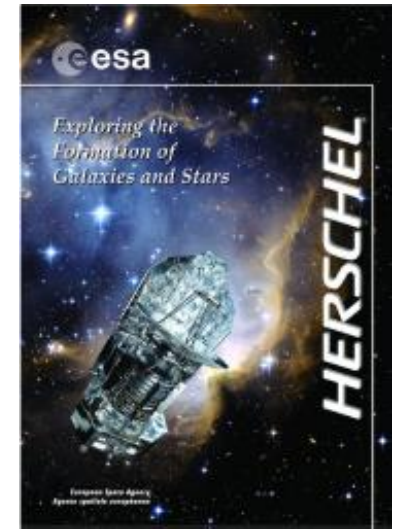




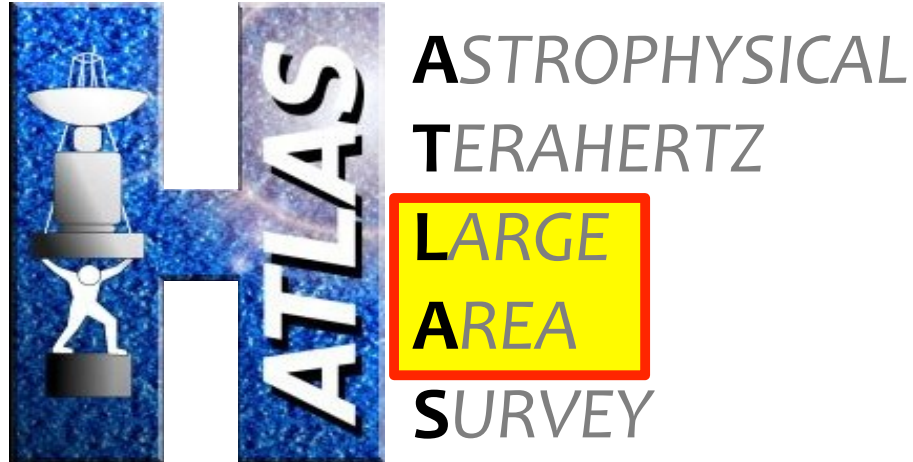
ASTROPHYSICAL  
TERAHERTZ  
LARGE  
AREA  
SURVEY



# The Herschel-ATLAS: Discovery and Properties of the Brightest 500 $\mu\text{m}$ Galaxies in the Universe

Mattia Negrello (The Open University, UK)  
on behalf of the H-ATLAS team

# WHY STUDYING SUB-MM BRIGHT GALAXIES?

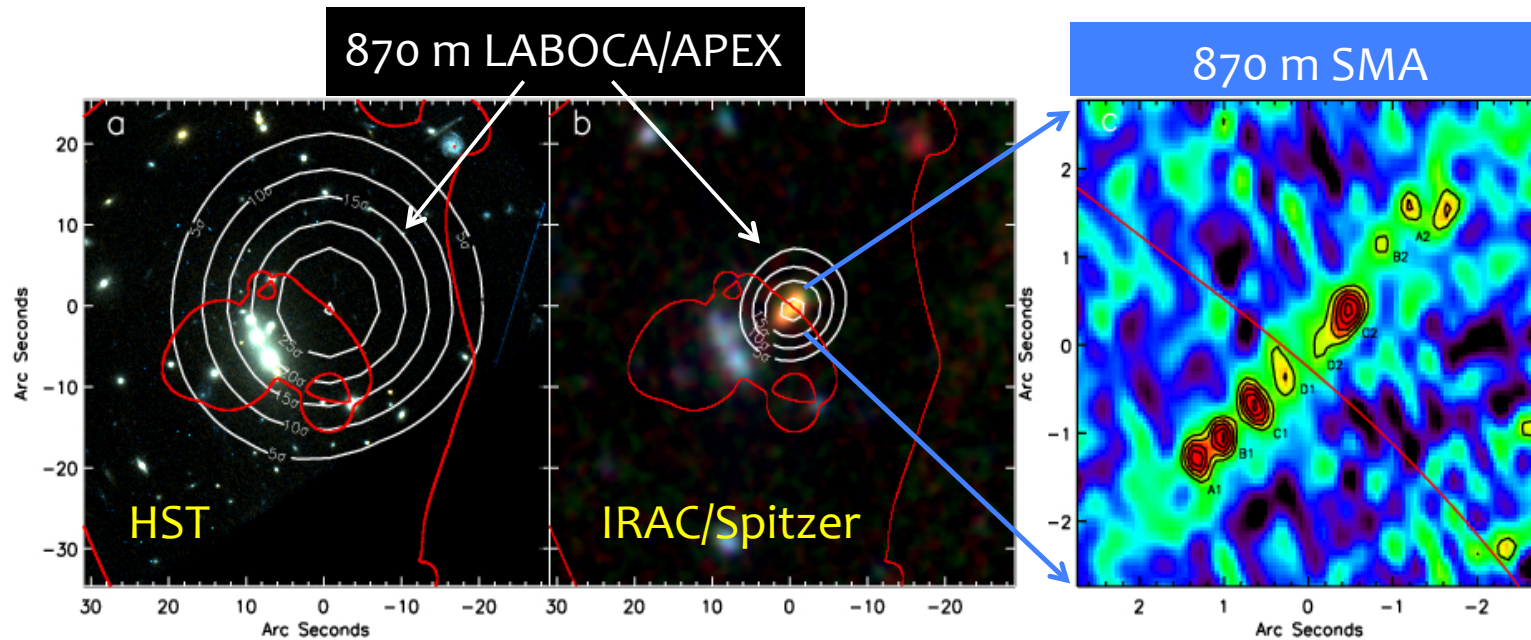


- constraints on galaxy formation/evolution models
- discovery of “extreme” ( $L_{\text{IR}} > \sim 10^{14}$ ) **HLIRGs**

# WHY STUDYING SUB-MM BRIGHT GALAXIES?

What's the role of **gravitational lensing** ?

dust-obscured star-forming galaxy at  $z=2.3$  lensed by a foreground galaxy cluster lens  $\Rightarrow S_{870\mu\text{m}}=107 \text{ mJy} !!$



