

The Herschel Lensing Survey (HLS)

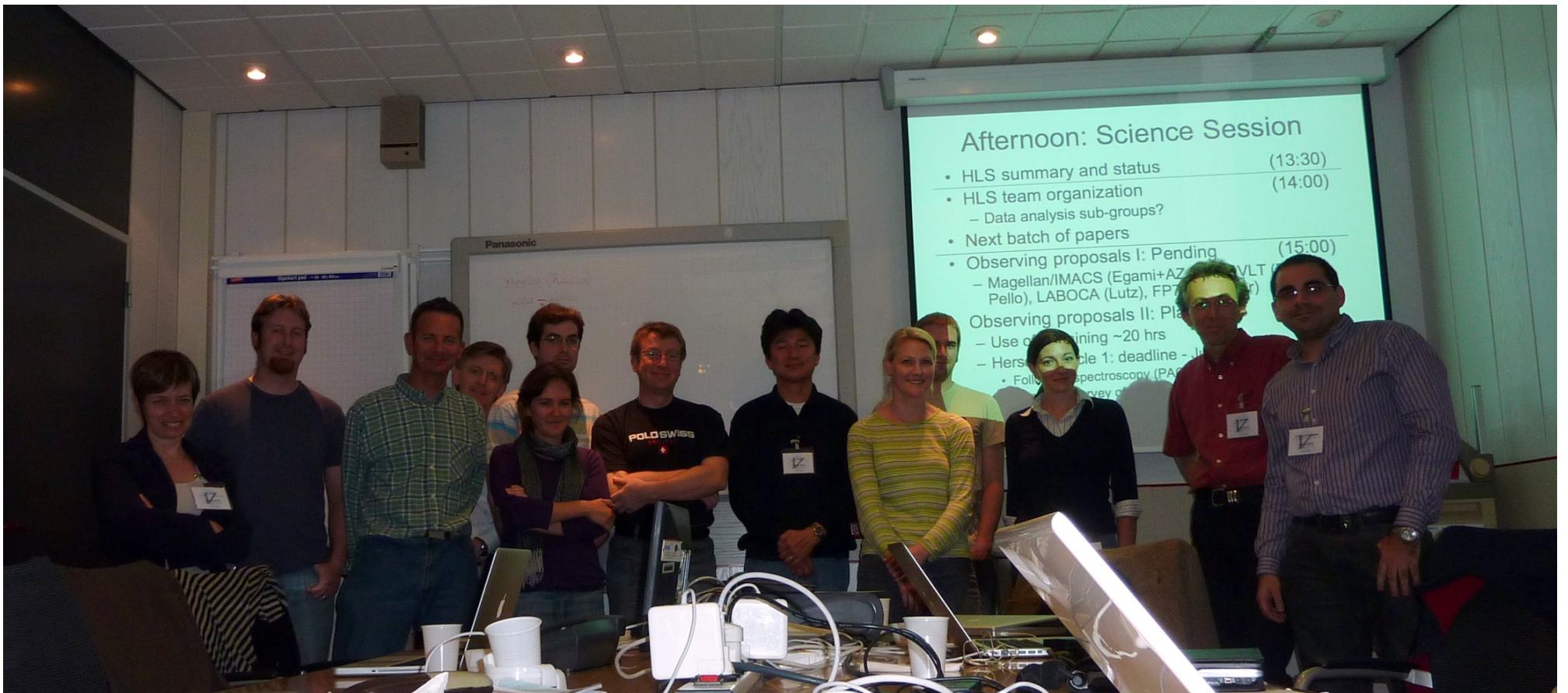
Eiichi Egami

Steward Observatory
University of Arizona

HLS Team Members

(Total: 35, US: 14, ESA 21)

E. Egami (PI)	Arizona	D. Dowell	JPL/Caltech	J. Richard	Durham
M. Rex	Arizona	M. Dessauges-Zavadsky	Geneva	G. Rieke	Arizona
T. Rawle	Arizona	D. Fadda	NHSC	G. Rodighiero	Padova
M. Pereira	Arizona	O. Ilbert	Marseille	D. Schaerer	Geneva
G. Walth	Arizona	R. Ivison	Edinburgh	I. Smail	Durham
B. Altieri	Blain	M. Jauzac	Marseille	G. Smith	Birmingham
A. Blain	Caltech	J.-P. Kneib	Marseille	G. Tramoy	Marseille
J. Bock	JPL/Caltech	D. Lutz	MPE	I. Valtchanov	ESAC
F. Boone	Toulouse	L. Metcalfe	ESAC	P. Van der Werf	Leiden
C. Bridge	Caltech	A. Omont	IAP	M. Werner	JPL/Caltech
B. Clement	Marseille	R. Pello	Toulouse	M. Zemcov	JPL/Caltech
F. Combs	Paris	P. Perez-Gonzalez	Madrid		



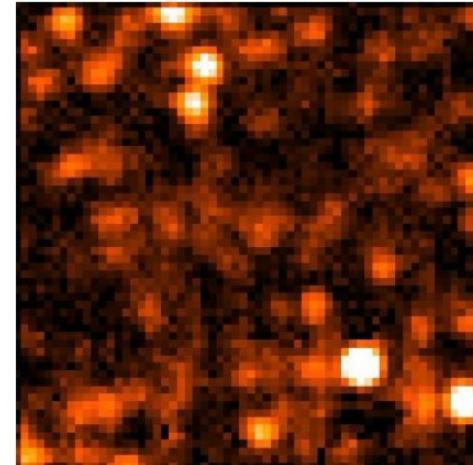
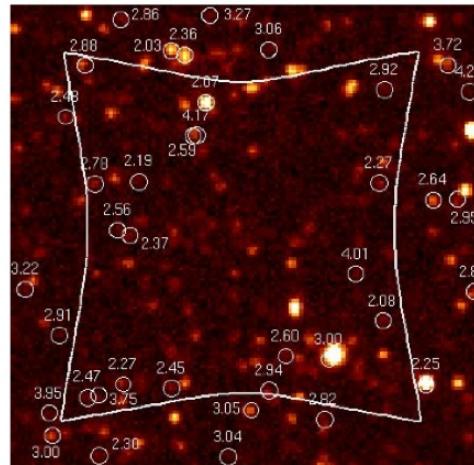
HLS Team Meeting at ESTEC
May 3, 2010

Scientific Goals

1. To detect and study IR/Submm sources that are **below the nominal confusion limit of Herschel** using the gravitational lensing power of massive galaxy clusters.
2. To study IR/submm properties of **galaxies in dense environment** (i.e., cluster members).
3. To investigate the **Sunyaev-Zel'dovich effect** (SPIRE can detect increment at 350/500 um).

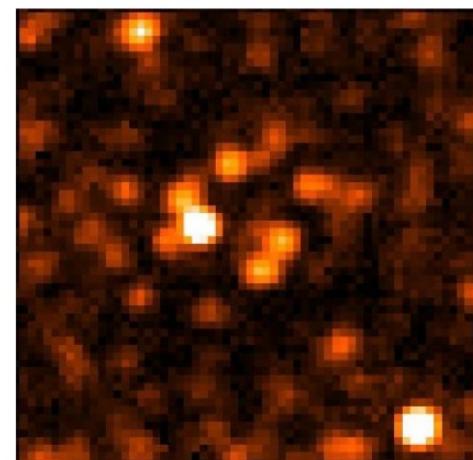
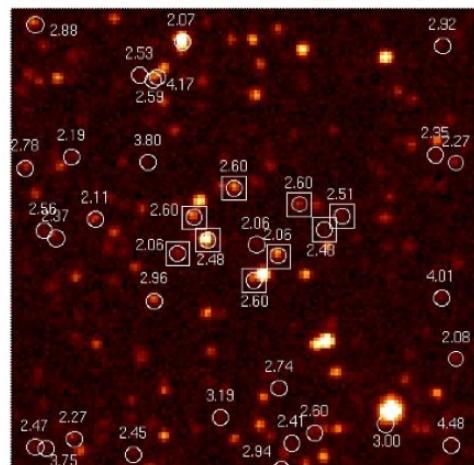
Penetrating through the Confusion with Cluster Cosmic Telescopes

Without Lensing



Lensing is more important for SPIRE images, which get confusion-limited quickly.

With Lensing



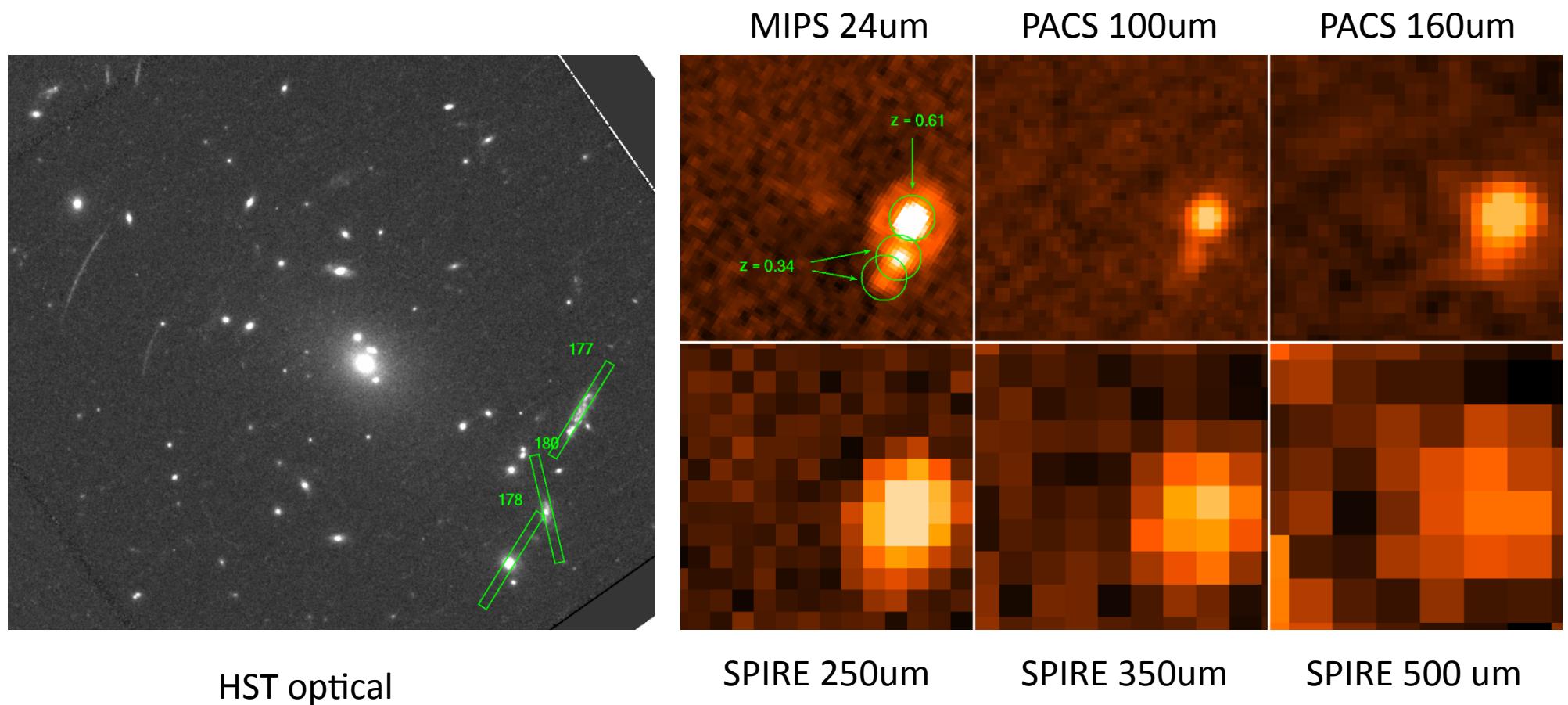
PACS 100 um

7'x7'

SPIRE 250 um

In IR/Submm, massive galaxy clusters
act as transparent lenses (well, almost...)

