

Environmental effects on the dust properties of Virgo cluster galaxies

Luca Cortese (Cardiff)

On behalf of SAG2 & HeViCS Teams

Thanks to

- Goran Pilbratt, ESA, the SPIRE and PACS consortia!

- SAG2 - SPIRE Local Galaxies Group

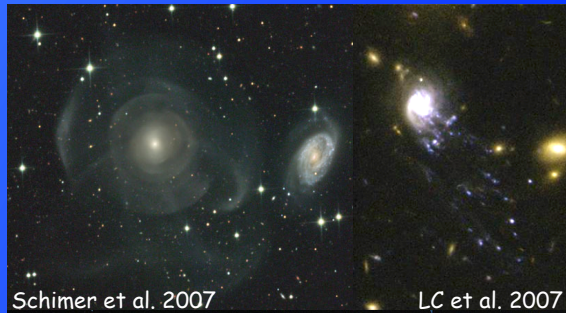
Auld R., Baes M., Barlow M., Bendo G., Bock J., Boselli A., Bradford M., Buat V., Castro-Rodriguez N., Chaniel P., Charlot S., Clements D., Cooray A., Cormier D., Cortese L., Davies J., Dwek E., Eales S., Elbaz D., Galametz M., Galliano F., Gear W., Glenn J., Gomez H., Griffin M., Hony S., Isaak K., Leboutteiller V., Levenson L., Lu N., Madden S., O'Halloran B., Okumura K., Oliver S., Page M., Panuzzo P., Papageorgiou A., Parkin T., Perez Fournon I., Pohlen M., Rangwala N., Rigby E., Roussel H., Rykala A., Sacchi N., Sauvage M., Schulz B., Schirm M., Smith M., Spinoglio L., Srinivasan S., Stevens J., Trichas M., Symeonidis M., Vaccari M., Vigroux L., Wilson C., Wozniak H., Wright G., Zeilinger W.

- The Herschel Virgo Cluster Survey (HeViCS) Team

Davies J. (P.I.), M. Baes, G. J. Bendo, S. Bianchi, H. Boehringer, D. Bomans, A. Boselli, P. Chaniel, M. Clemens, E. Corbelli, L. Cortese, A. Dariush, I. De Looze, S. Dye, S. Eales, D. Fadda, J. Fritz, D. Garcia-Appadoo, G. Gavazzi, C. Giovanardi, M. Grossi, T. Hughes, L. Hunt, A. Jones, S. Madden, D. Pierini, M. Pohlen, M. Putman, S. Sabatini, M. Smith, S. di Serego Alighieri, J. Verstappen, C. Vlahakis, E. Xilouris, S. Zibetti

Environmental effects on galaxy evolution

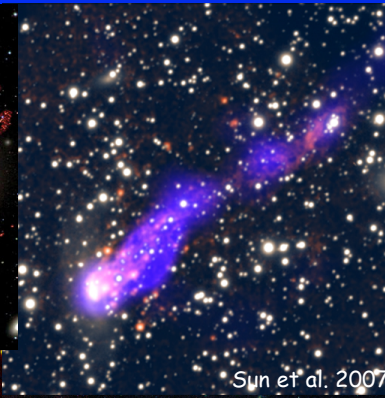
So far a plethora of info about stars and gas (atomic, molecular, ionized)...