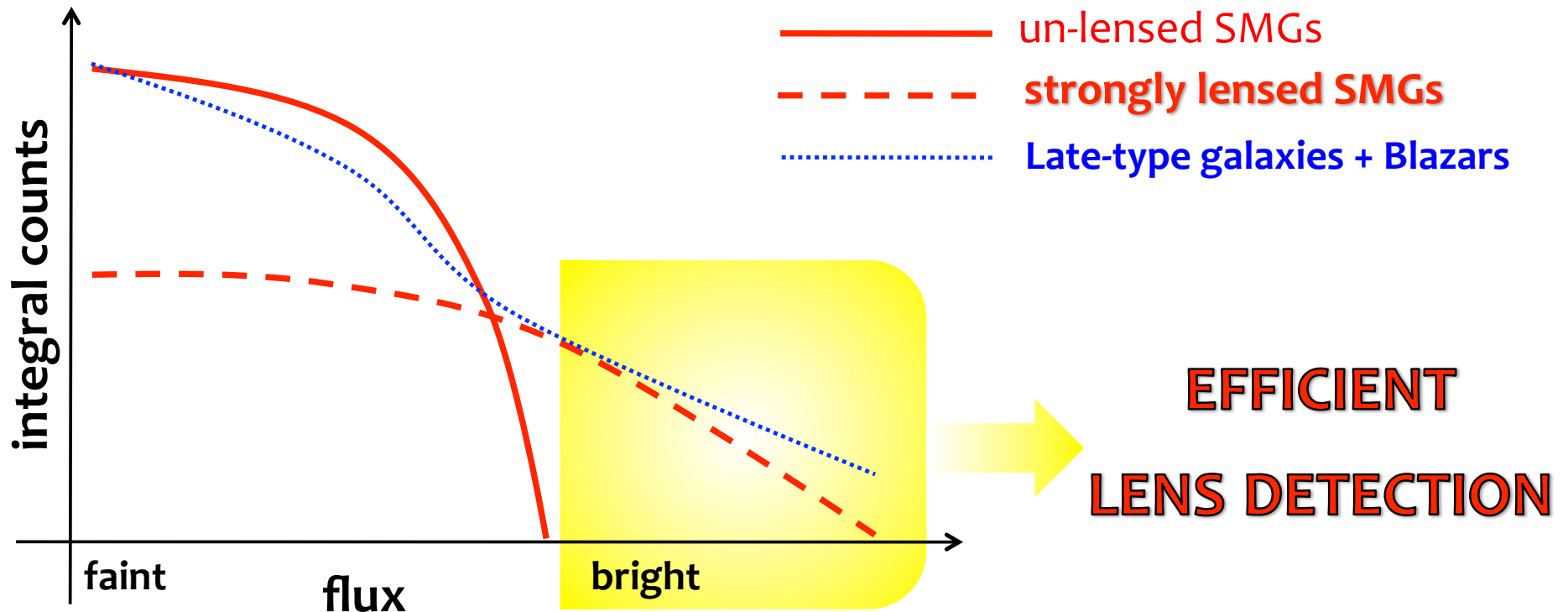
*Swinbank et al., Nature (2010)*

# A LENSING SCIENCE CASE FOR H-ATLAS

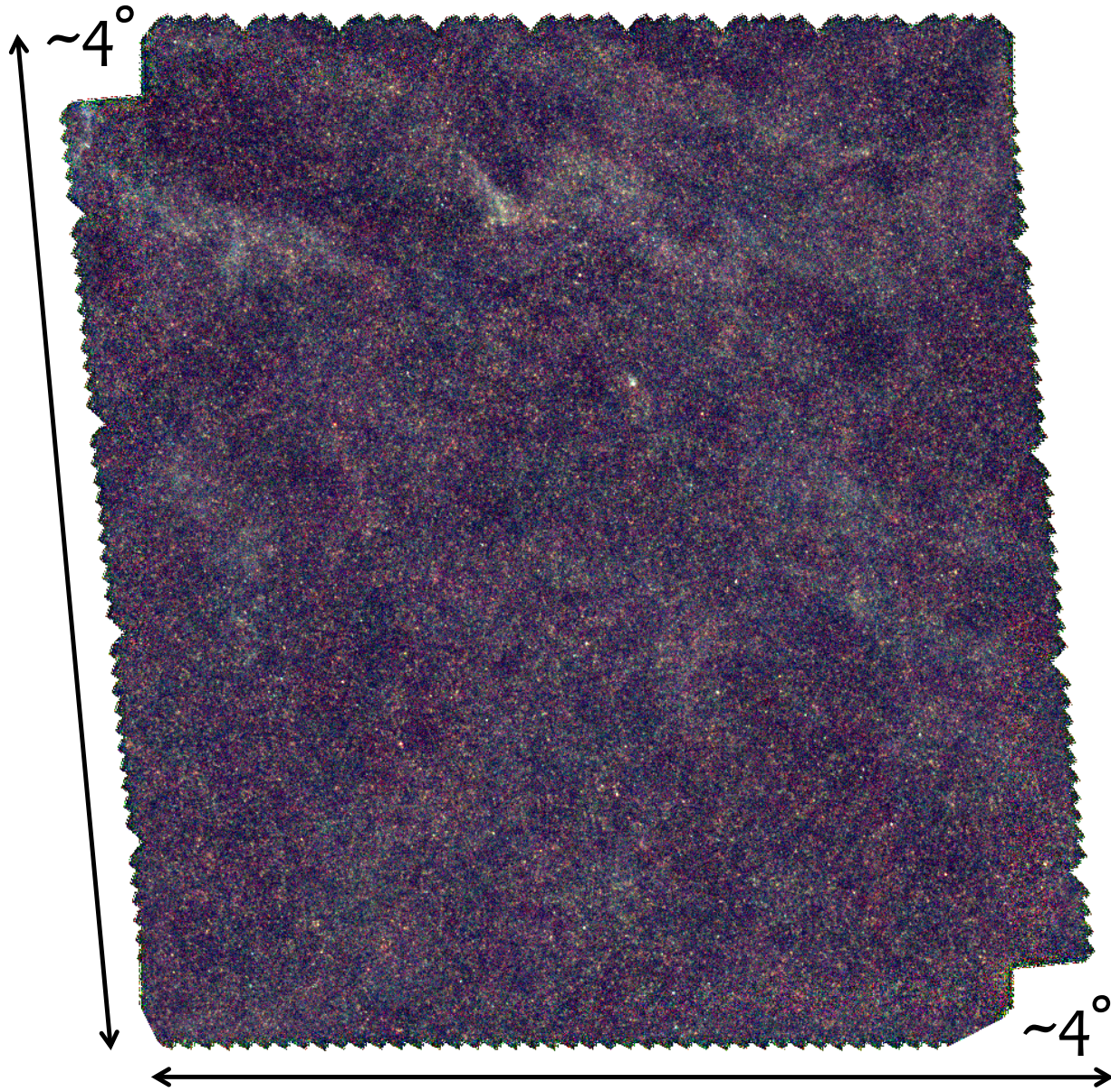
## Sub-mm surveys are ideal for finding lenses

*Blain (1996), Perrotta et al. (2003), Negrello et al. (2007)*

- **high redshift** ➔ **high efficiency for lensing**  
*Chapman et al. (2005)*
- **steep counts** ➔ **strong magnification bias**  
*Coppin et al. (2006)*



# 500 $\mu\text{m}$ BRIGHTEST GALAXIES IN H-ATLAS SDP



H-ATLAS SDP field

➤  $\sim 14.4 \text{ deg}^2$

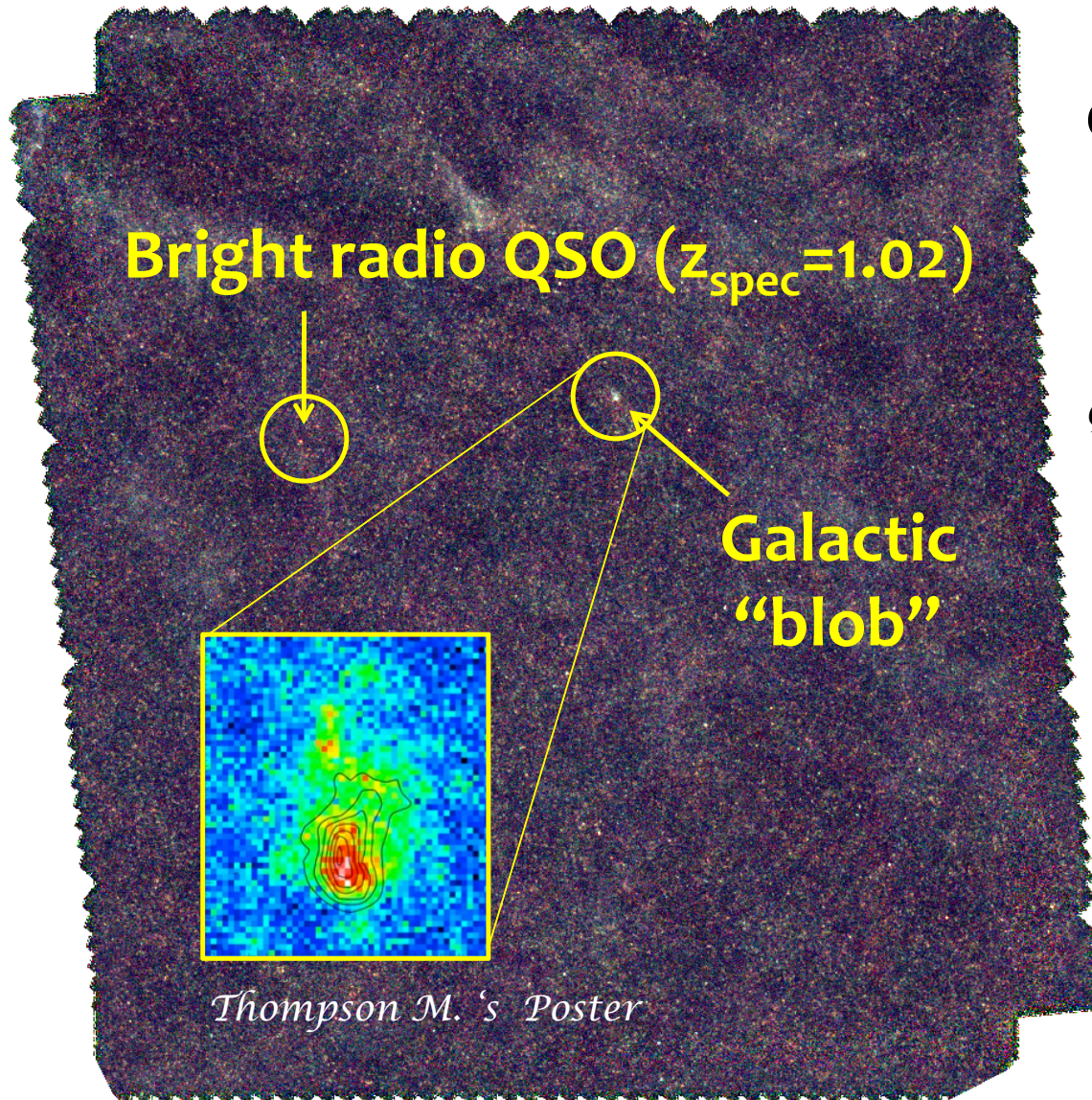
➤  $\sim 7000$  sources

*see posters by E. Rigby, D. Smith*



**11** sources with  
 $S_{500\mu\text{m}} > 100 \text{ mJy}$

# 500 $\mu\text{m}$ BRIGHTEST GALAXIES IN H-ATLAS SDP



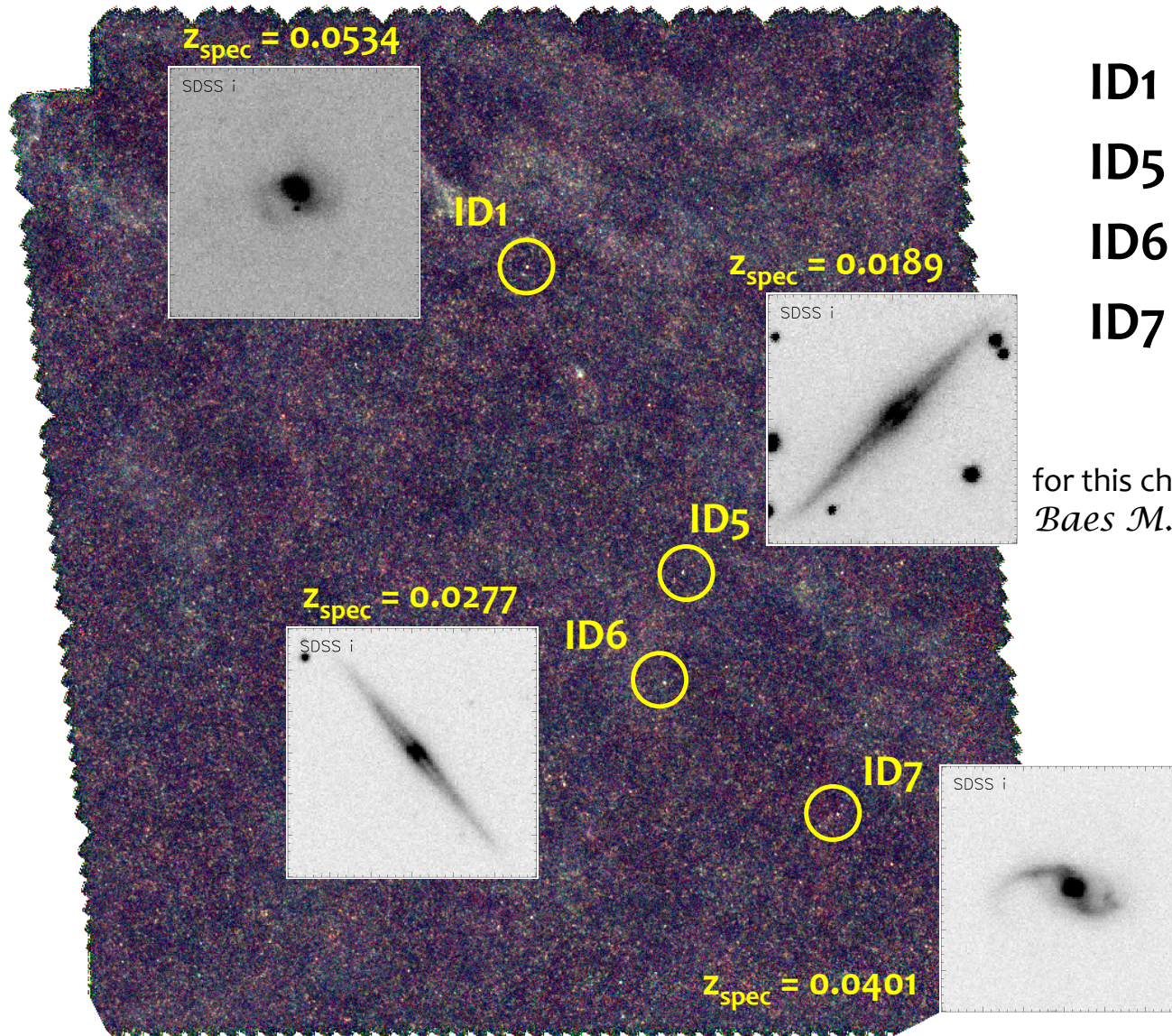
$$\begin{aligned} \text{QSO: } S_{250\mu\text{m}} &= 159.6 \text{ mJy} \\ S_{350\mu\text{m}} &= 193.8 \text{ mJy} \\ S_{500\mu\text{m}} &= 265.8 \text{ mJy} ! \\ S_{1.4\text{GHz}} &= 571.7 \text{ mJy} \end{aligned}$$