AS 1063 ($z = 0.34$)



See the poster by Greg Walth

Survey Strategy & Design

- Target: ~40 massive galaxy clusters (GT surveys observe 10 clusters -> OT+GT~50)
- Selected X-ray-luminous clusters with good ancillary data and cluster mass models.
- Close collaboration with two other cluster OTKPs (LoCuSS – PI: G. Smith; BCGs – PI: A. Edge)
- PACS: 100/160um; FOV 8'x8'; 5/10 mJy (5 σ)
- SPIRE: 250/350/500 um; FOV 17'x17'; confusion limit ~30 mJy (5 σ)
- Total observing time: 292.3 hrs

Five Herschel Special-Issue Papers on the Bullet Cluster (SDP target)

- The Herschel Lensing Survey: Overview
(Egami et al.)
- **Sources behind the Cluster:**
 - 1. Far-IR/Submm SED properties **(Rex et al.)**
 - 2. Multi-wavelength source matching/photometry and far-IR/submm phot-z's **(Perez-Gonzalez et al.)**
- **Cluster Galaxies:** 3. Far-IR/Submm properties of galaxies in dense environment **(Rawle et al.)**
- **Sunyaev-Zel'dovich effect:** 4. First detection of SZ increment at < 650 um **(Zemcov et al.)**

The Bullet Cluster: X-ray-luminous merging cluster at $z=0.3$

Special thanks to,

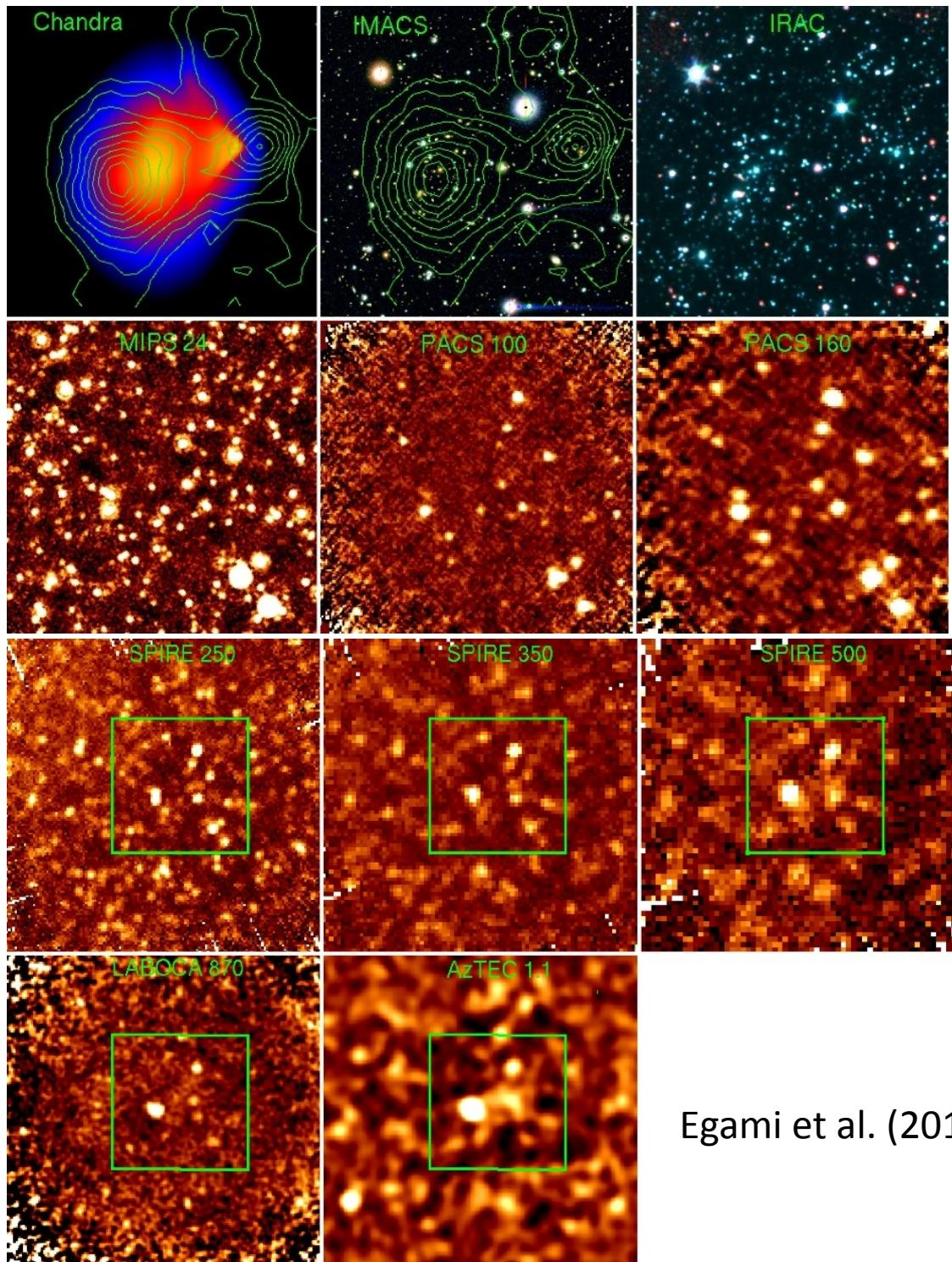
Doug Clowe
(Magellan/IMACS images)

Jean-Gabriel Cuby
(VLT/HAWKI images)

Anthony Gonzalez, Sun Mi Chung
(Magellan/IMACS redshifts)

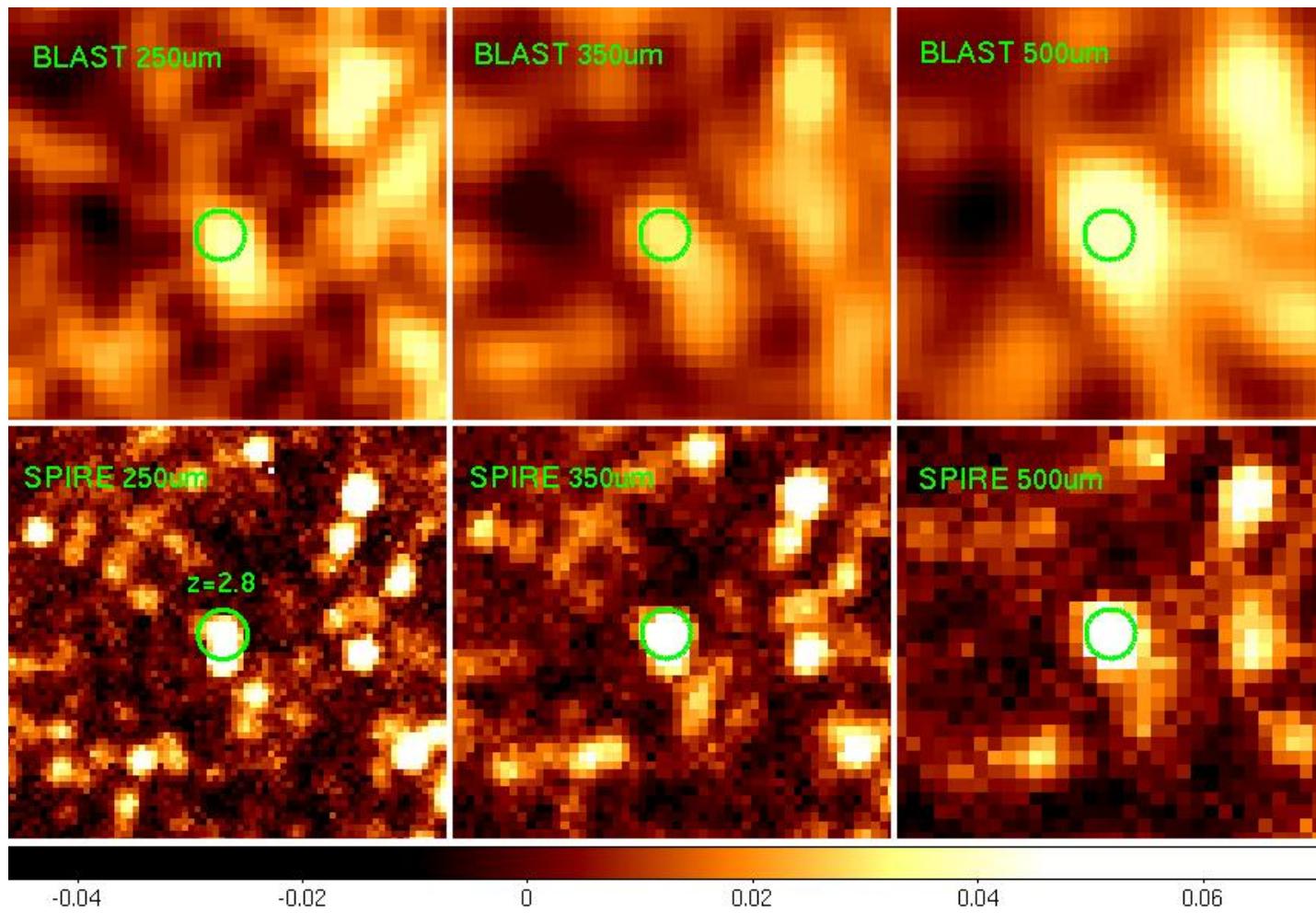
Cathy Horellou, Daniel Johansson and
LABOCA team