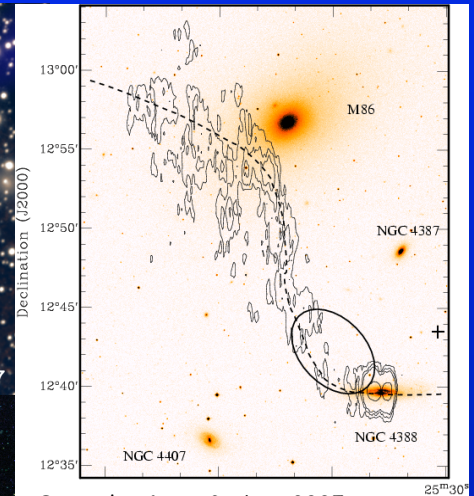
LC et al. 2007



Kenney et al. 2008



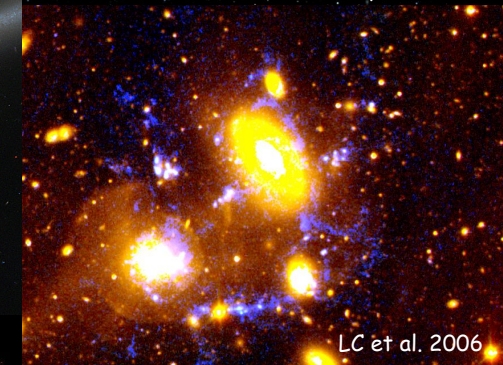
Sun et al. 2007



Oosterloo & van Gorkom 2005



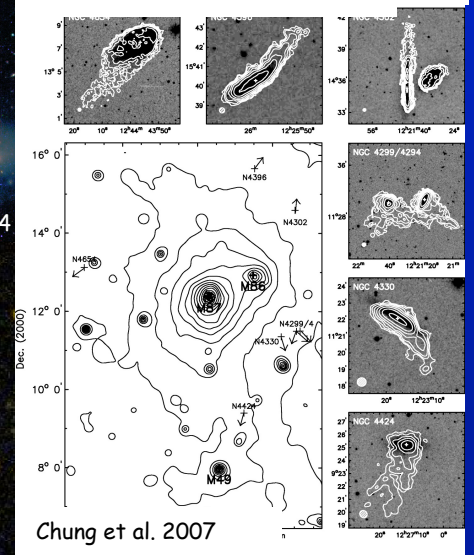
LC et al. 2006



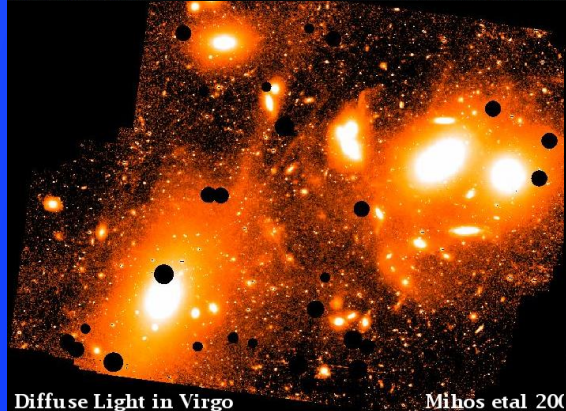
LC et al. 2006



Yoshida et al. 2004



Chung et al. 2007



Mihos et al. 2006



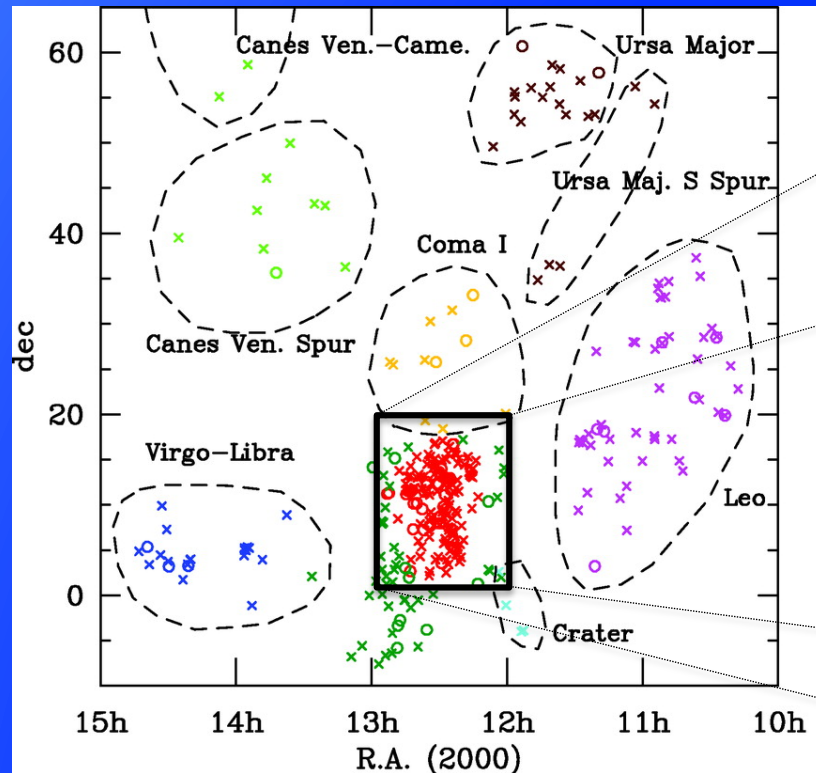
Duc et al. 2007

What about dust?

HRS

Herschel Reference Survey

Boselli et al. 2010, PASP, 122, 261



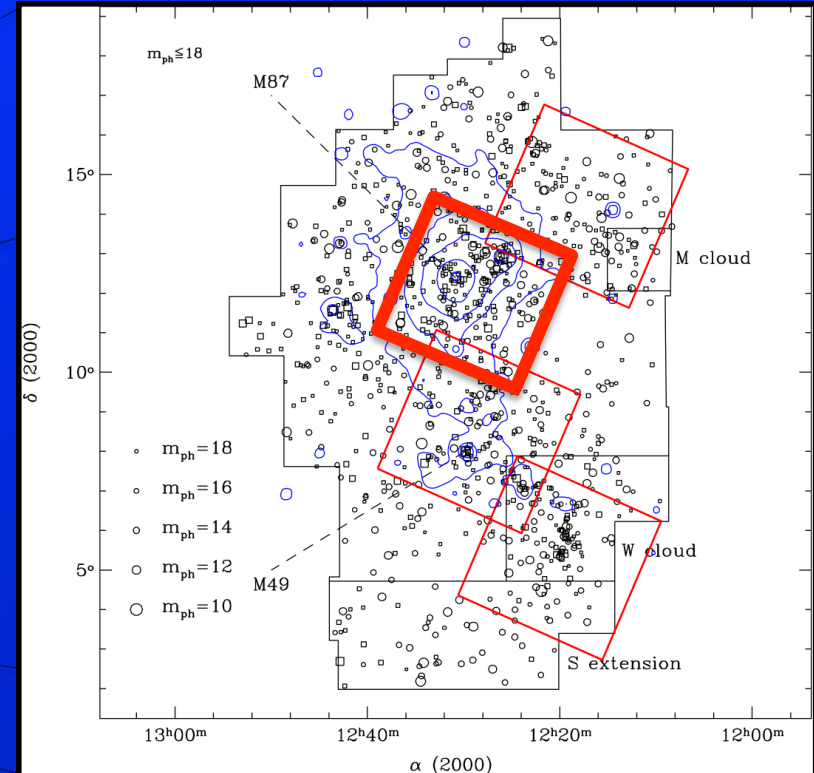
- SPIRE 250, 350, 500 μ m maps
- $15 < D < 25$ Mpc
- $K < 12$ for Sp. -- $K < 8.7$ for E/S0
- Gal. lat. $> +55$ deg & $A(B) < 0.2$ mag
- 323 obj. (65 E/S0, 258 Sp./Irr)

See talk by Eales (tomorrow)