*González-Nuevo et al. Special Issue*

prediction:  $\sim 0.1 \text{ deg}^{-2}$

*De Zotti et al. (2005)*

# 500 $\mu\text{m}$ BRIGHTEST GALAXIES IN H-ATLAS SDP



$$\text{ID1} : S_{500\mu\text{m}} = 177 \pm 28 \text{ mJy}$$

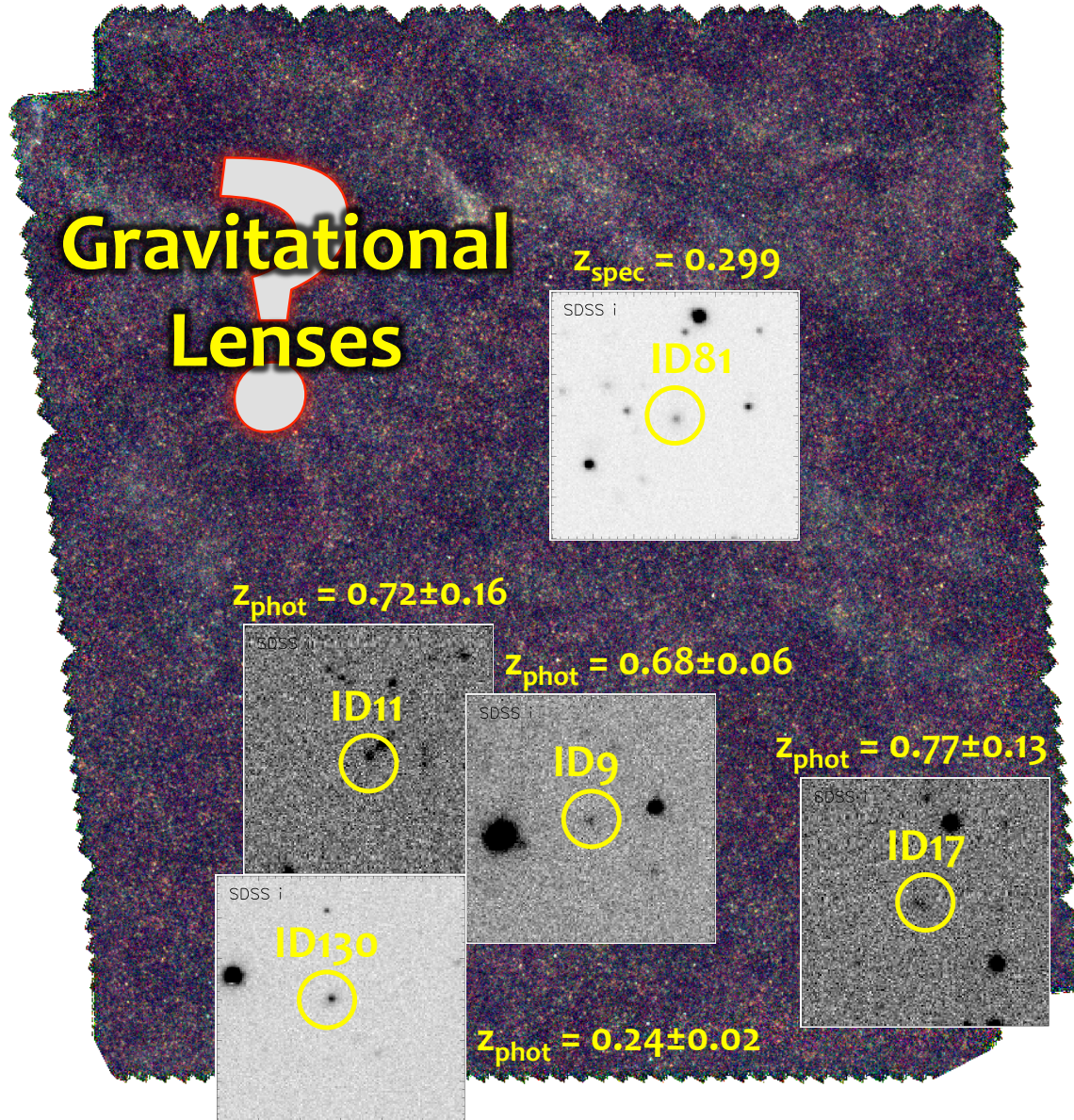
$$\text{ID5} : S_{500\mu\text{m}} = 122 \pm 20 \text{ mJy}$$

$$\text{ID6} : S_{500\mu\text{m}} = 112 \pm 19 \text{ mJy}$$

$$\text{ID7} : S_{500\mu\text{m}} = 104 \pm 18 \text{ mJy}$$

for this check  
*Baes M. et al. Special Issue*

# 500 $\mu\text{m}$ BRIGHTEST GALAXIES IN H-ATLAS SDP



ID9 :  $S_{500\mu\text{m}} = 175 \pm 28$  mJy

ID11 :  $S_{500\mu\text{m}} = 238 \pm 37$  mJy

ID17 :  $S_{500\mu\text{m}} = 220 \pm 34$  mJy

ID81 :  $S_{500\mu\text{m}} = 166 \pm 27$  mJy

ID130 :  $S_{500\mu\text{m}} = 108 \pm 18$  mJy

optical counterparts

$$z_{\text{phot/spec}} < 1.0$$

...

what about the  
sub-mm SED?

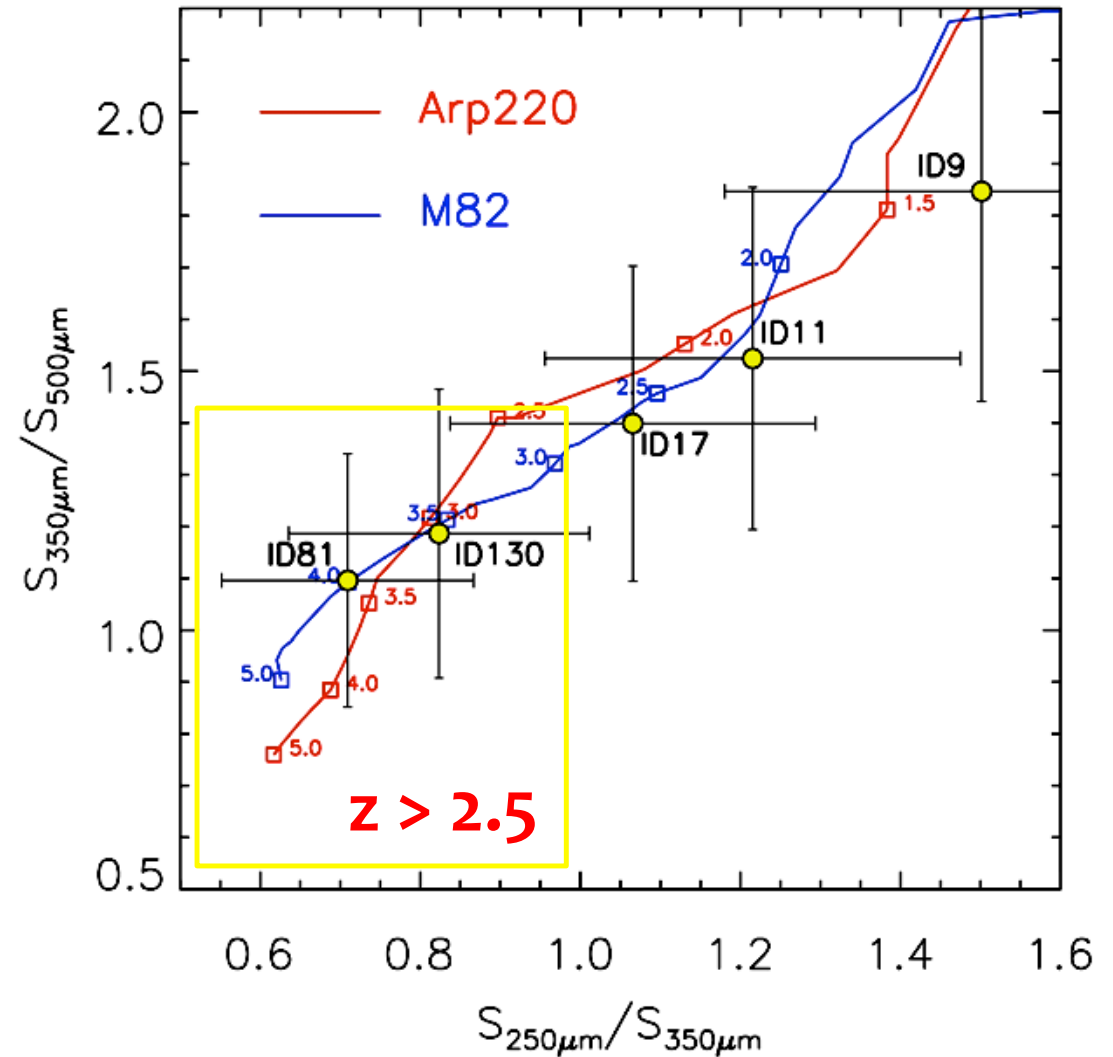


# GRAVITATIONAL LENS CANDIDATES

**ID81 - ID130:**  
UV/optical/near-IR SED  
inconsistent with  
sub-mm SED !



**best lens candidates  
for DDT follow-ups**

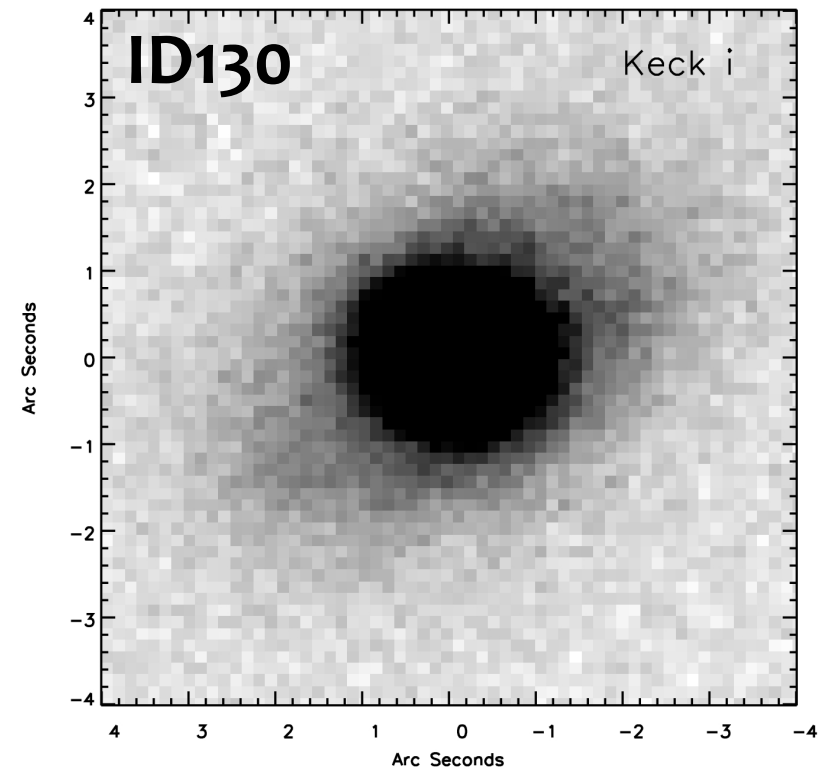
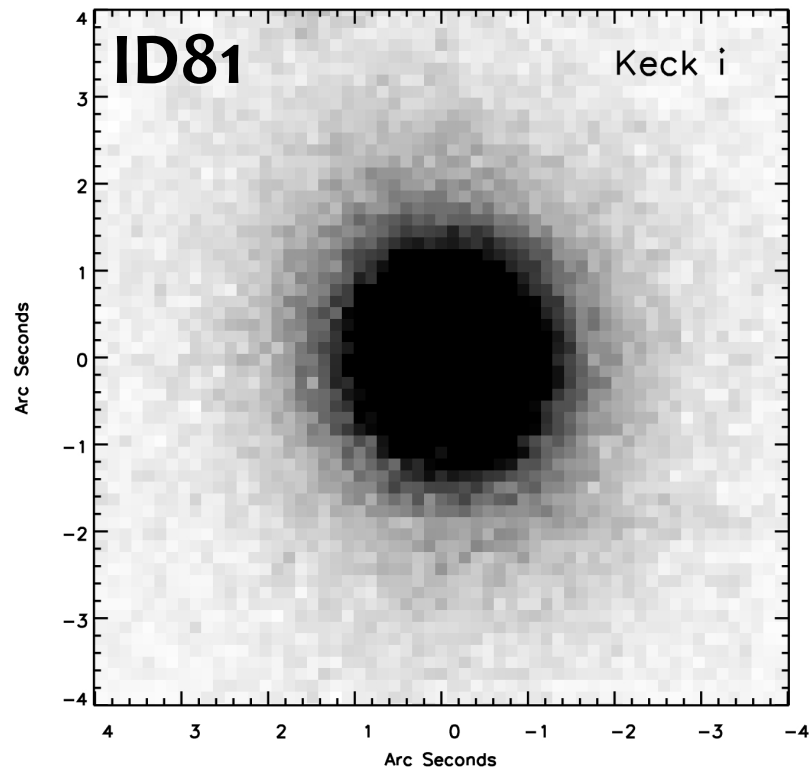