David Hughes, Itziar Artxaga and
AzTEC team

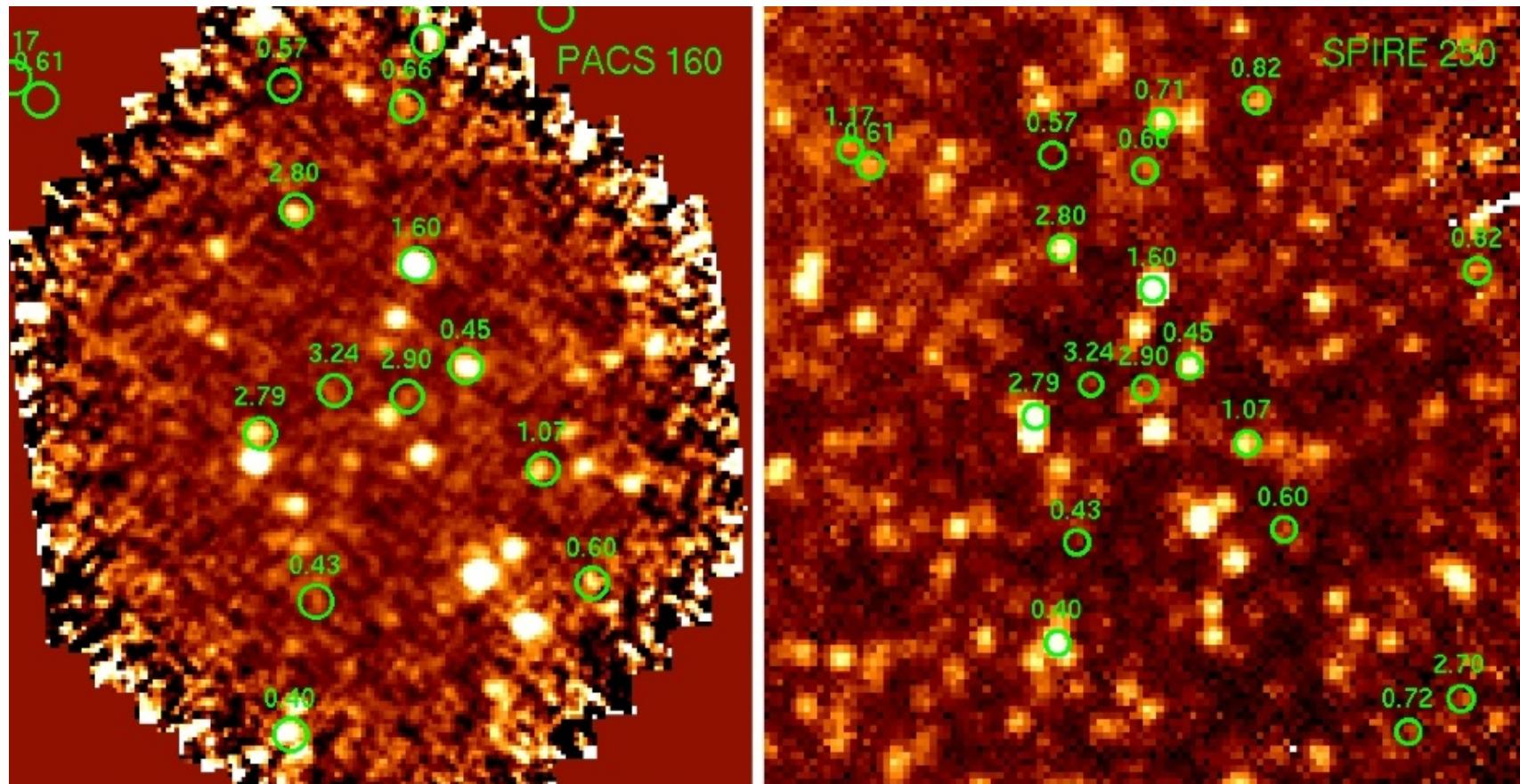


Egami et al. (2010)

BLAST vs. SPIRE



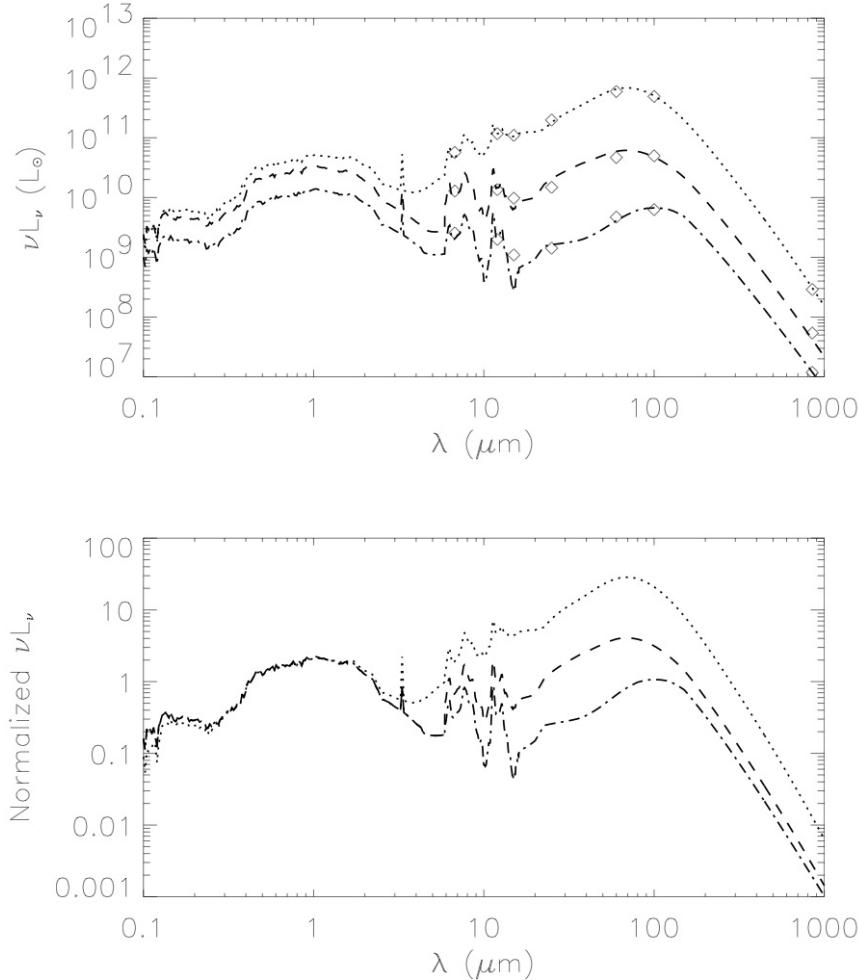
Galaxies behind the Bullet Cluster ($z > 0.4$)



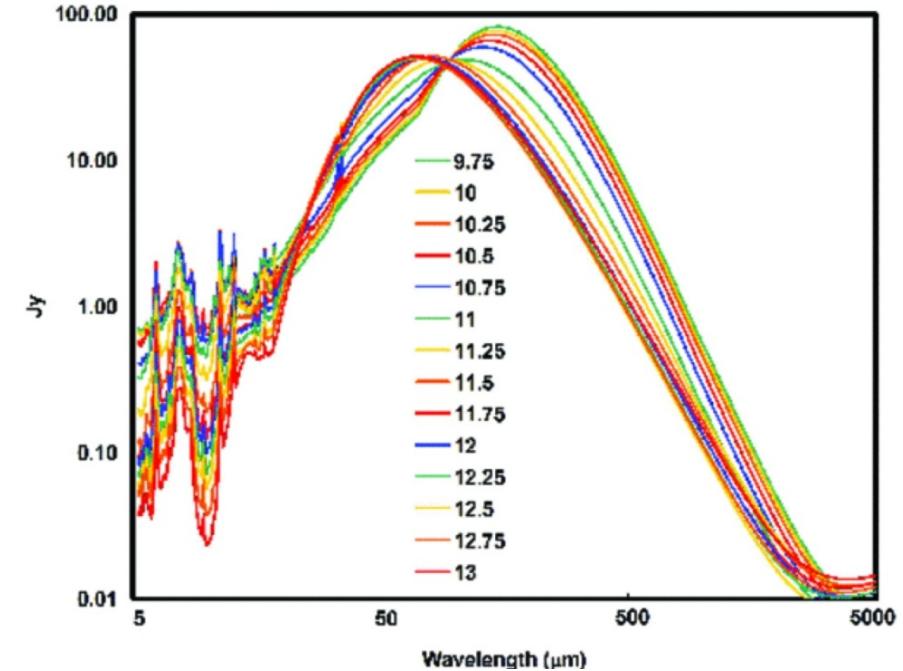
For the SDP papers, we limited the analysis to well-isolated,
• 15 sources with spectroscopic redshifts
• 4 sources with good photometric redshifts

1. Far-IR/Submm SED Properties

Rex et al. (2010); see her poster!



Chary & Elbaz (2001)

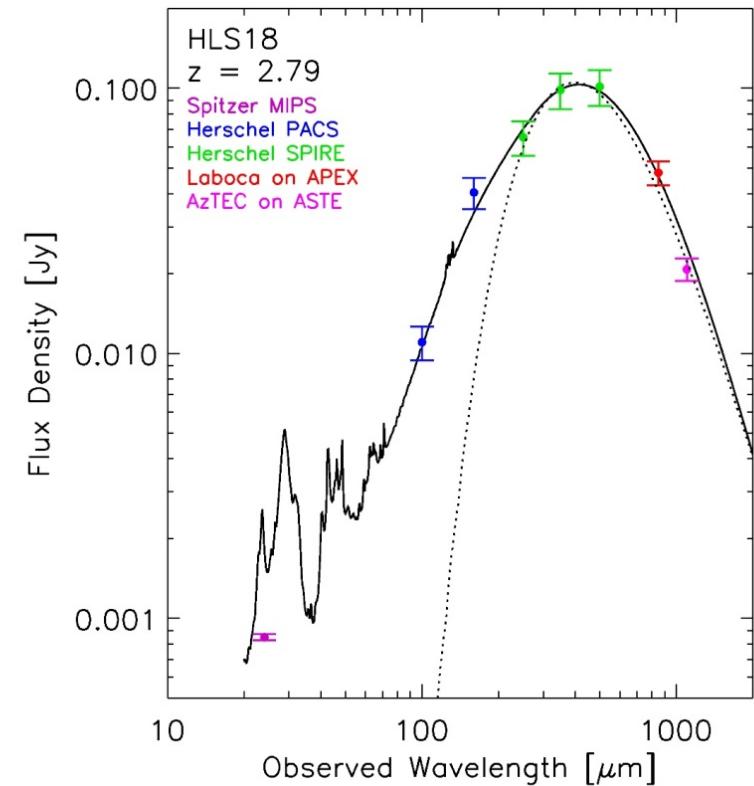
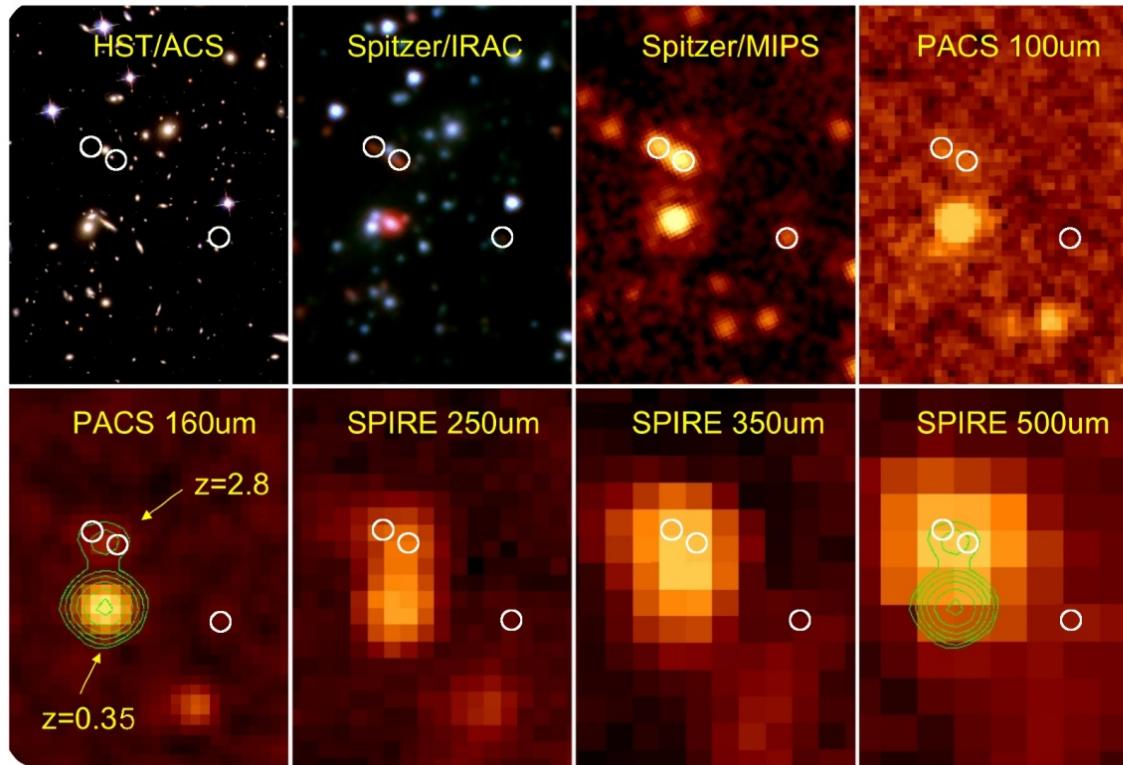