HeViCS

Herschel Virgo Cluster Survey

Davies et al. 2010, A&A in press



- SPIRE+PACS maps
- 100, 160, 250, 350, 500 μ m
- ~ 64 deg 2

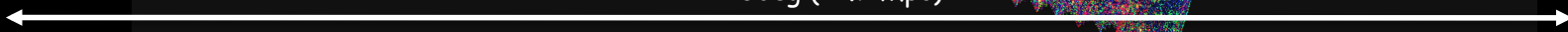
See talks by
Bianchi, Baes, De Looze

NGC4569

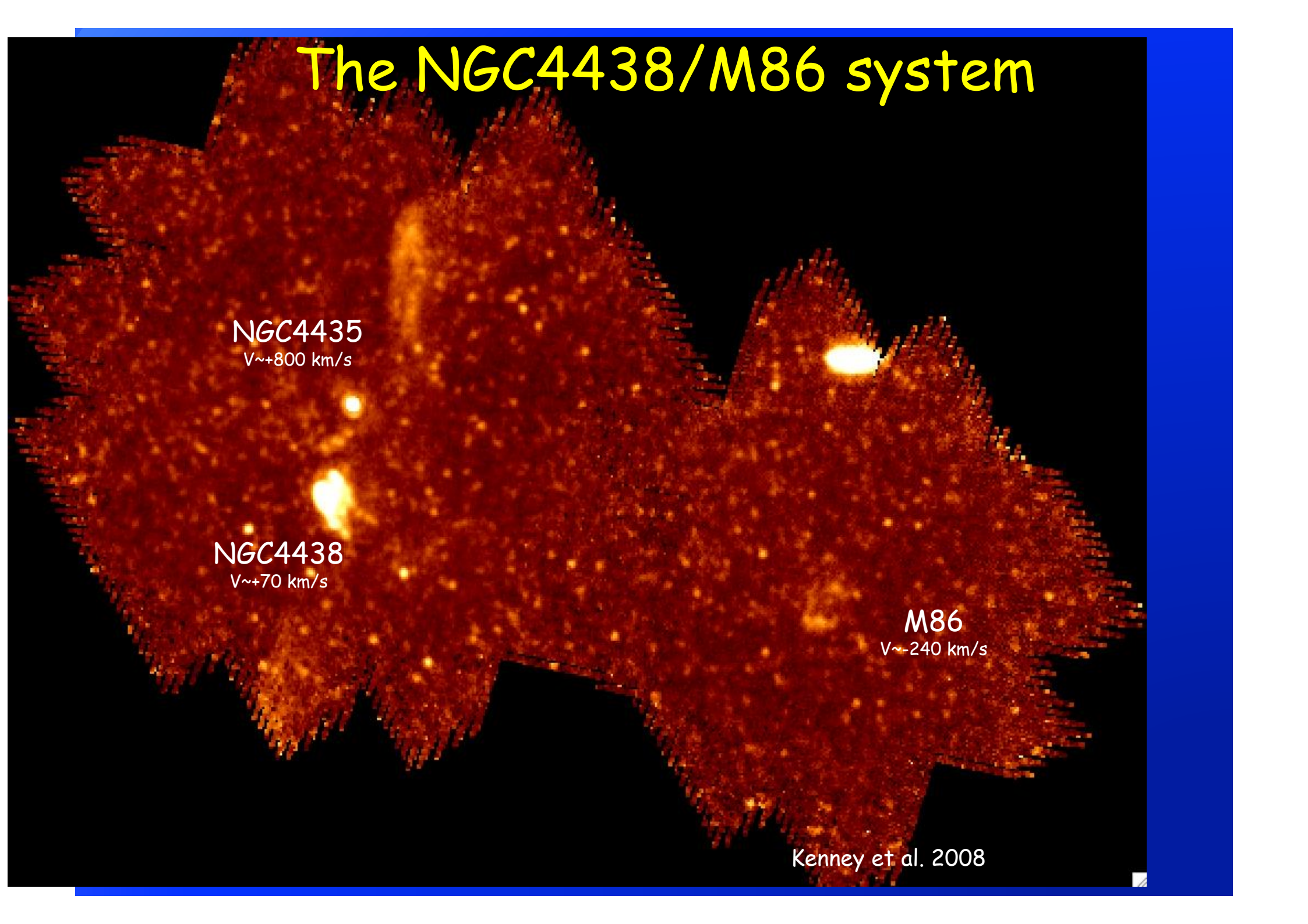
M86 group

M87

~6deg (~1.7 Mpc)