# FOLLOW-UP PROPOSALS (A LOT!!)

- **IRAM/PdBI** : high resolution mm imaging
- **CARMA** : high resolution mm imaging
- **SMA** : very-extended + compact + sub-compact imaging at 870  $\mu\text{m}$
- **Keck** : imaging g & i bands
- **VLT** : NIR high-resolution imaging
- **SPIRE FTS + PACS** minipas
- **VLT** : NIR imaging
- **CSO/Z-spec** : CO lines
- **GBT/Zspectrometer** : CO lines
- **PdBI** : CO lines
- **SCUBA-2** : 450 & 850  $\mu\text{m}$
- **MAMBO** : 1.2 mm flux for all the 5 sources

# GRAVITATIONAL LENS CANDIDATES ID81 – ID130

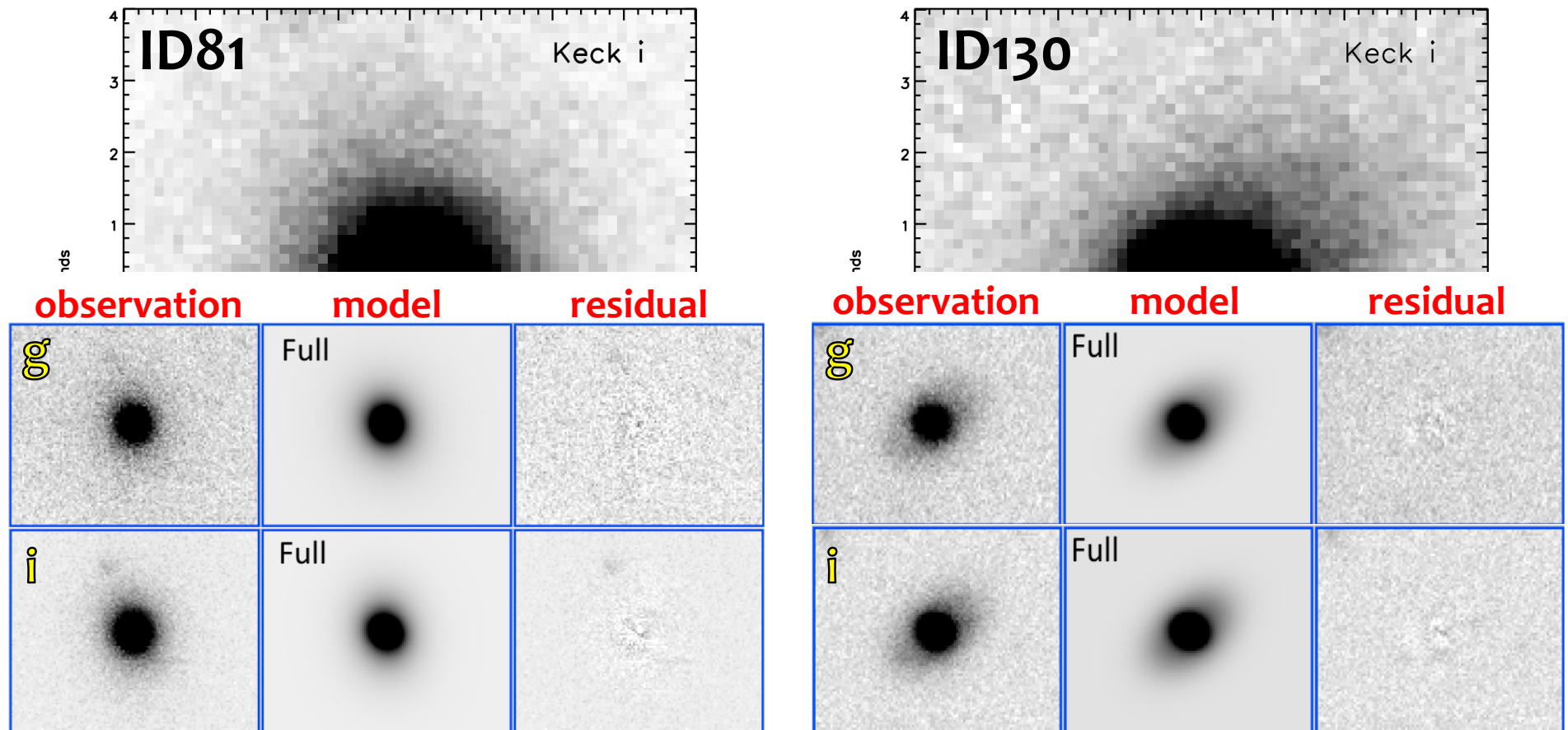
Keck imaging in g and i bands



CREDITS: Betsy Barton, Jeff Cooke, Asantha Corray, Sam Kim

# GRAVITATIONAL LENS CANDIDATES ID81 – ID130

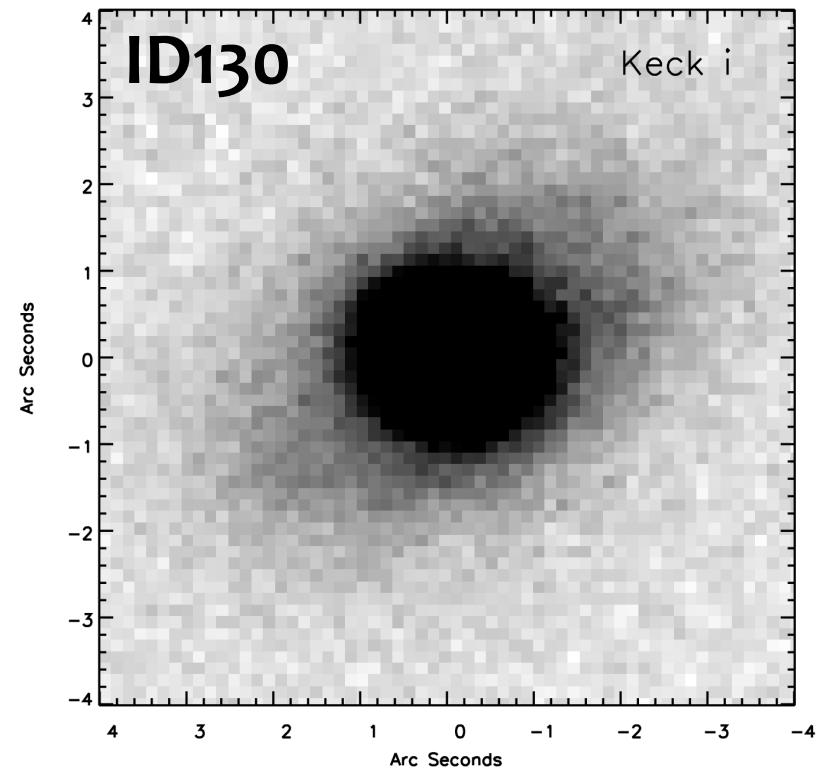
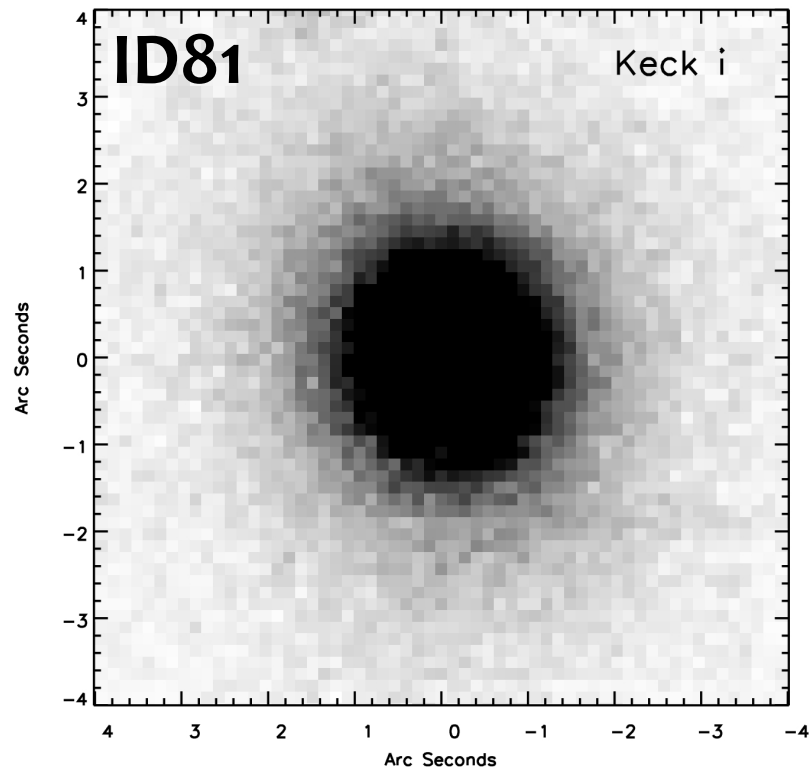
Keck imaging in g and i bands



CREDITS: Rosalind Hopwood (OU), Sam Kim (Irvine U.)