Rieke et al. (2009)

$$L_{\text{TIR}} \uparrow \rightarrow T_{\text{dust}} \uparrow \rightarrow L_{\text{FIR}}/L_{\text{MIR}} \uparrow$$

Q: Do the local galaxy SED templates work at high redshift?

Mapping the Full Far-IR/Submm SED of a ULIRG ($2 \times 10^{12} L_\odot$) at $z=2.8$!



Magnification factor $\sim x50\text{-}100$

Observed flux densities : 7.0, 24.5, 65.3, 98.6, 101.4 mJy

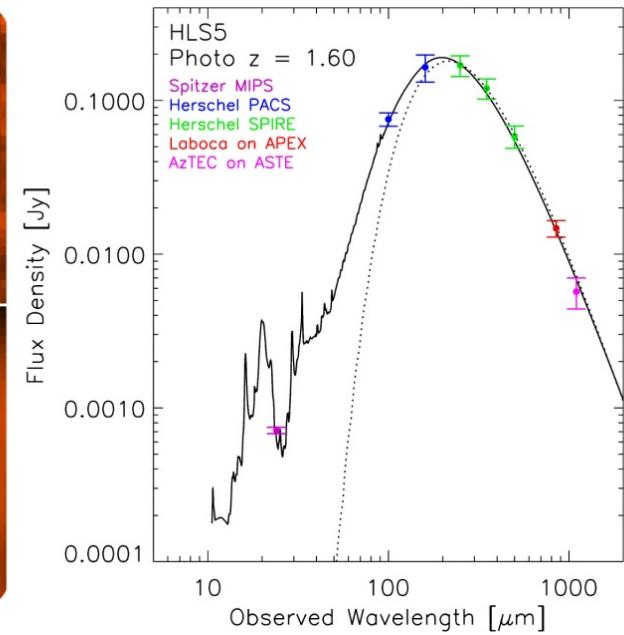
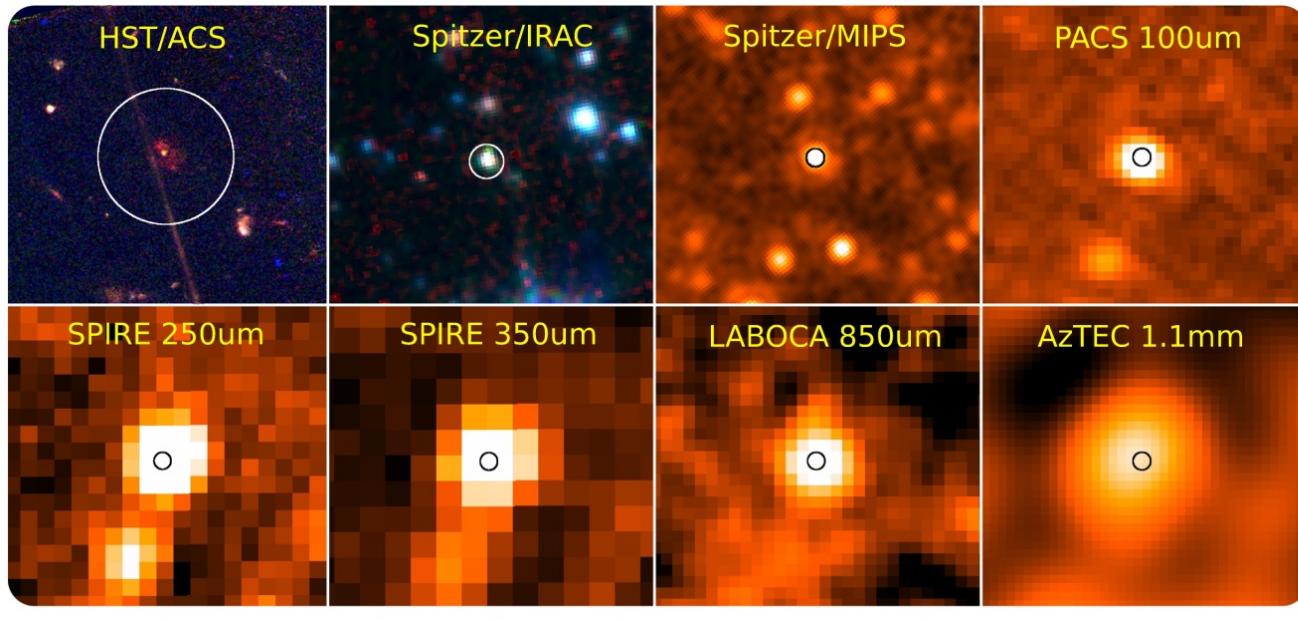
Corrected for lensing (x75): 0.09, 0.3, 0.9, 1.3, 1.4 mJy

Also, see Rex et al. (2009) for BLAST

Impossible to detect
without lensing!!

HyLIRG ($4.4 \times 10^{13} L_\odot$) at $z=1.6$

(But not significantly lensed...more like a typical SMG)

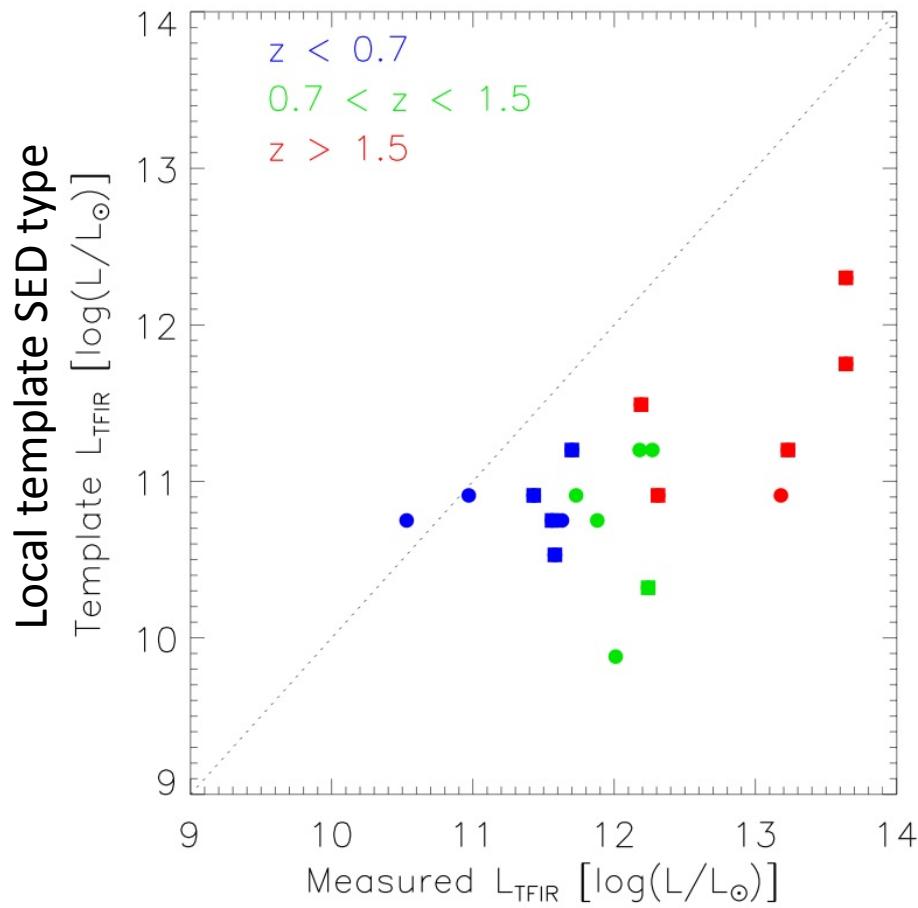


Magnification factor $\sim x1.2$

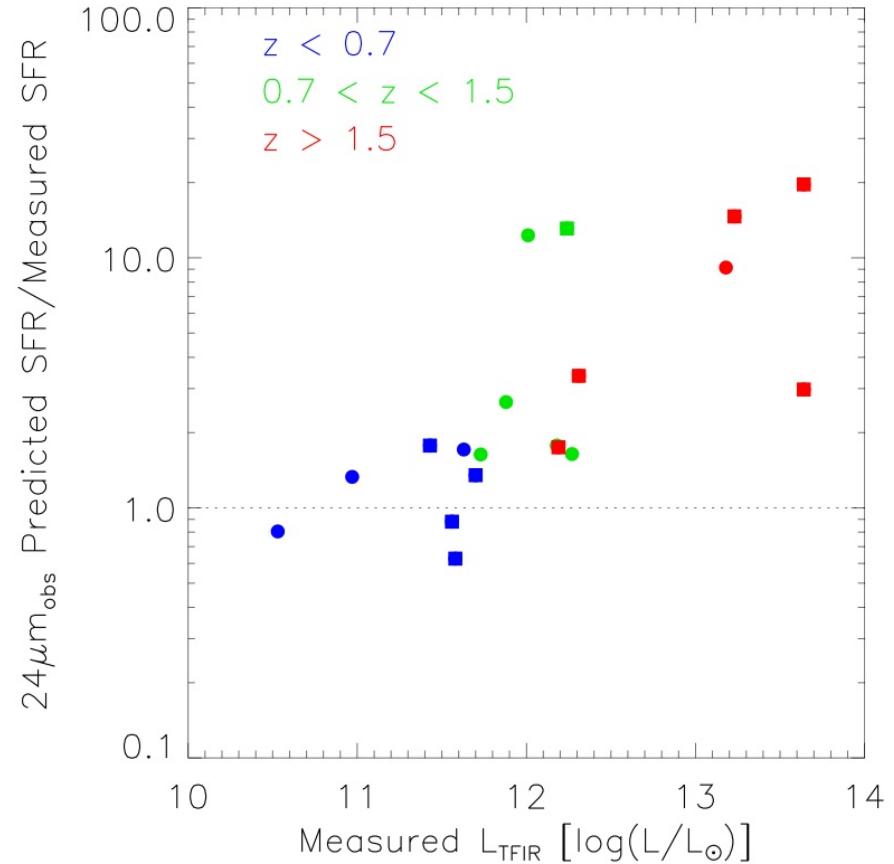
Observed flux densities : 75.4, 164.4, 168.9, 120.0, 58.4 mJy