The NGC4438/M86 system



NGC4435

$V \sim +800$ km/s

NGC4438

$V \sim +70$ km/s

M86

$V \sim -240$ km/s

Kenney et al. 2008

The NGC4438/M86 system

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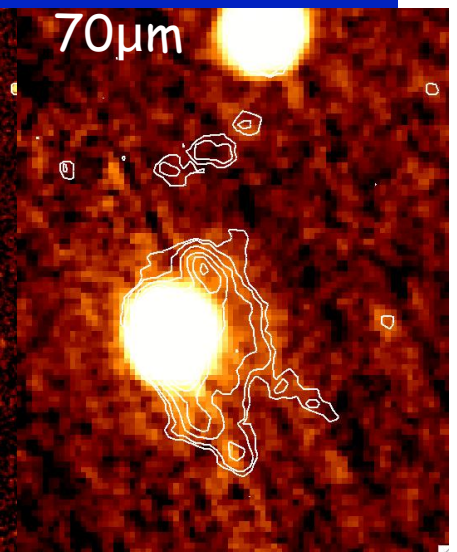
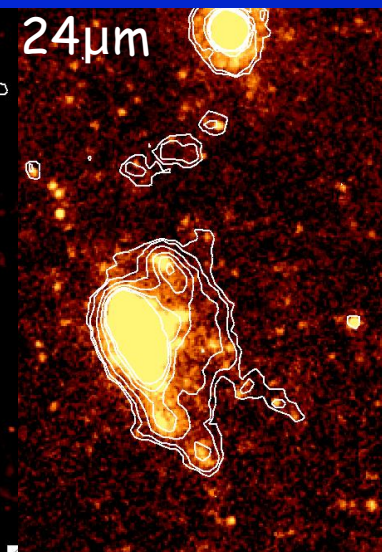
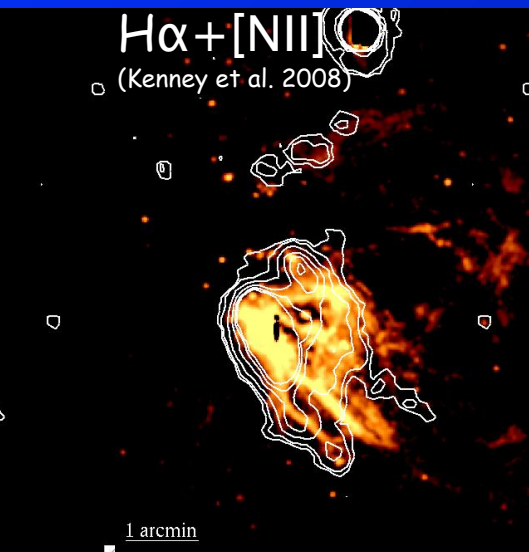
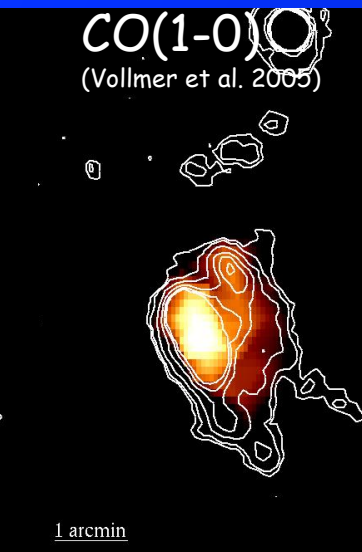
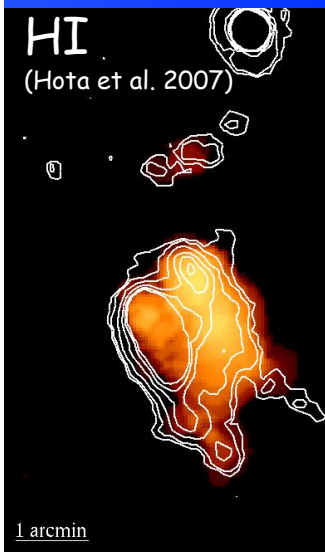
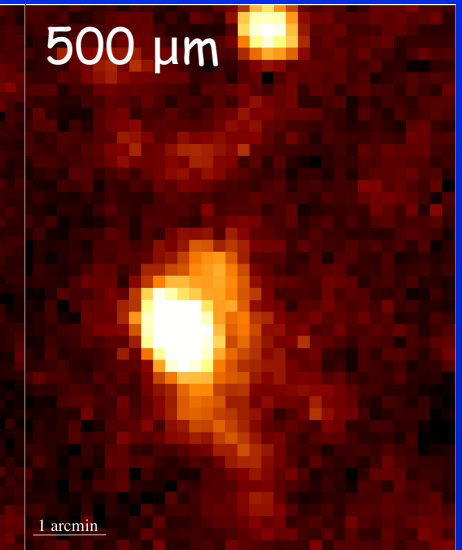
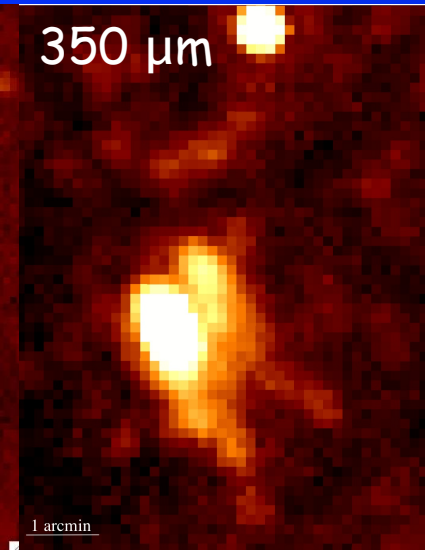
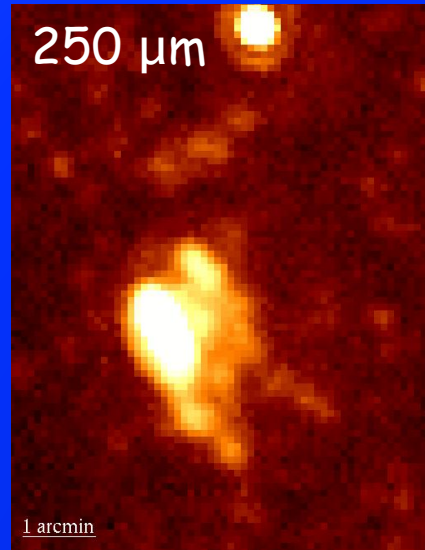
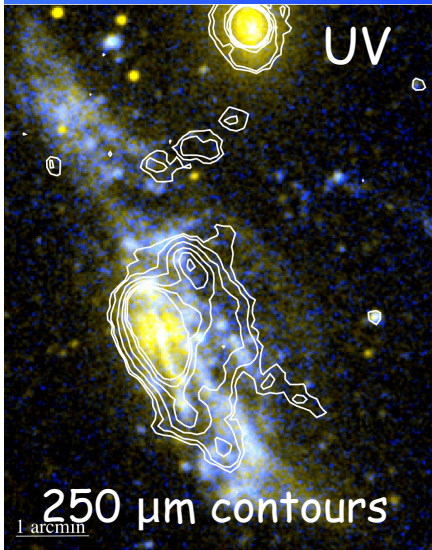
$V \sim +70$ km/s