# GRAVITATIONAL LENS CANDIDATES ID81 – ID130

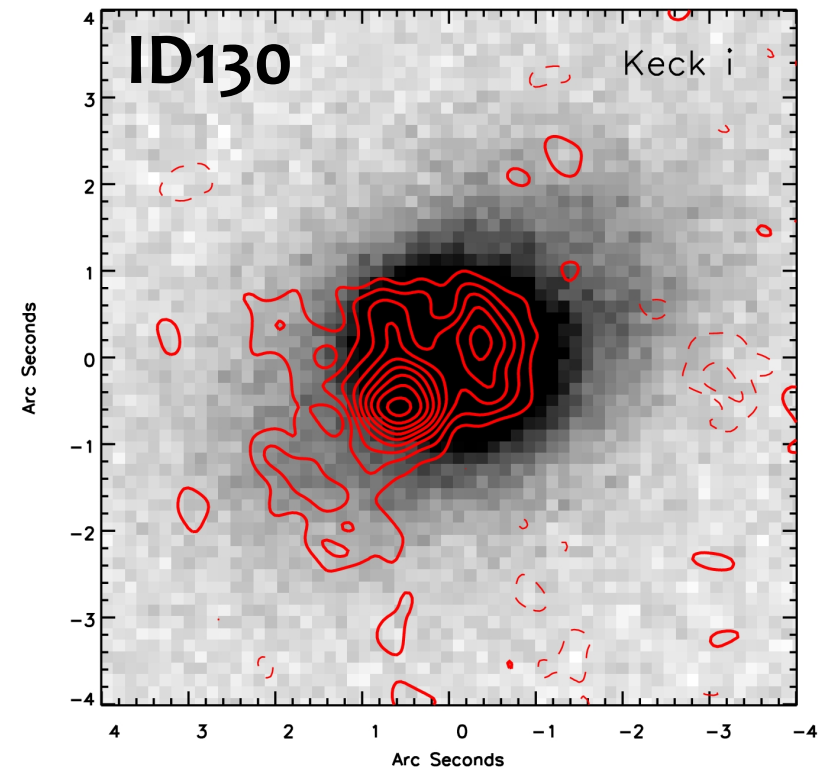
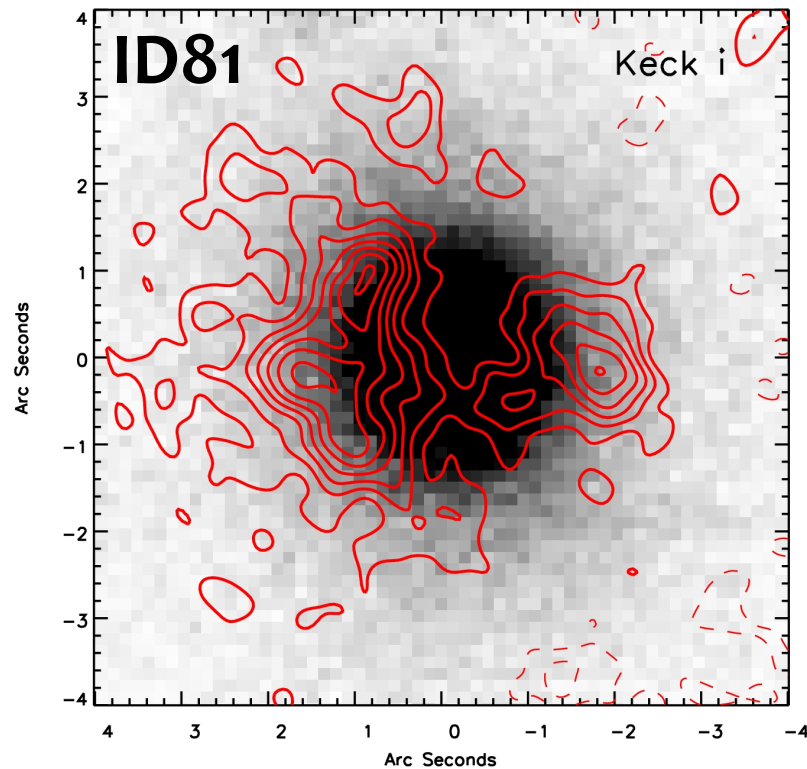
Keck imaging in g and i bands



CREDITS: Betsy Barton, Jeff Cooke, Asantha Corray, Sam Kim

# GRAVITATIONAL LENS CANDIDATES ID81 – ID130

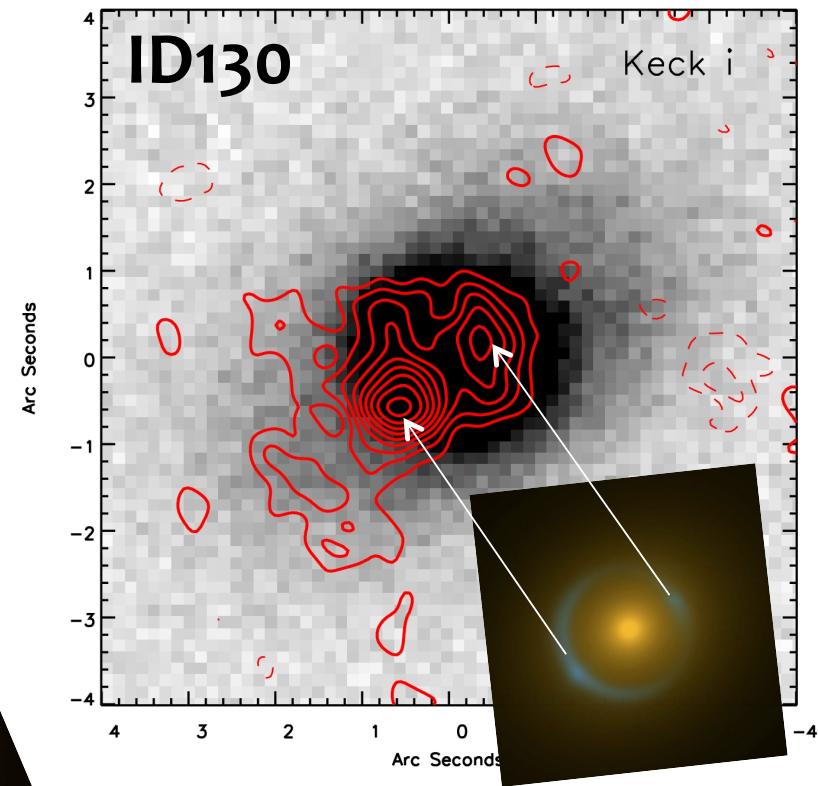
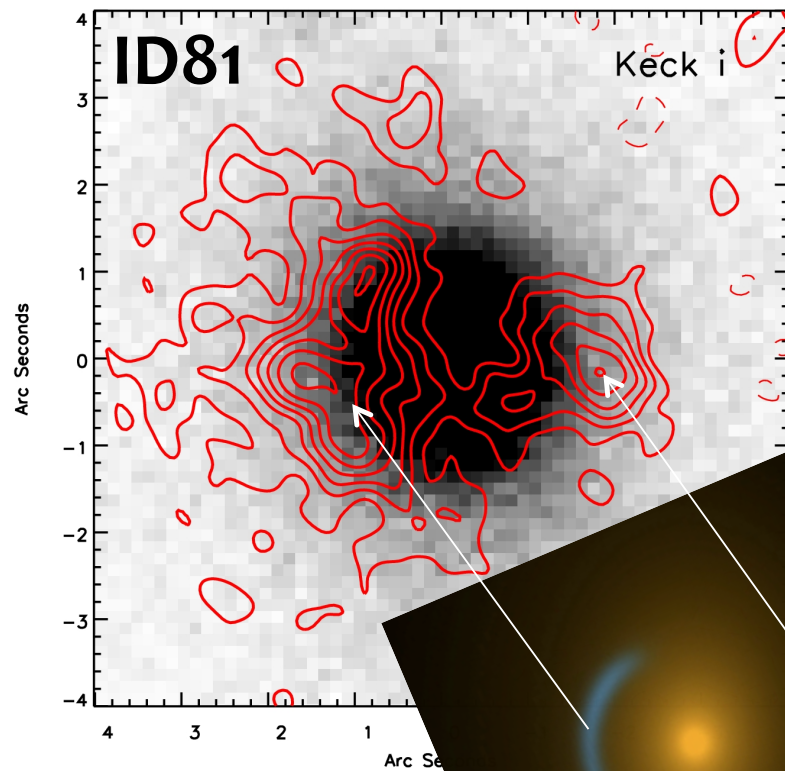
Sub Millimeter Array follow-up at  $870\ \mu\text{m}$   
(very-extended, sub-compact and compact configurations)



CREDITS: Mark Gurwell (CfA)

# GRAVITATIONAL LENS CANDIDATES ID81 – ID130

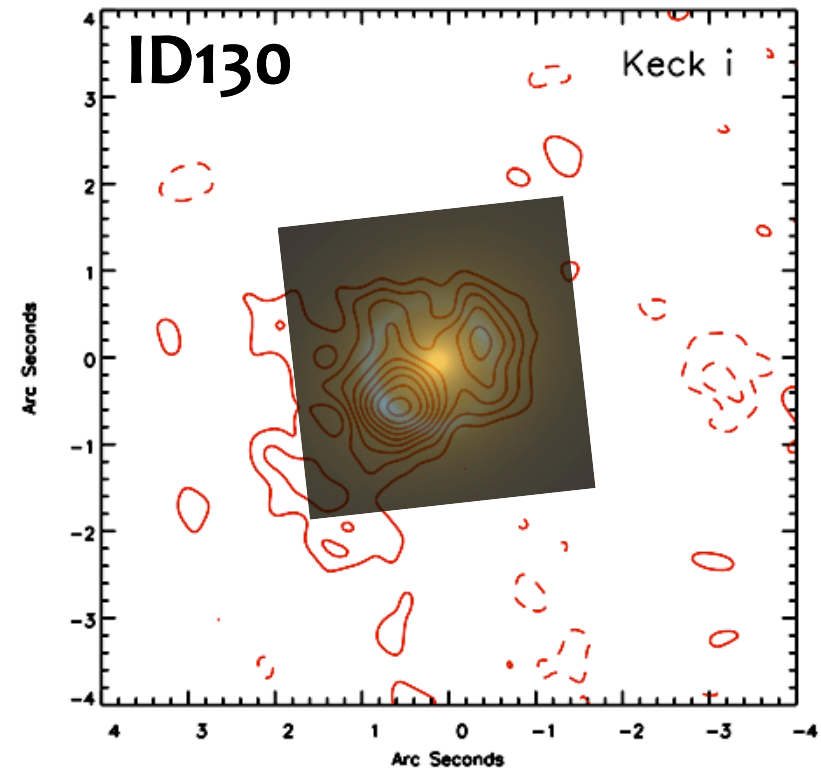
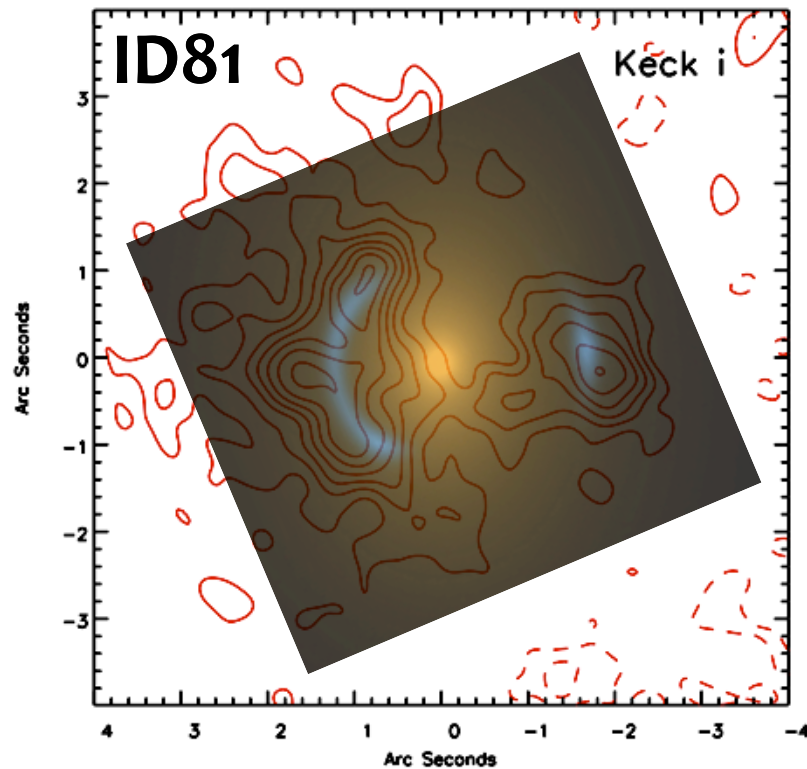
Sub Millimeter Array follow-up at  $870\ \mu\text{m}$   
(very-extended, sub-compact and compact configurations)