Star-forming galaxy
SED

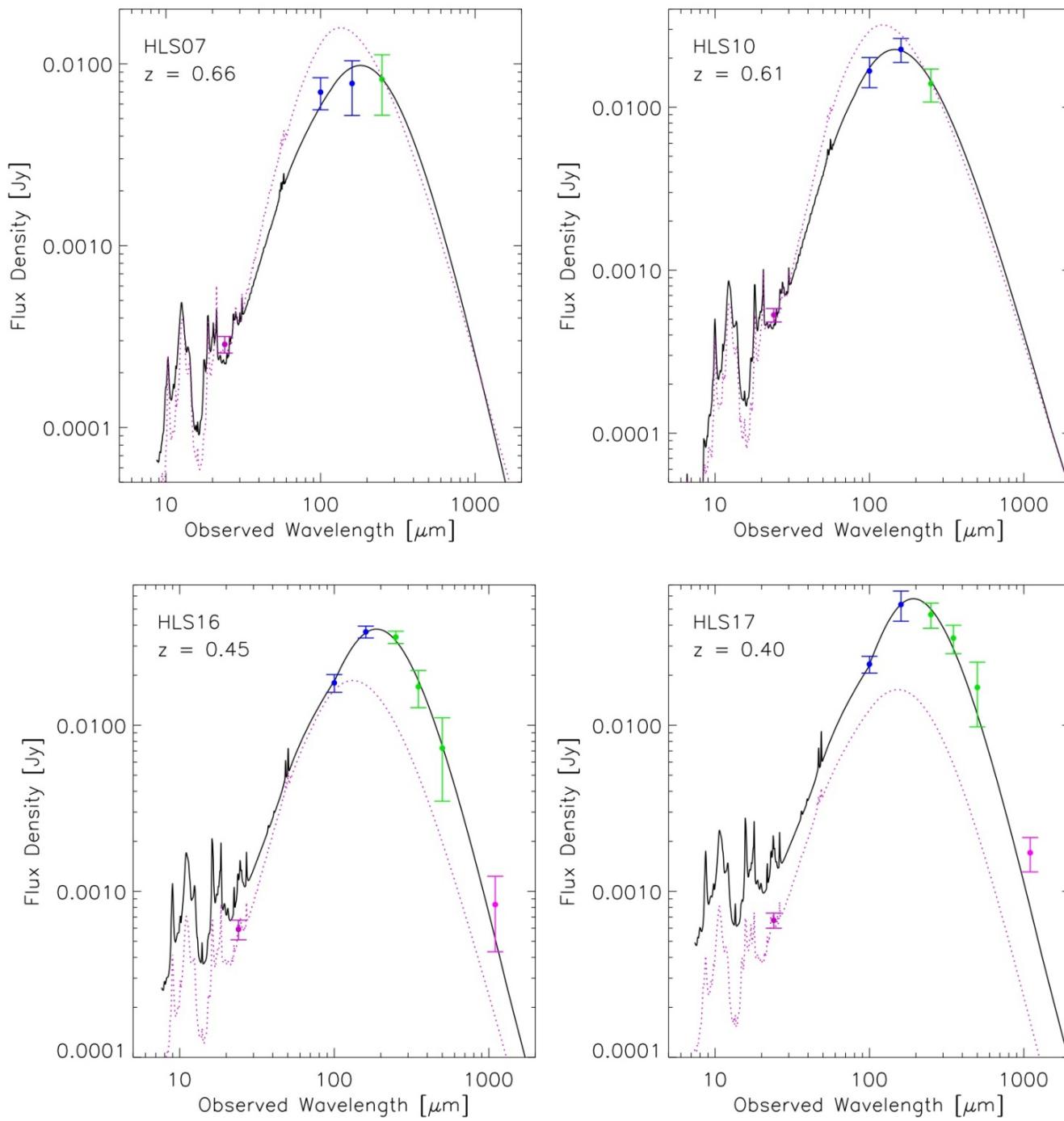
Properties of High-z Galaxy SEDs



IR/Submm SEDs of high-redshift galaxies appear colder than their local galaxies with Similar IR luminosities.



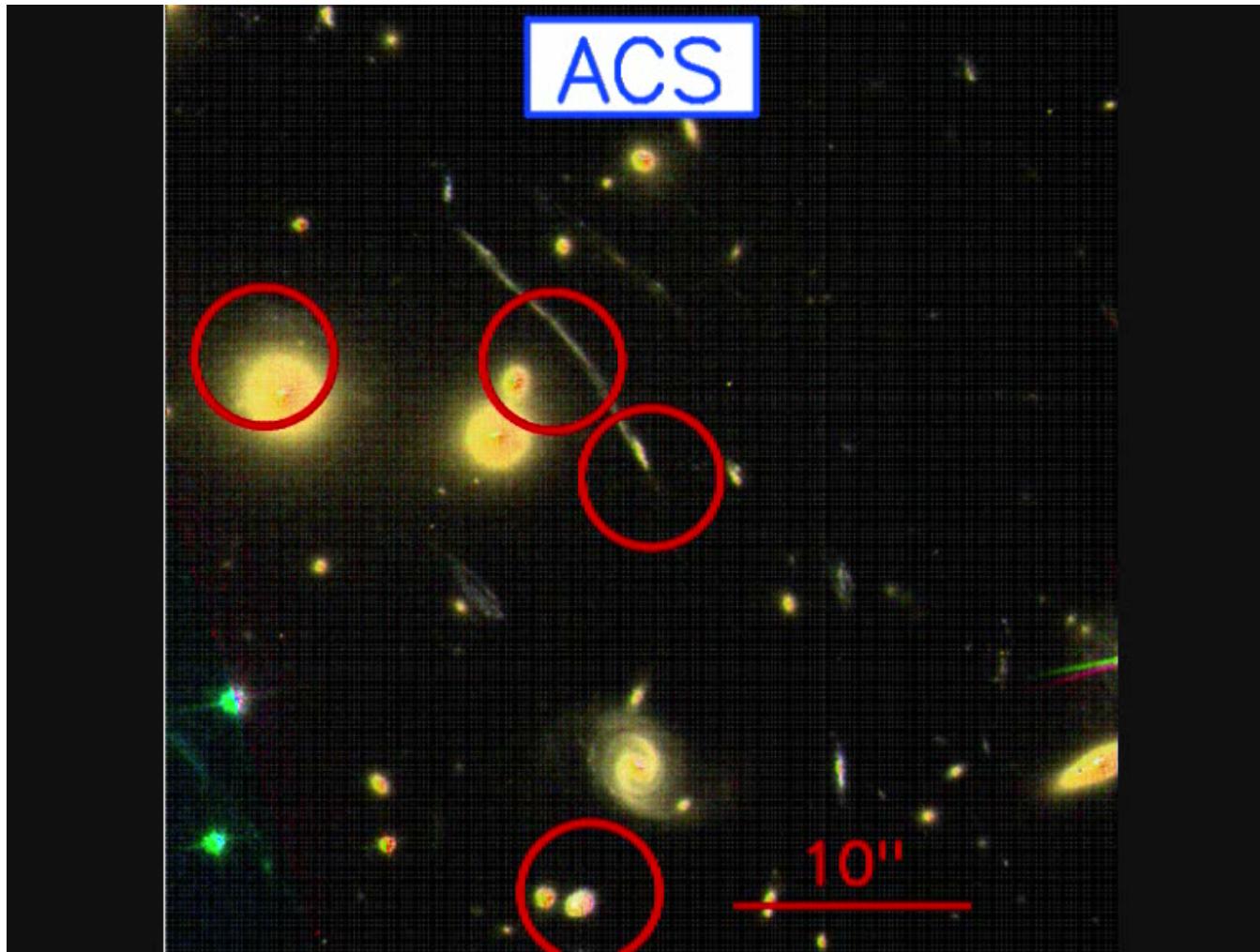
24 μm -derived SFRs/ L_{TIR} tend to overestimate the true values.



Suggest that
there's still
a lot to learn
about the
properties of
Far-IR/Submm
galaxy SEDs.

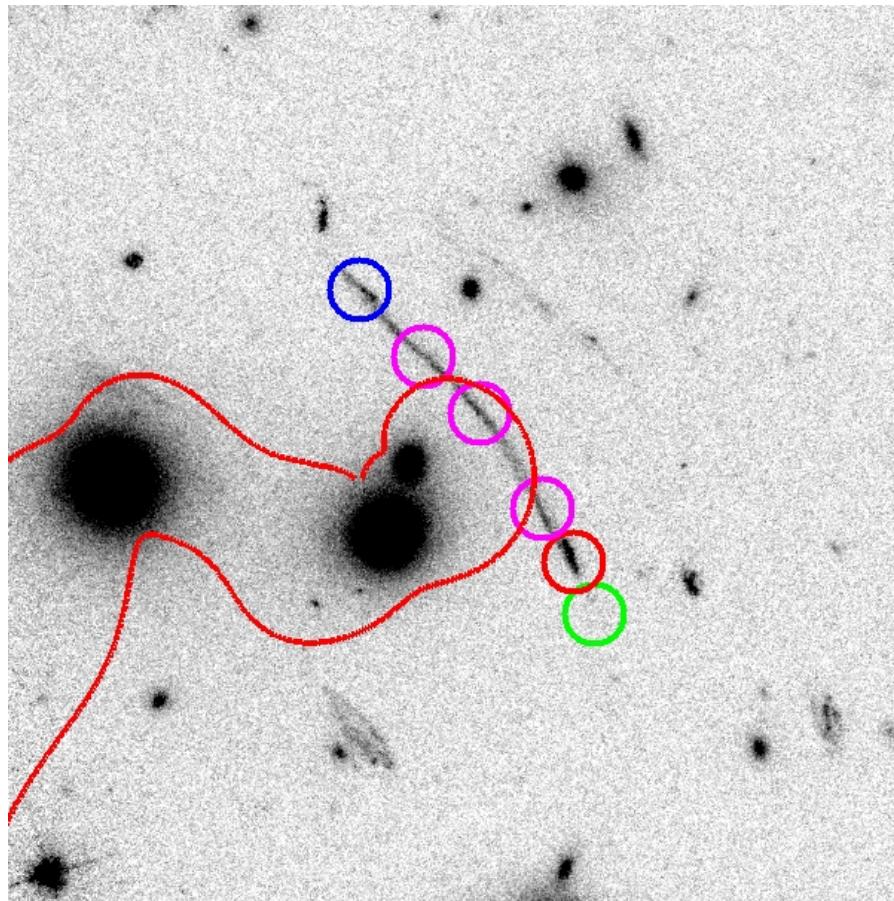
2. Multi-Wavelength Source Matching, Photometry & Far-IR/Submm Phot-z's