M86

$V \sim -240$ km/s

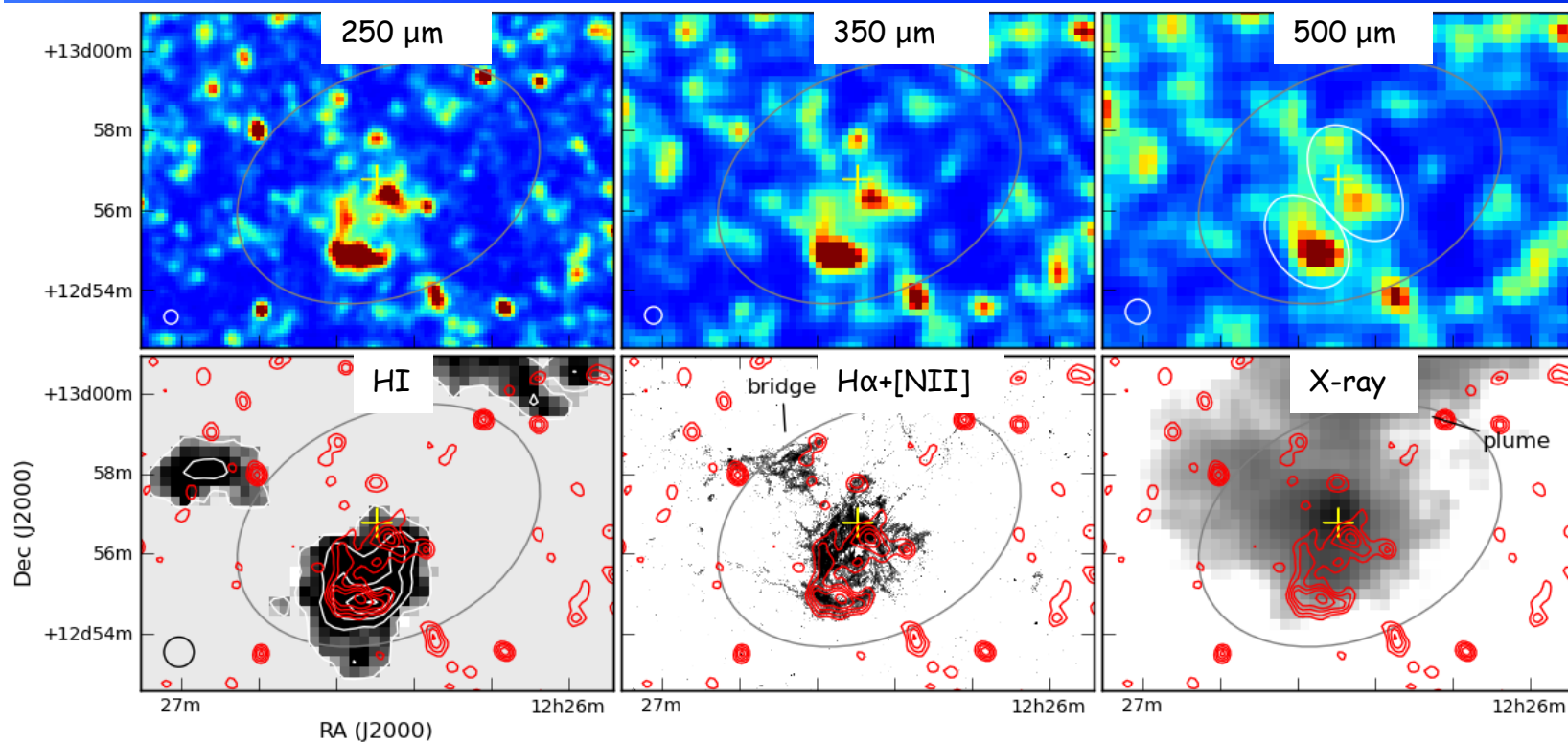
Kenney et al. 2008

NGC4438



Extra-planar cold dust associated with stripped gas

M86

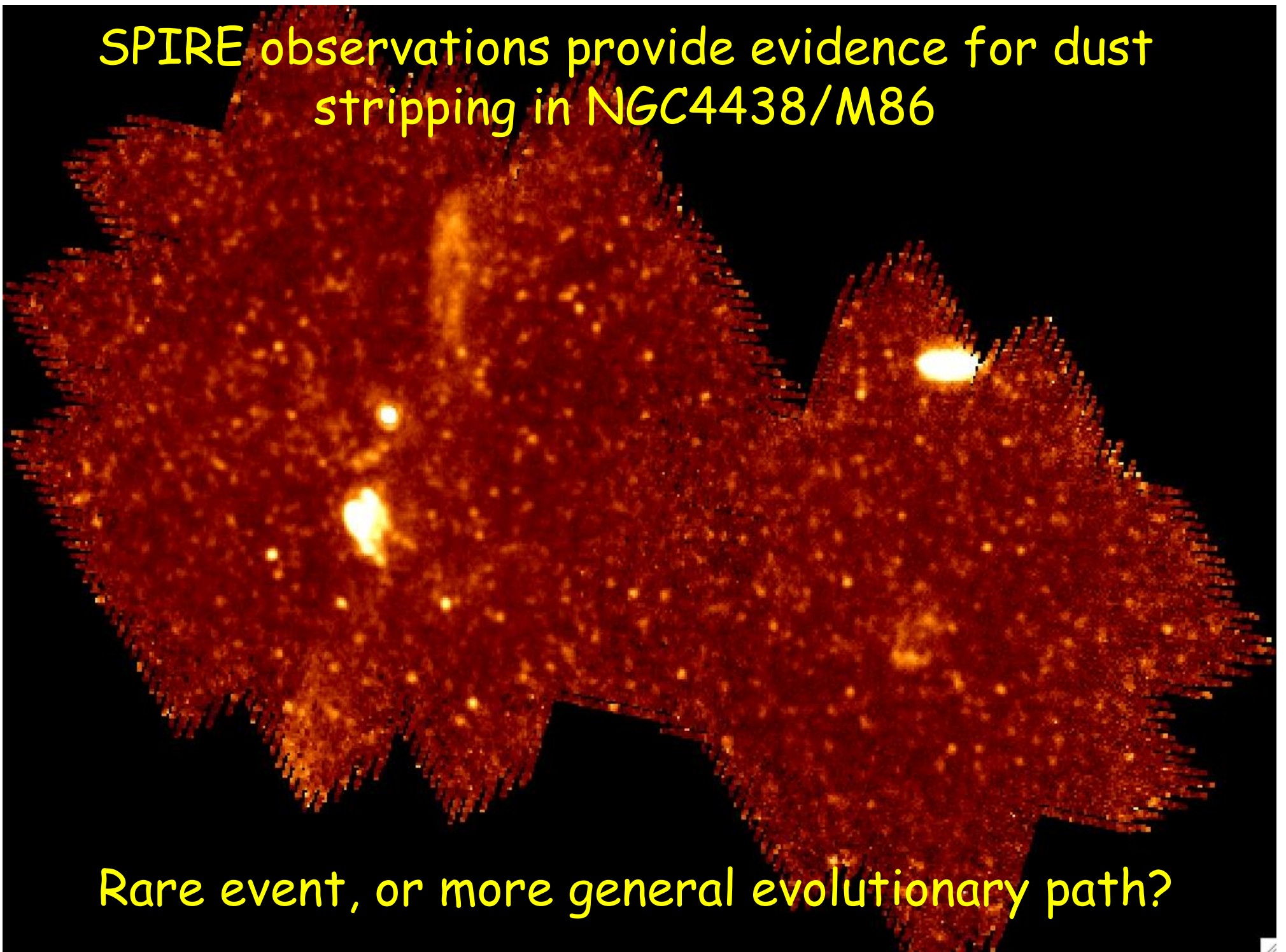


Cold dust follows closely ionized ($H\alpha+[NII]$) gas
Likely material stripped from NGC4438

Heating mechanism not clear, likely the same responsible for $H\alpha+[NII]$ emission: e.g., tidal heating

See poster by Haley Gomez (P1.39) for more details!

SPIRE observations provide evidence for dust stripping in NGC4438/M86