from <http://www.slacs.org>

# GRAVITATIONAL LENS CANDIDATES ID81 – ID130

Sub Millimeter Array follow-up at  $870\ \mu\text{m}$   
(very-extended, sub-compact and compact configurations)

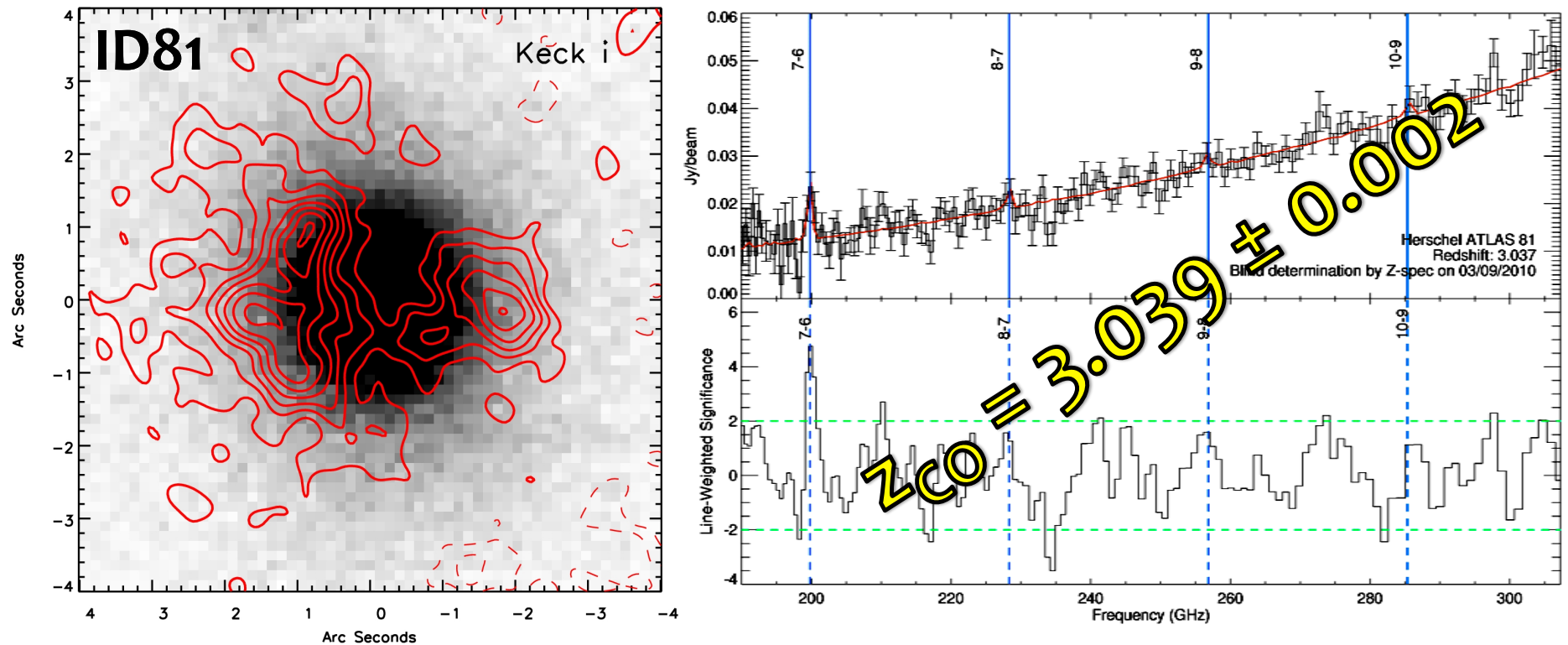


from <http://www.slacs.org>



# GRAVITATIONAL LENS CANDIDATES ID81

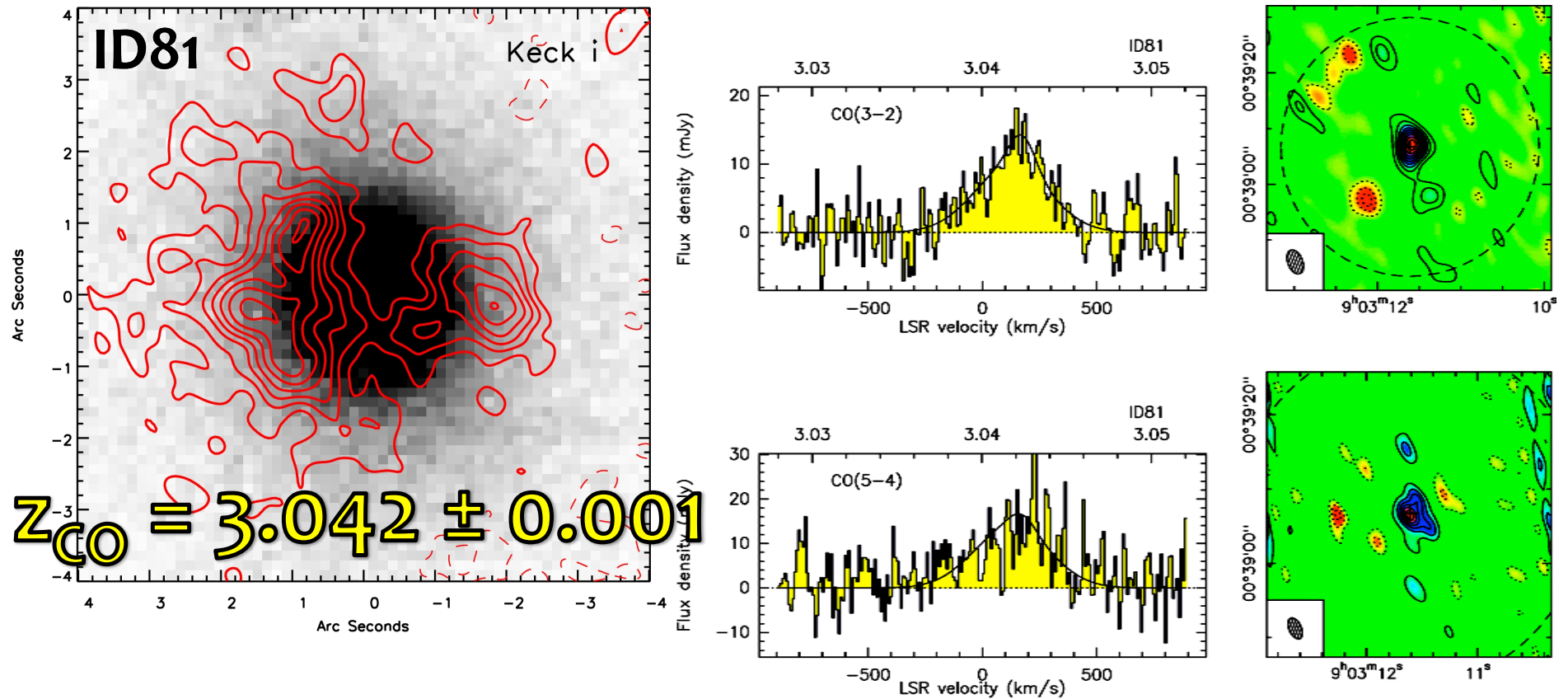
CSO/Z-spec blind redshift determination for ID81 (March 09 2010)  
from observations of the CO ladder



CREDITS: J. E. Aguirre, J. Bock, C. M. Bradford, L. Earle, J. Glenn, J. R. Kamenetzky, R. E. Lupu, P. Maloney, E. Murphy, H. Matsuhara, B. Naylor, H. T. Nguyen, K. S. Scott, J. Zmuidzinas