Perez-Gonzalez et al. (2010)

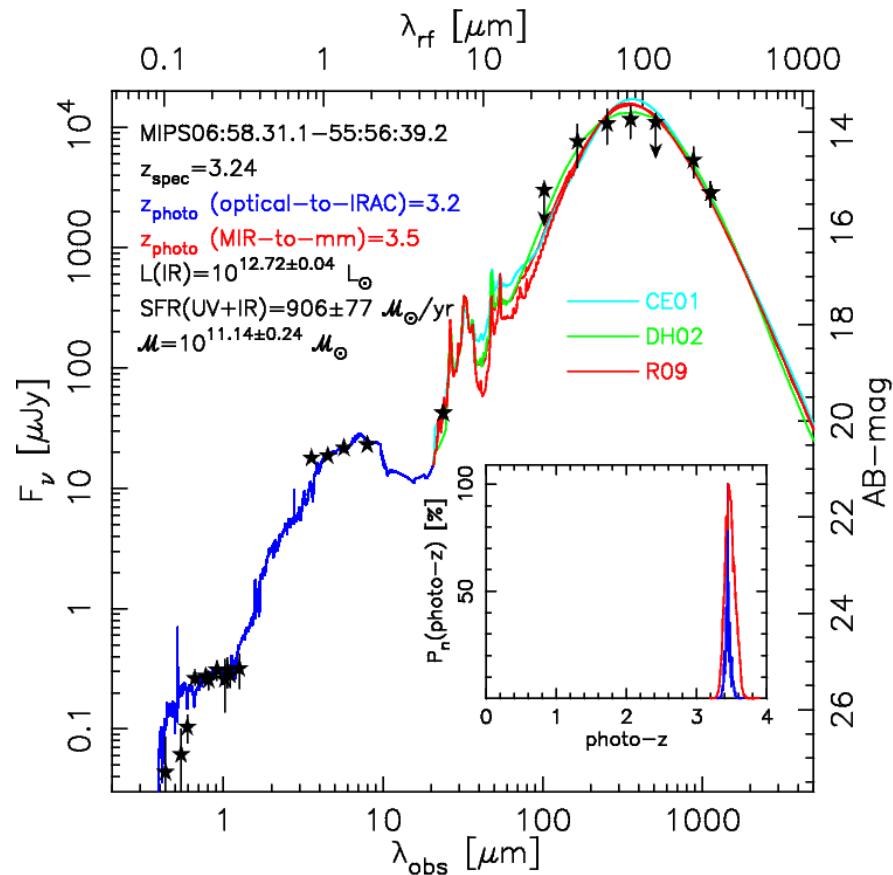
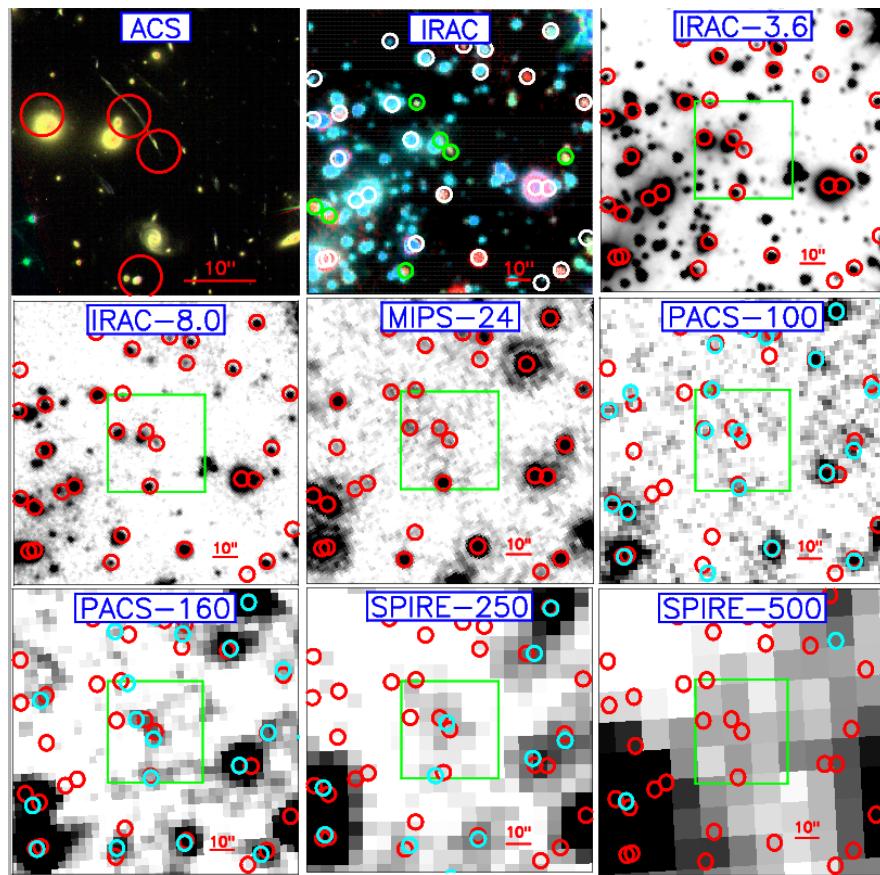


$z=3.24$ giant
lensed
arc

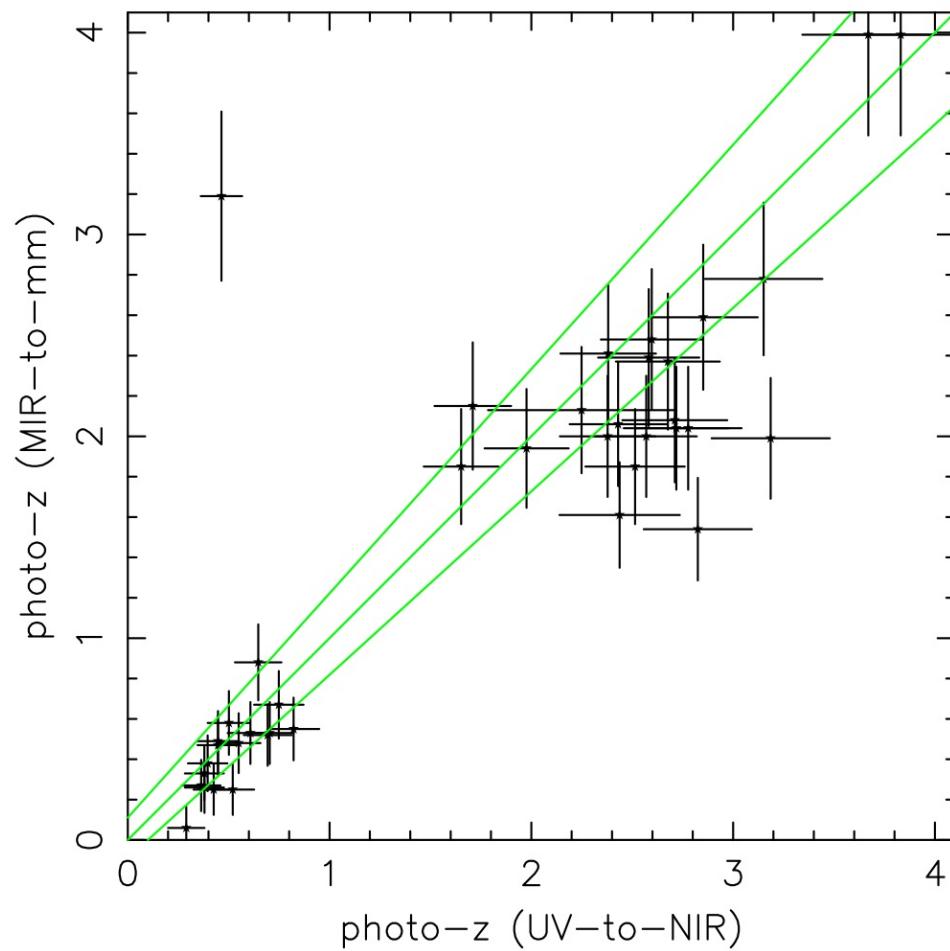
HST Image Reconstruction in the Source Plane at $z=3.24$



Full SED of the z=3.24 Lensed Galaxy



IR/Submm Phot-z's



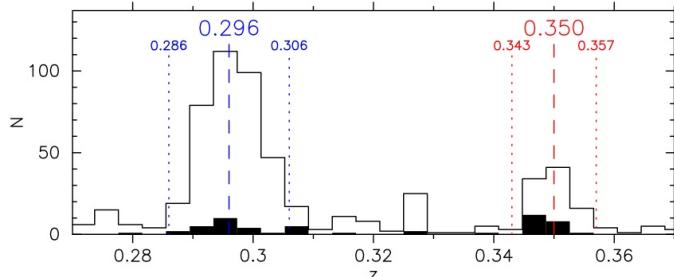
Preliminary
results

Perez-Gonzalez, in prep

3. IR/Submm Properties of Galaxies in Dense Cluster Environment

Rawle et al. 2010; See his poster!