Rare event, or more general evolutionary path?

Ideally, we want to estimate a
"dust-deficiency" parameter...

HI deficiency

$$\text{def}(\text{HI}) = \log(\langle M(\text{HI}, D_{\text{opt}}, \text{Type}) \rangle) - \log(M(\text{HI})_{\text{obs}})$$

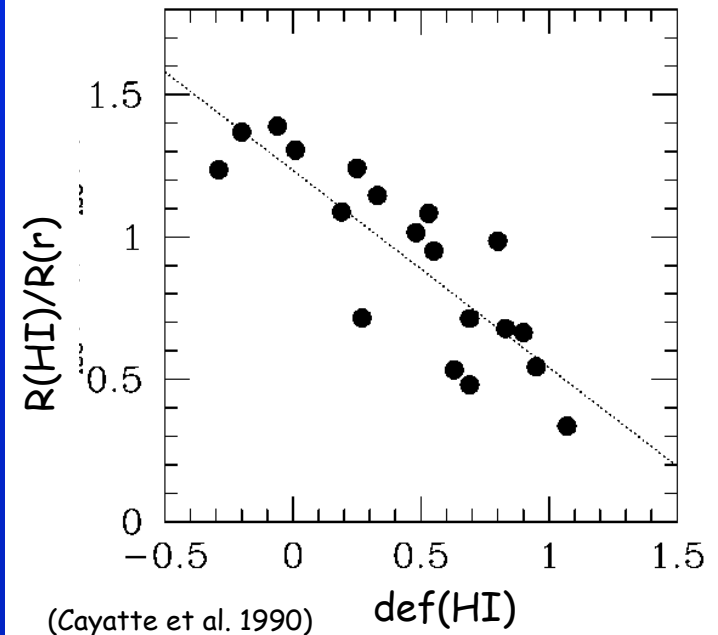
(Haynes & Giovanelli 1984)

$\text{def}(\text{HI}) = 0 \implies$ Normal HI content

$\text{def}(\text{HI}) = 1 \implies$ Lost 90% of HI

HI deficiency associated with "truncation" in the HI disk

Dust is stripped \implies
"truncation" in the dust disk ?