# GRAVITATIONAL LENS CANDIDATES ID81

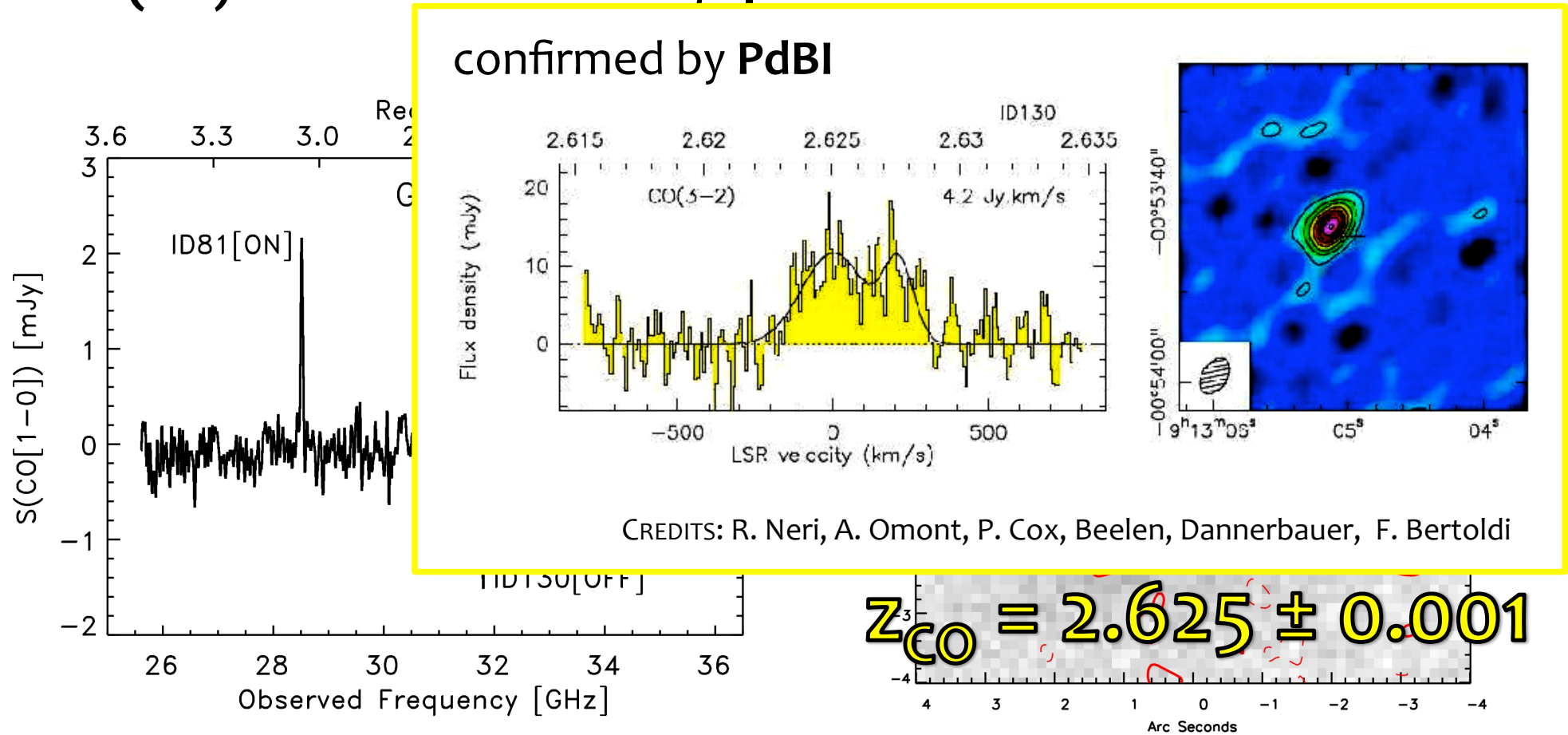
Redshift confirmed by **follow-ups** with the **PdB Interferometer** (March 23 2010) and **GBT/Zspectrometer** (March 25 2010)



CREDITS: R. Neri, A. Omont, P. Cox, Beelen, Dannerbauer, F. Bertoldi

# GRAVITATIONAL LENS CANDIDATES ID130

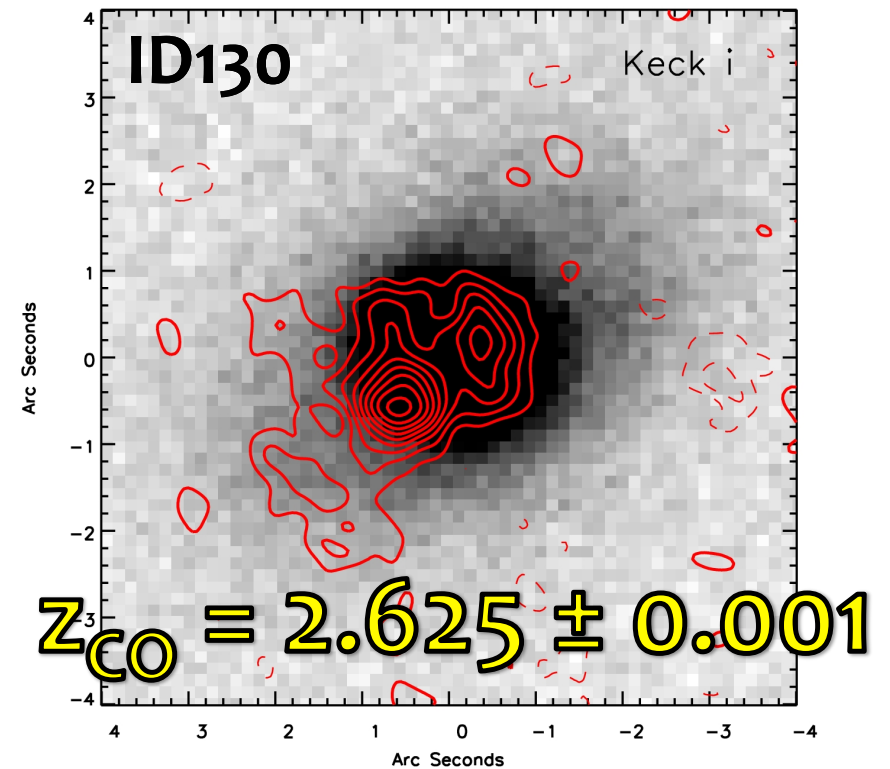
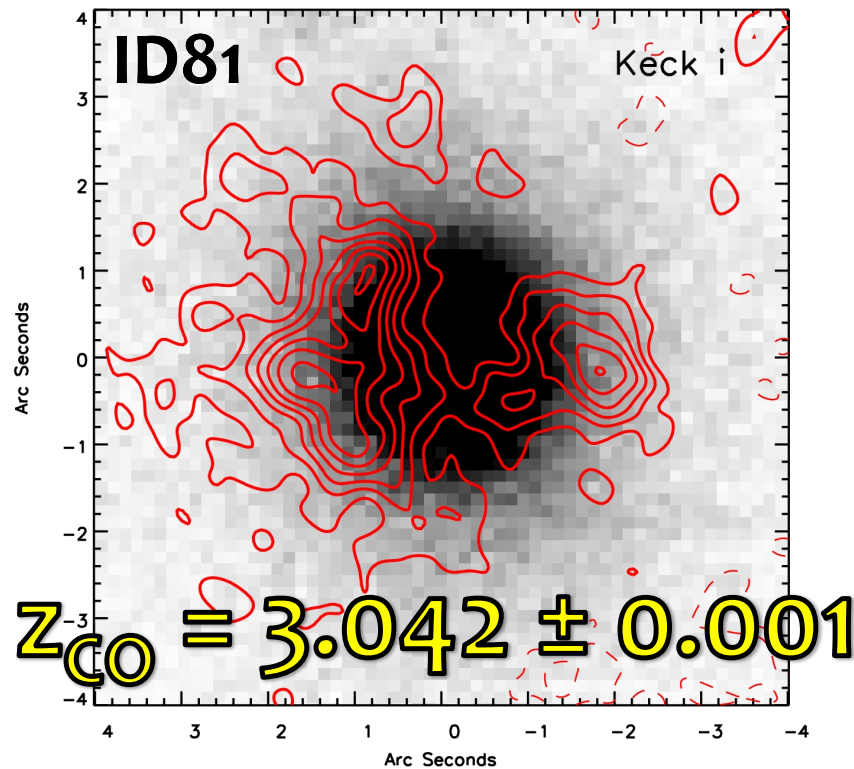
## CO(1-0) redshift from GBT/Zpectrometer



CREDITS: D. Frayer, A. Harris, A. Baker, R. Maddalena et al.

# GRAVITATIONAL LENS CANDIDATES ID81 – ID130

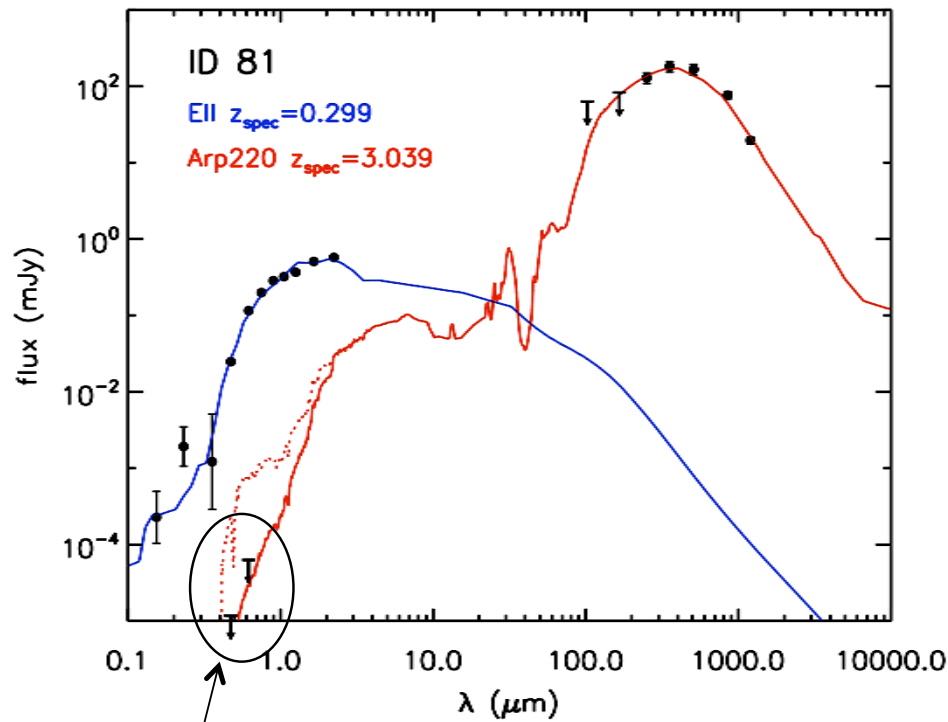
CSO/Z-spec + GBT/Zspectrometer + PdBI  $\longrightarrow$   $z_{CO}$



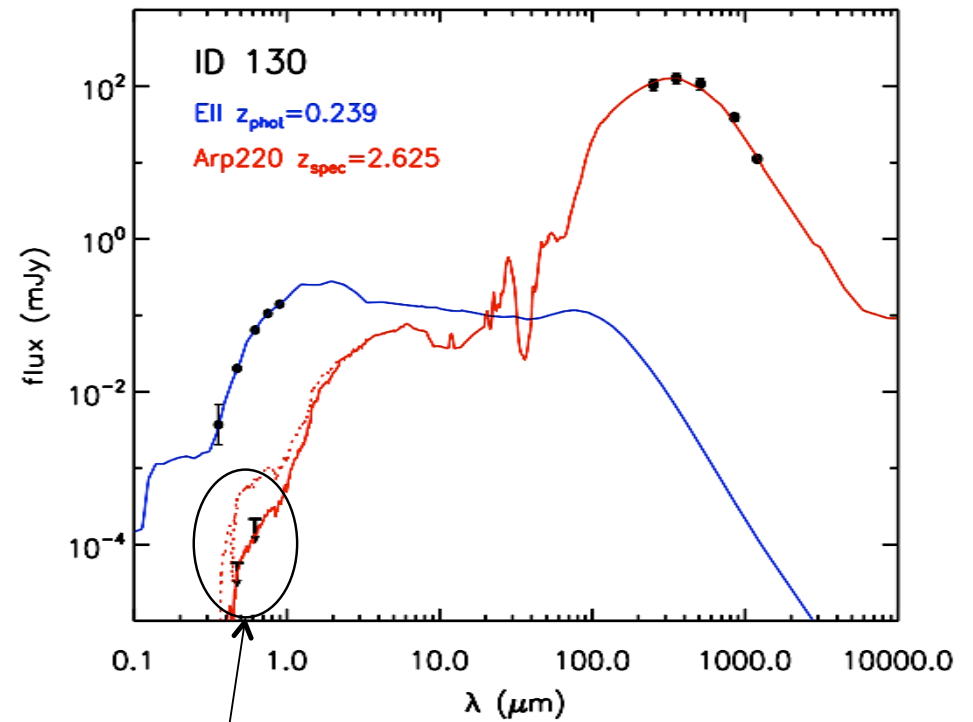
**! STRONG GRAVITATIONAL LENS EVENTS !**

# GRAVITATIONAL LENSES ID81 – ID130

These systems are **missed in the optical** !