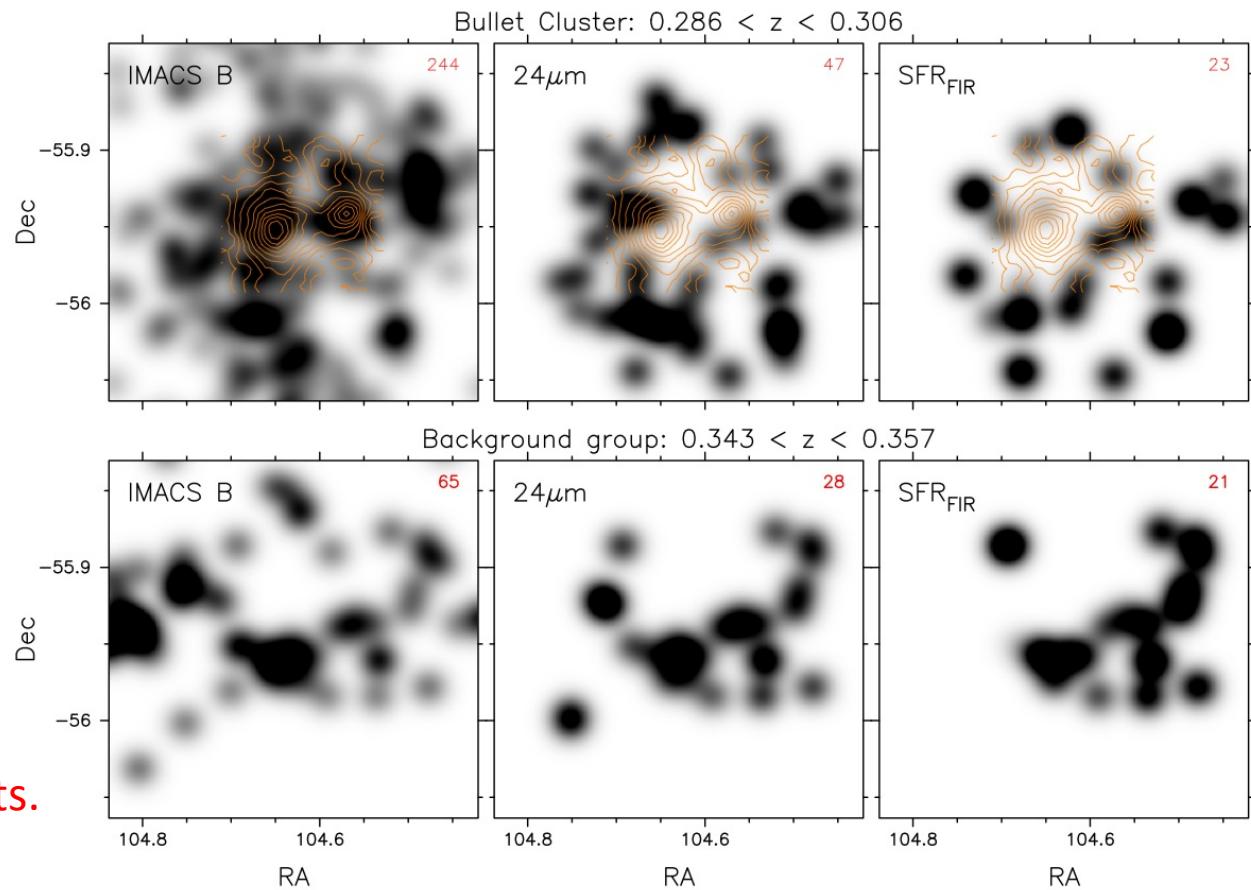
Comparing the Bullet Cluster ($z=0.3$) and the $z=0.35$ background system



$144 M_{\odot} \text{ yr}^{-1}$ $207 M_{\odot} \text{ yr}^{-1}$

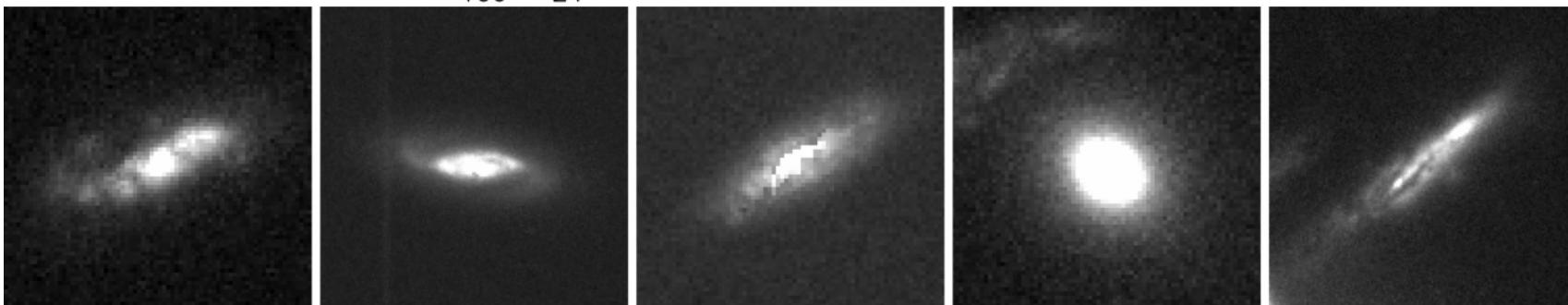
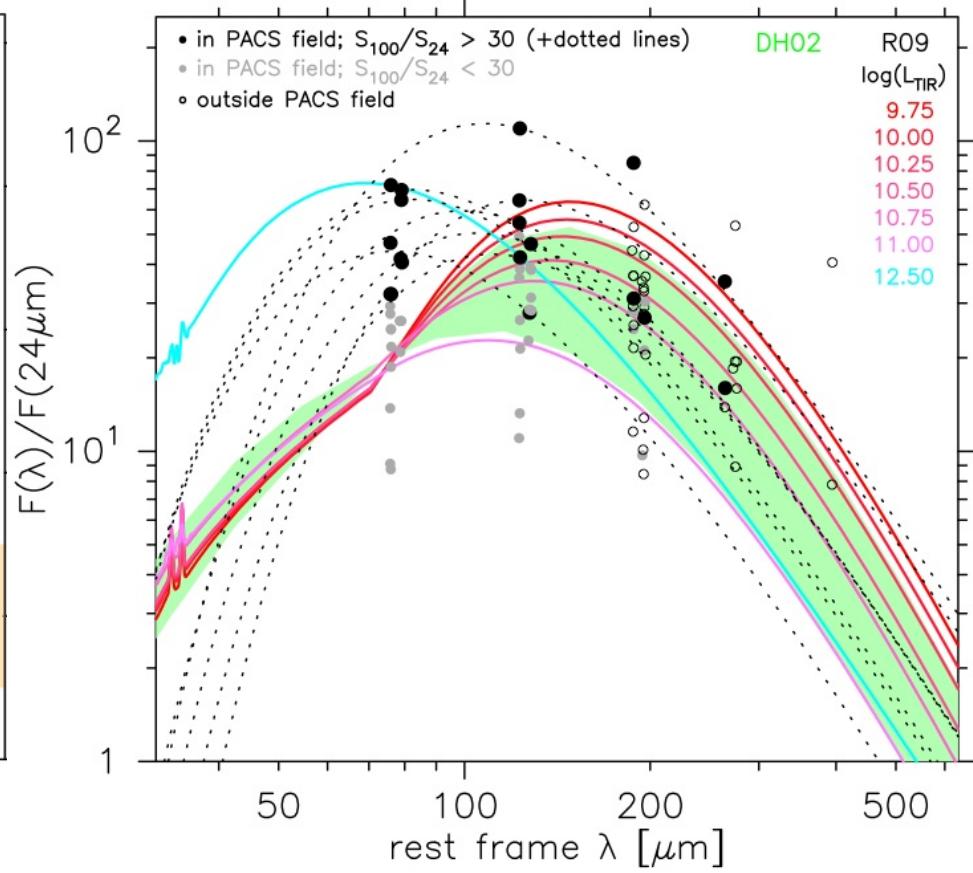
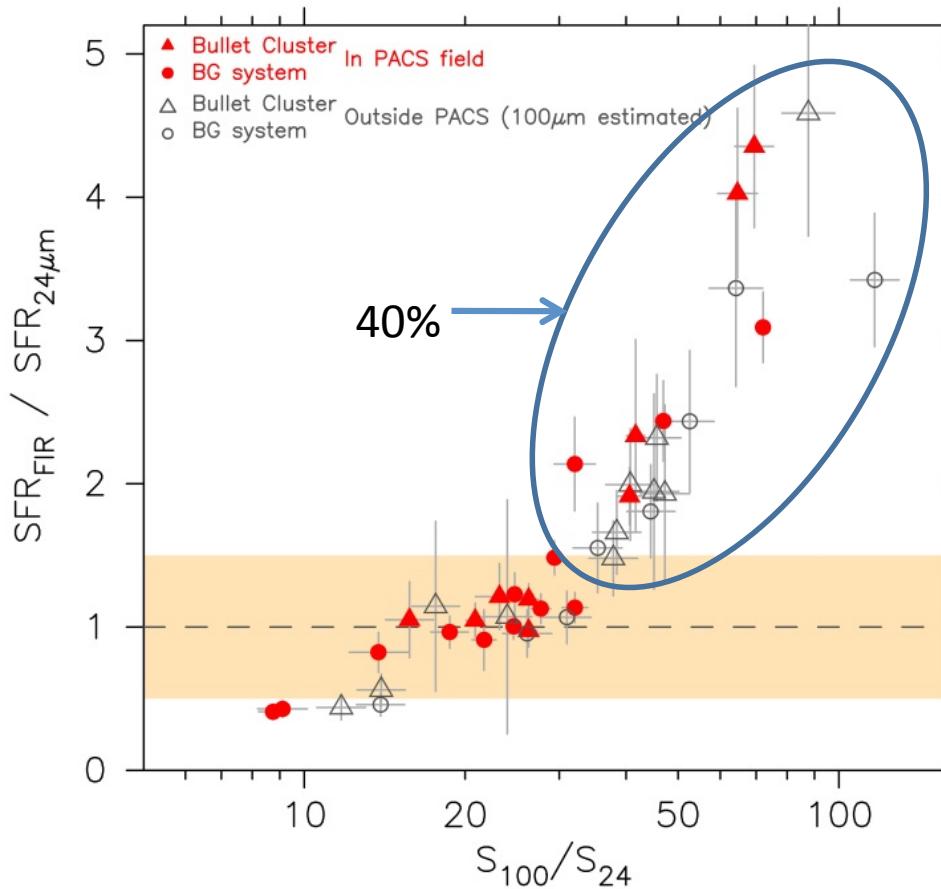
IR-derived total SFRs

Cluster-cluster mergers
may not be important for
triggering IR-luminous starbursts.