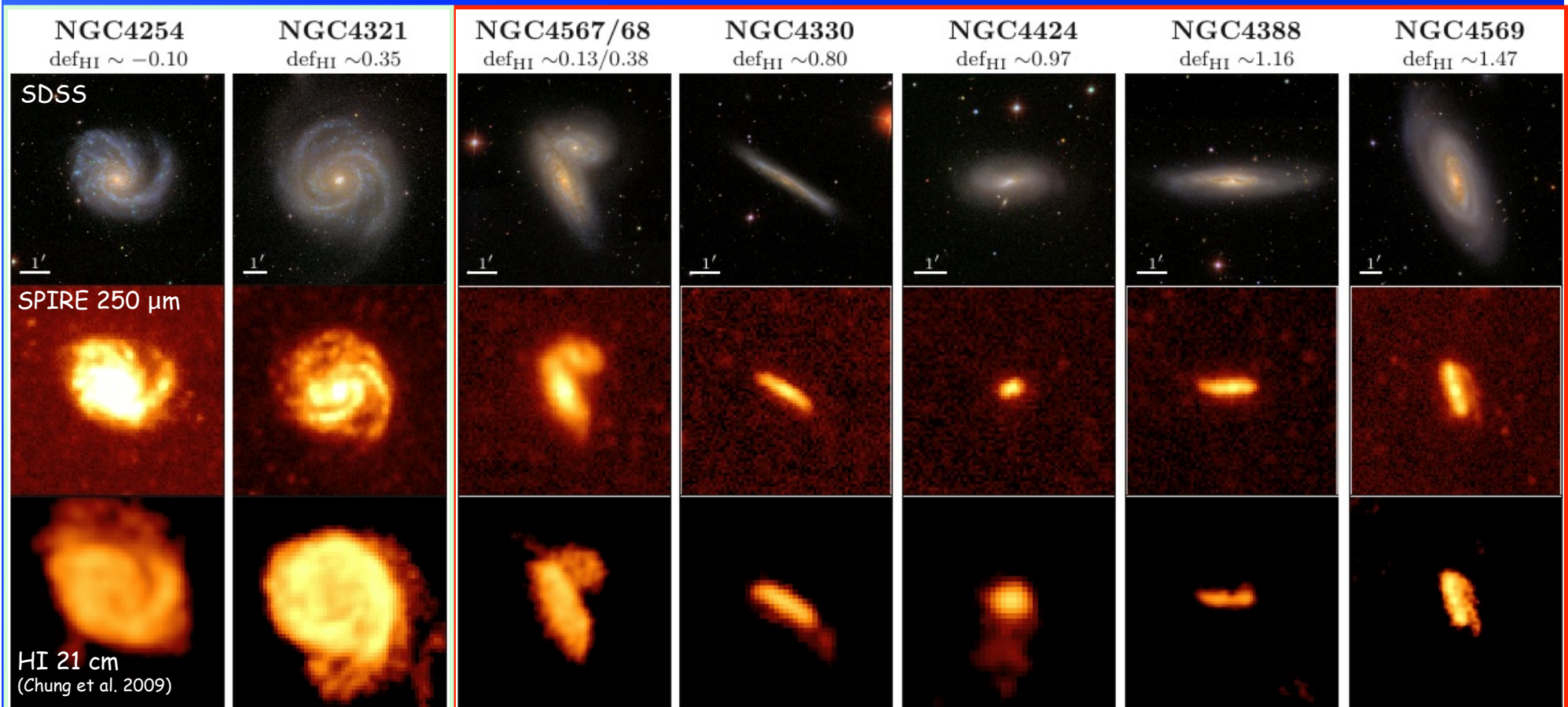


"Truncated" dust disk in HI-deficient Virgo spirals

HI-rich



HI-poor



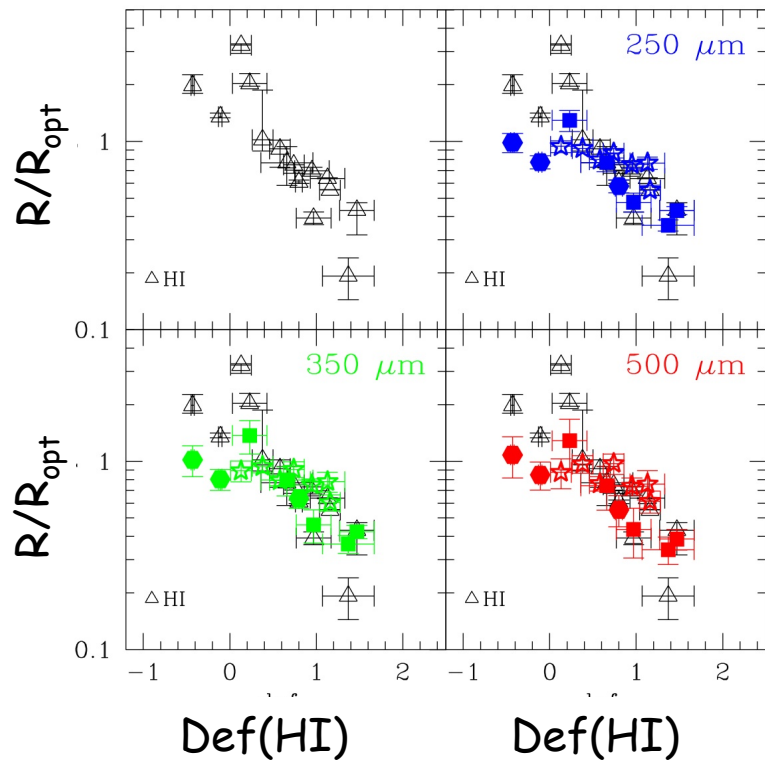
Pohlen et al. 2010 A&A, in press

LC et al. 2010b A&A, subm.

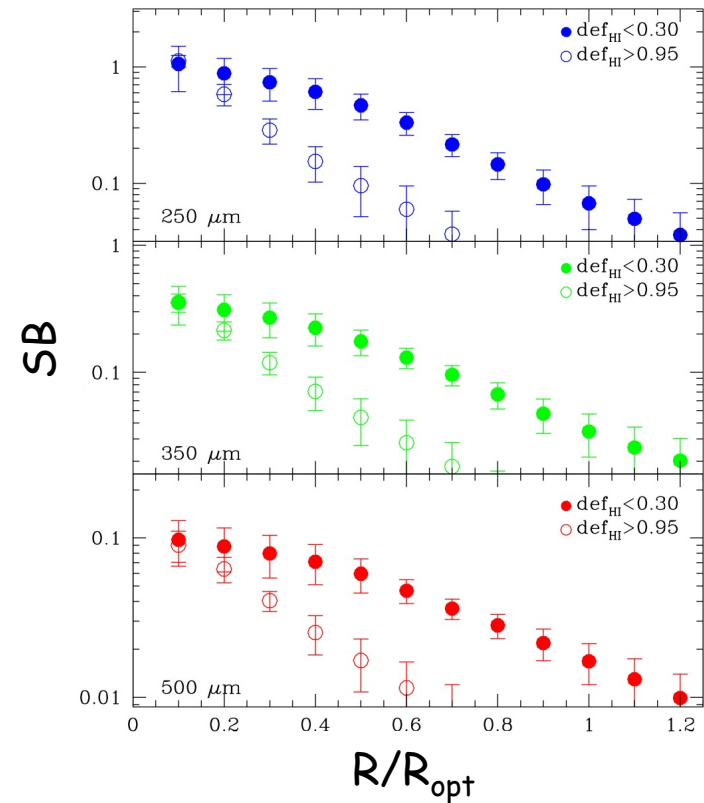
See Poster by Matt Smith (P2.44)