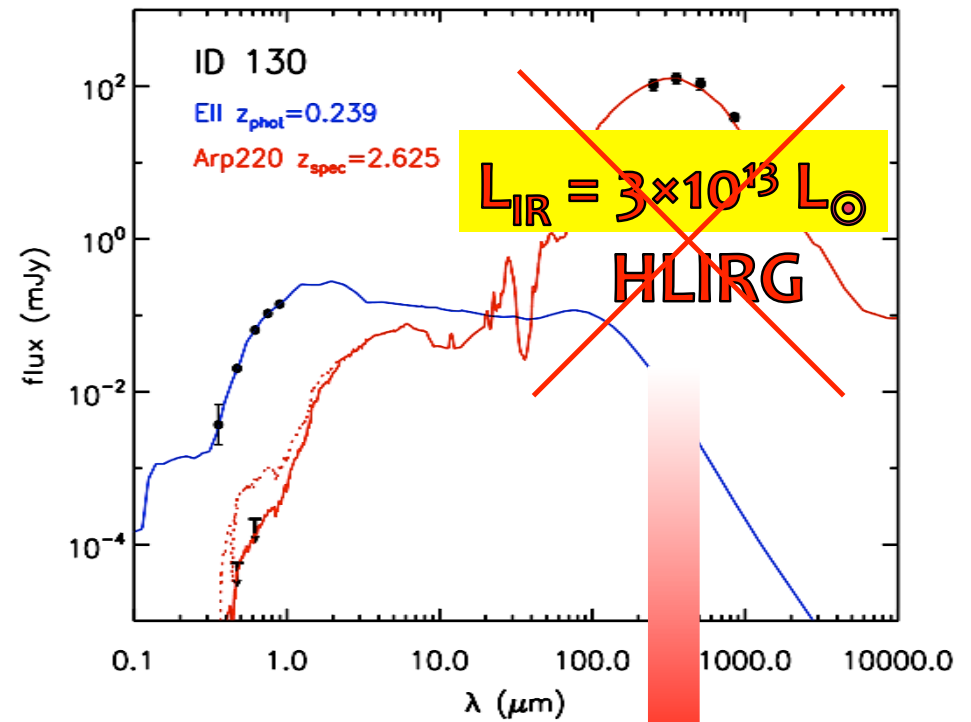
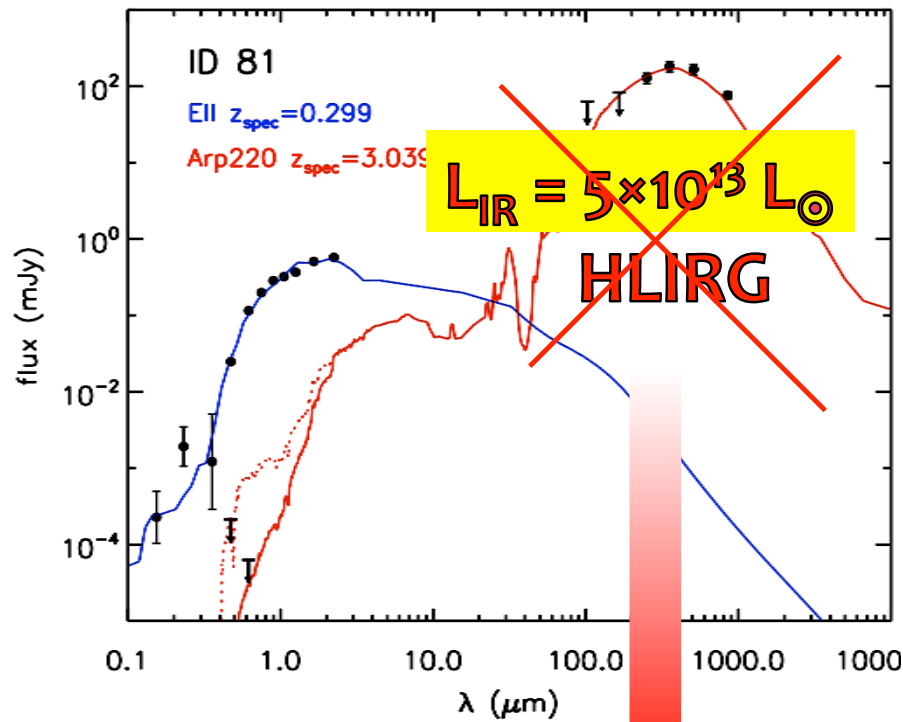
Keck 3 $\sigma$  upper limits



Keck 3 $\sigma$  upper limits

# GRAVITATIONAL LENSES ID81 – ID130

These systems are missed in the optical !



$L_{\text{IR}} < 10^{13} L_{\odot}$

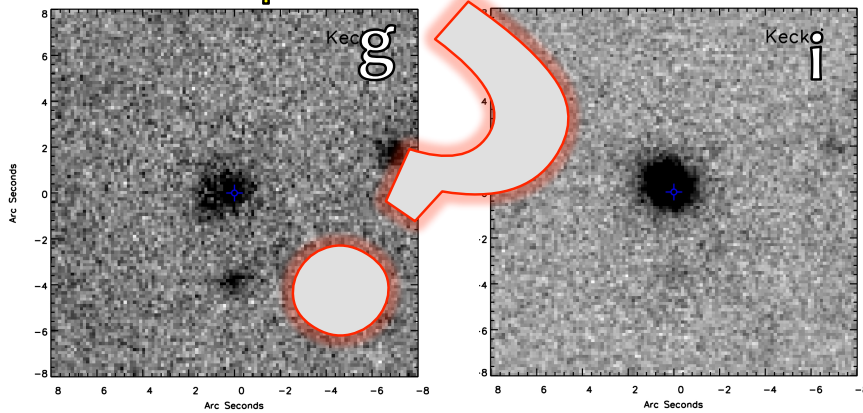
**LIRGs/ULIRGs**

$L_{\text{IR}} < 10^{13} L_{\odot}$

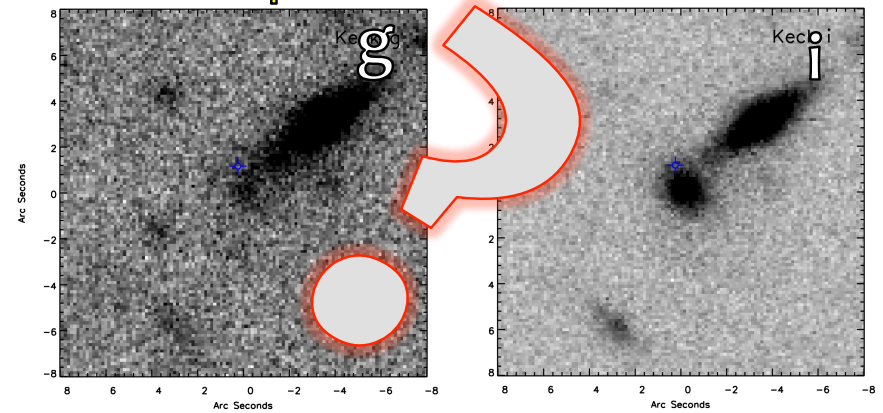
# GRAVITATIONAL LENS CANDIDATES

Keck + MAMBO + Z-spec (ongoing...)

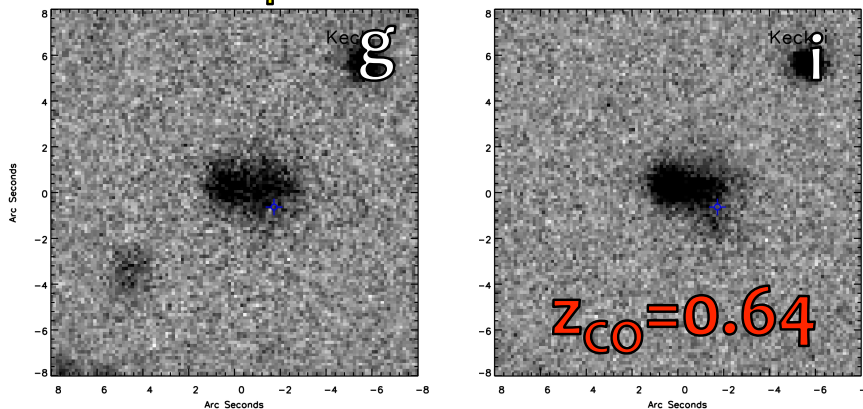
ID9  $z_{\text{phot}}=0.68\pm0.06$



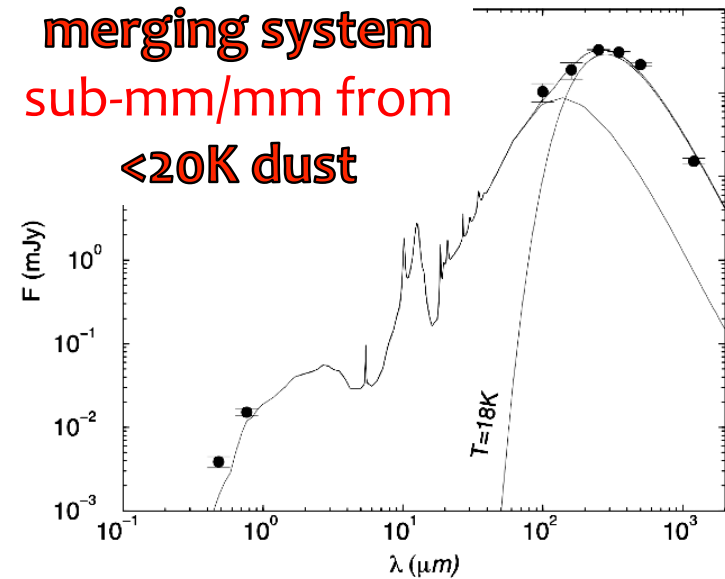
ID11  $z_{\text{phot}}=0.72\pm0.16$



ID17  $z_{\text{phot}}=0.77\pm0.13$



merging system  
sub-mm/mm from  
<20K dust



# CONCLUSIONS



ASTROPHYSICAL  
TERAHERTZ  
LARGE  
AREA  
SURVEY



discovery of the  
**Sub-mm Brightest  
Galaxies**

## STRONG GRAVITATIONAL

### LENS SYSTEMS

**proof of concept**  
no serendipitous  
discovery

**lenses invisible**  
in the **optical**

a hundred of  
**bright lenses**  
in full H-ATLAS

**blind survey**  
no cluster field

**galaxy-galaxy**  
lensing events

**very high**  
lens selection  
**efficiency**