Discovery of 100 μm -Excess Galaxies

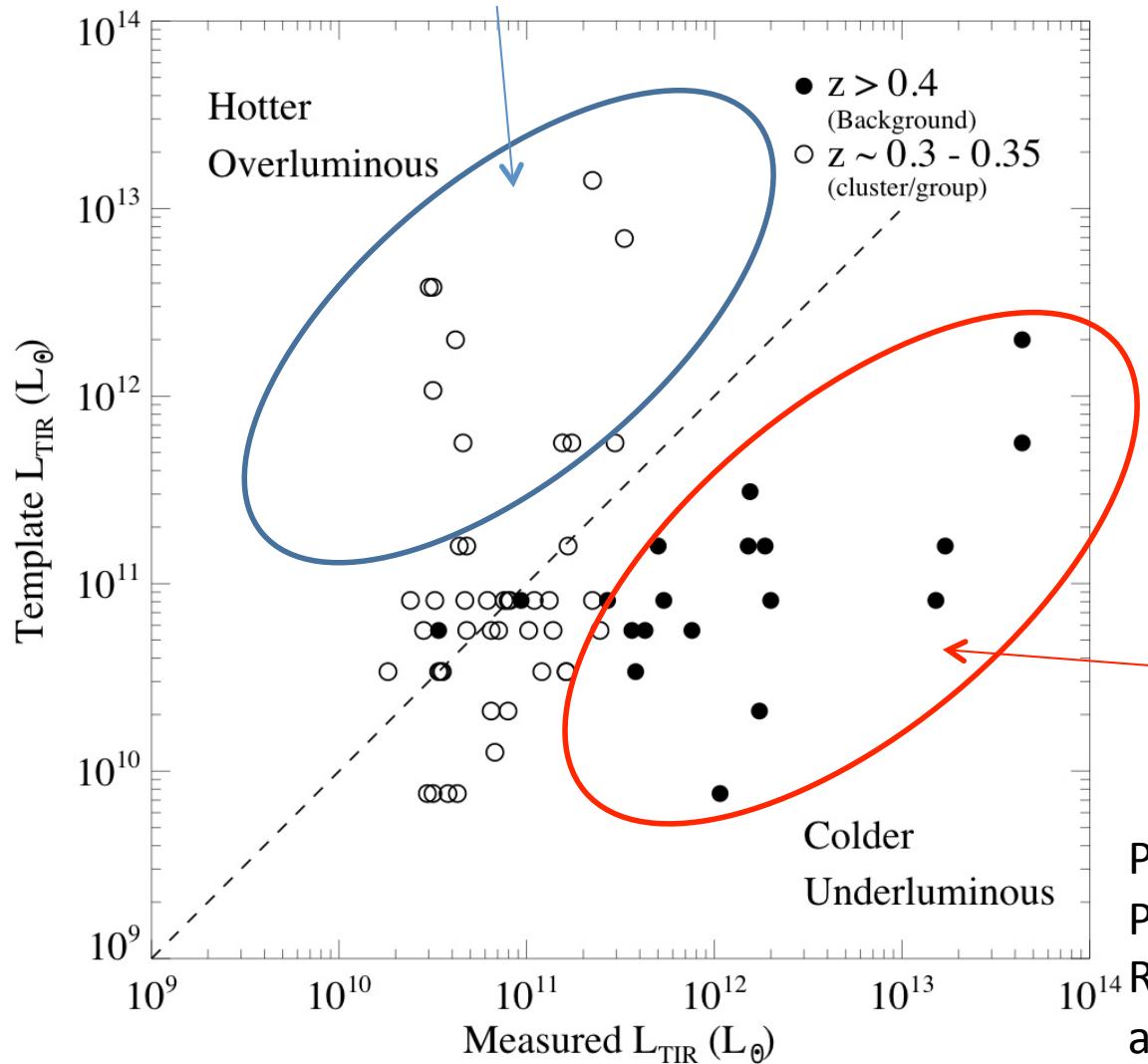
Does the dense cluster environment modify the far-IR/Submm SED properties?



Diverse morphology

IR/Submm SED Properties: Summary

Surprise!!



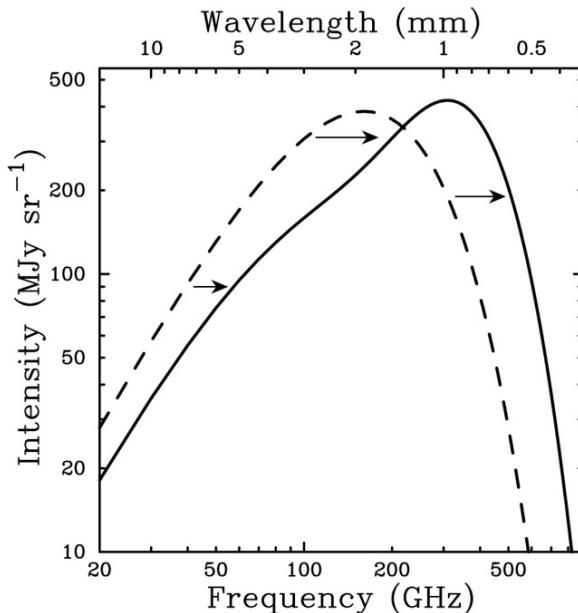
Consistent
with
Spitzer
results.

Pope et al. (2006)
Papovich et al. (2007)
Rigby et al. (2008)
and etc...

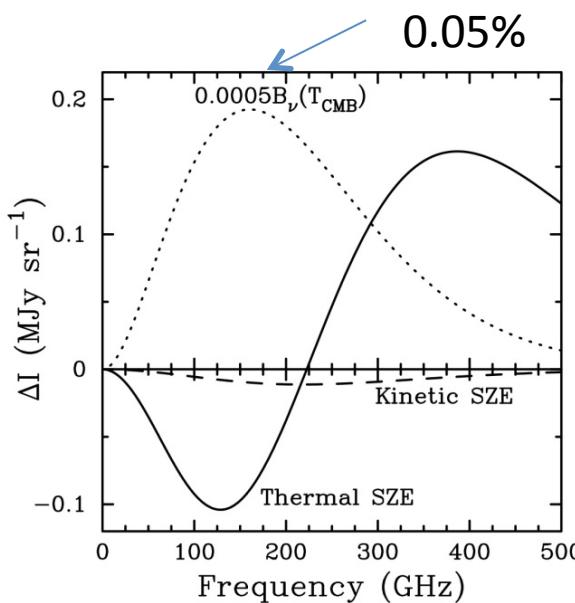
4. First Detection of the Sunyaev-Zel'dovich Effect Increment at < 650 um

Zemcov et al. (2010)

First, some introduction....

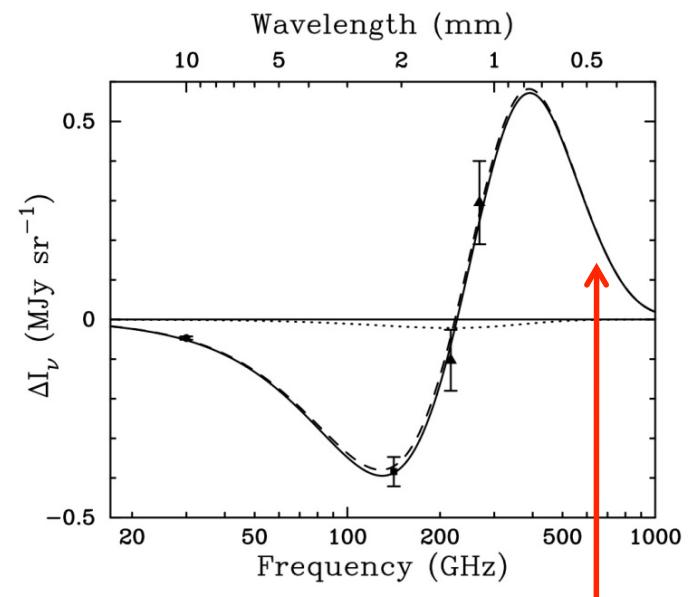


CMB photons gain energy when scattered by high energy electrons in ICM. (The effect exaggerated in the figure above.)

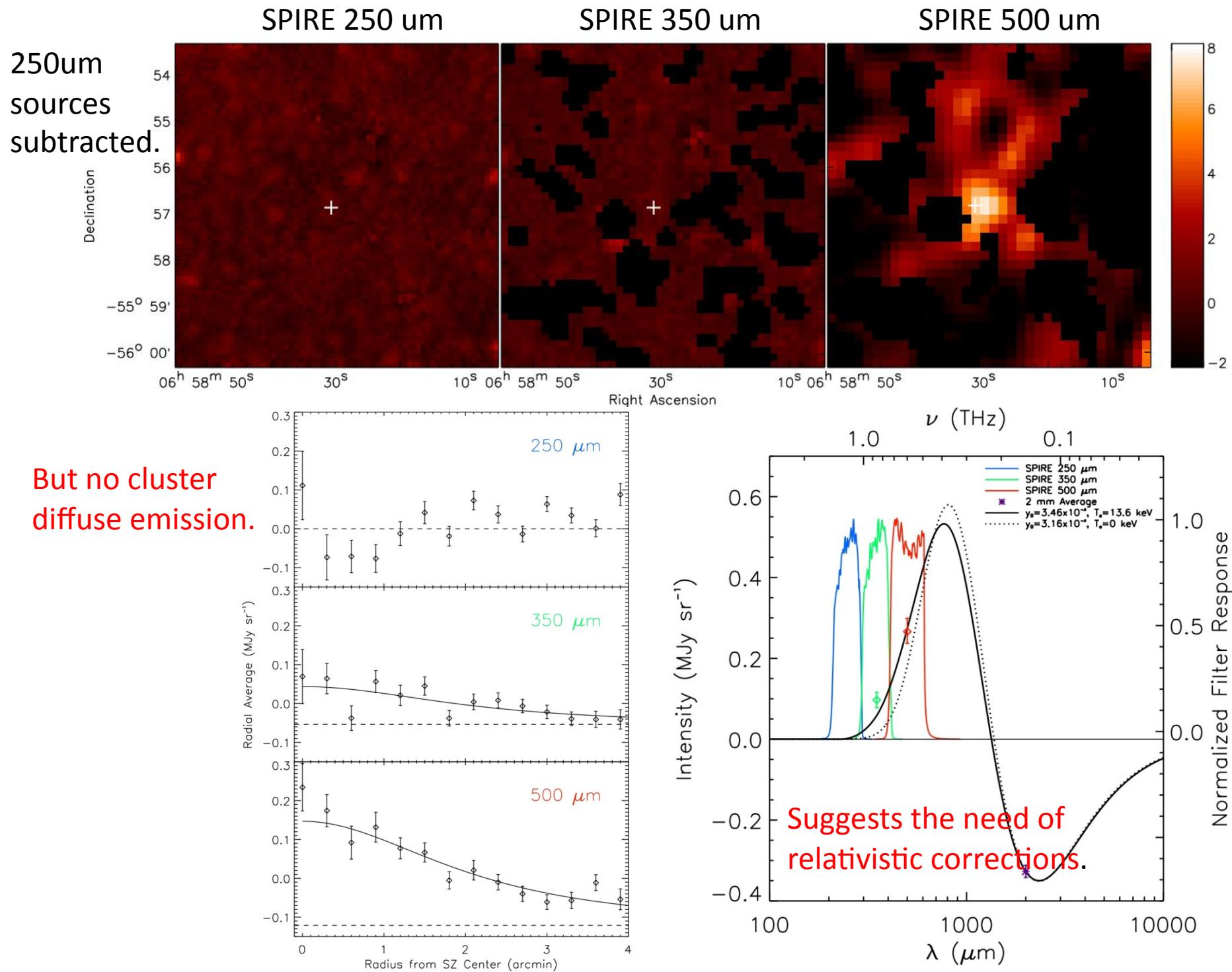


Resultant spectral distortion of the CMB due to the SZ effect.

(from Carlstrom et al. 2002)



SPIRE will see this part of the SZ signal.



Summary

- The Herschel Lensing Survey (HLS) is delivering what it promised (lensed high-z galaxies)...and more (cluster members, SZ effect)!
- SEDs of high-redshift IR-bright galaxies → colder ($F_{24\mu m}$ overpredicts L_{TIR} with local SEDs)
- SEDs of 100 μm -excess cluster-member galaxies → hotter ($F_{24\mu m}$ underpredicts L_{TIR} with local SEDs)
- SZ-effect increment clearly detected at 350/500 μm
 - Fits the expected SZ spectrum, but suggests relativistic corrections due to fast-moving electrons.
- Road ahead:
 - 39 more clusters to analyze in detail!
 - LABOCA, SCUBA2, LMT and various optical/near-IR observations proposed/planned for the near future.