"Truncated" dust disk in HI-deficient Virgo spirals

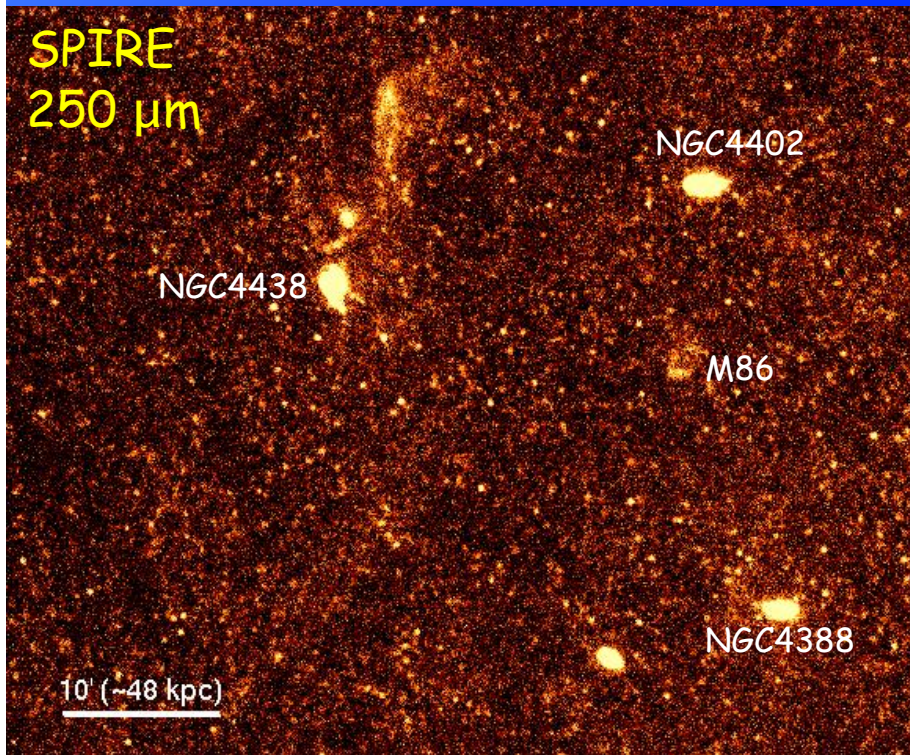
Isophotal radii



Average SB profiles



Looking for stripped dust - Beware of cirrus!

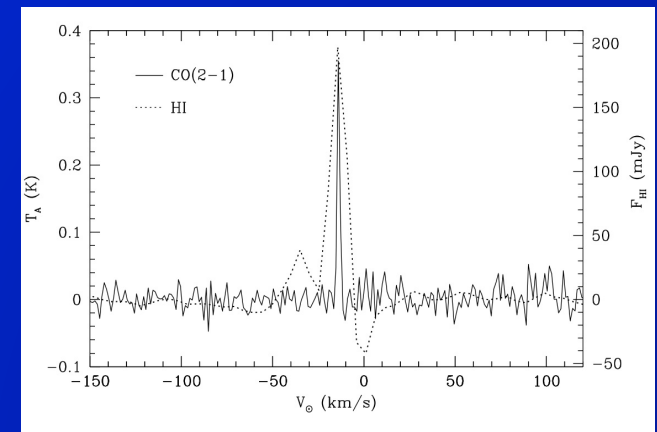
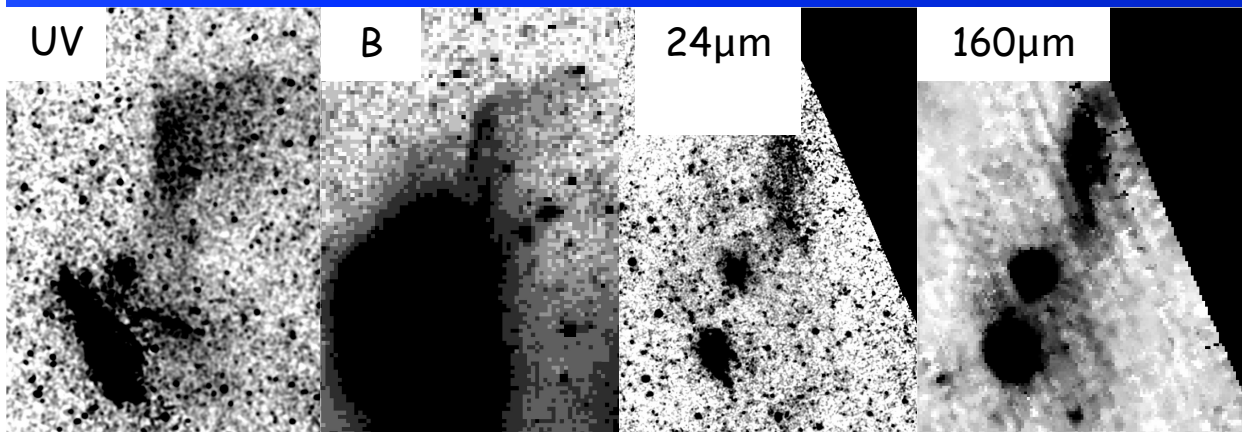


Diffuse Plume near NGC4438

FUV/opt/24 μm /160 μm /HI/CO counterpart

Previously interpreted as a tidal stream

Properties and kinematics more consistent with cirrus than tidal debris



Summary

- HeViCS+HRS observations of the Virgo cluster: first glimpse of the effects of the environment on the dust properties of cluster spirals.

Very promising data to:

- Look for “clear” examples of dust stripping
- Quantify the amount of dust-deficiency in cluster spirals
- Look for any diffuse intra-cluster dust stripped from infalling spirals

