



Unbiased HIFI spectral survey of SgrB2(M)

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00^s

Dowell et al. 1999

 $L = 5 \times 10^6 L_{\odot}$

 $M = 5 \times 10^5 M_{\odot}$

17^h44^m20^s

Centimeter continuum



Vol. 474



Observing SgrB2(M) allows a simultaneous study of

- a very massive, cluster-forming core
 - Including the velocity field
 - Density and temperature structure
 - Chemical structure
- a large collection of translucent and diffuse clouds across the Galaxy

- By serving as background illumination





1a



1Ь



2a



b



Spiral arm cloud molecules



Schilke et al., 2010

H₂O ortho/para ratio



Lis et al. 2010

Fits by XLASS and MAGIX using data from CDMS and JPL And from Müller, priv. comm.

CH



Qin et al. 2010

Fits by XLASS and MAGIX using data from CDMS and JPL And from Müller, priv. comm.



 H_2CI^+

Unbiased surveys: Expect the unexpected!



Comparing column densities



Physical structure of core: HCN

Green: Infall only Red: Infall outside, expansion at center

HCN HCN 6-5 HCN 7-6 HCN 8-7 HCN 12-11 620.3 GHz 531.7 GHz 708.8 GHz 1062 GHz 2 2 2 NW/M 0 0 H¹³CN 7-6 H¹³CN 6-5 H¹³CN 8-7 H¹³CN 13-12 604.2 GHz 690.5 GHz 517.9 GHz 1121 GHz 0.5 0.5 0.5 0.5 0 0 0.5 2 HC¹⁵N 6-5 HC¹⁵N 8-7 HC¹⁵N 7-6 2 HC¹⁵N 13-12 1.5 516.2 GHz 602.2 GHz 688.2 GHz 1117 GHz -0.2 1 0.5 0 0 0.2 -0.5 20 80 100 20 40 60 80 Velocity (km/s) 100 20 40 60 80 Velocity (km/s) 80 100 20 40 60 80 Velocity (km/s) 40 60 100 Velocity (km/s)

Rolffs et al. 2010

Spherical modeling with RATRAN



6 50 K $5 \ 10^5 \ cm^{-3}$ 4 Velocity [km/s] 2 Expansion 0 Infall -2 2 10⁵ Radius [AU] 4 10⁵ 0

HCN

Chemical and physical structure: few lines - HDO

HDO Ground state line, CSO



Comito et al. 2003





Chemical and physical structure: many lines - SO₂



SO₂ in Band 4: continuum shutting down lines



Schilke et al. 2010

Outlook

- More bands in SgrB2(M)
 - Finalizing the molecular inventory
- SgrB2(N) complete survey
 - Similar weight class, in a different stage of evolution
- High-resolution ground based followups
 needed to interpret results
 - SMA
 - ALMA!