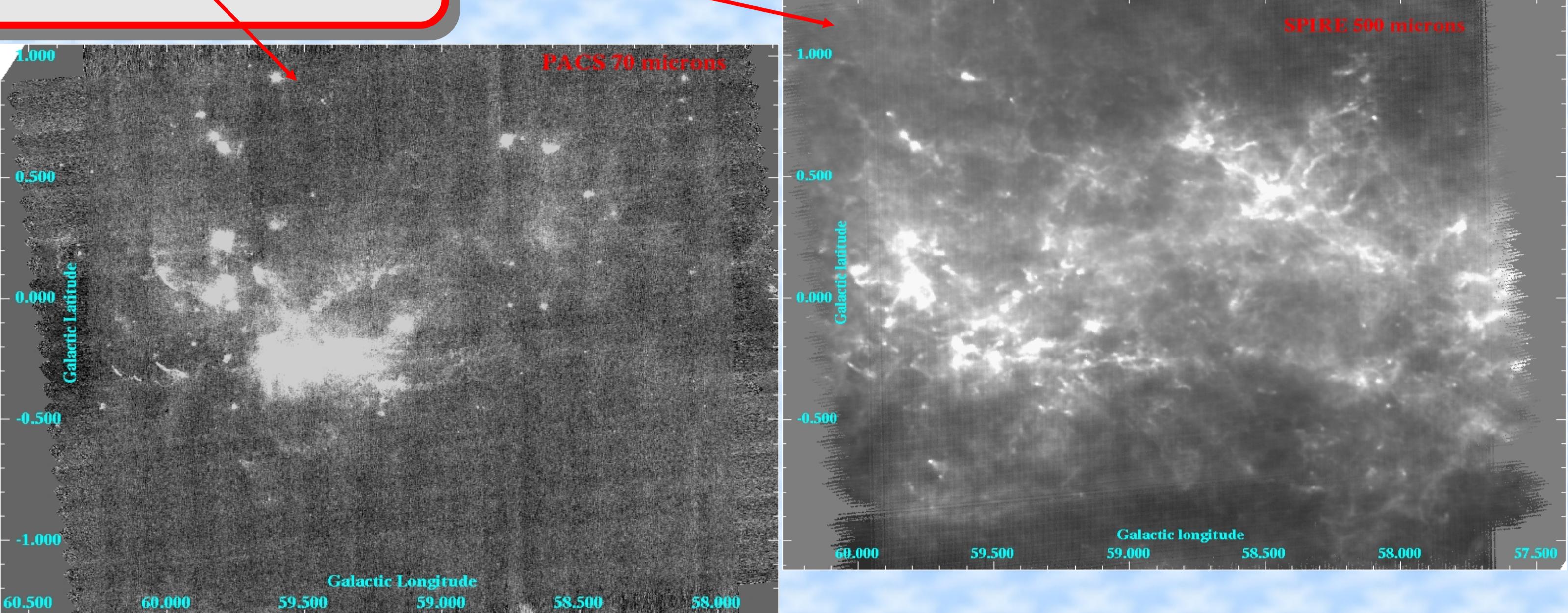


Distance determination for the Hi-GAL Survey

Russeil D., M. Pestalozzi, J. C. Mottram, S. Bontemps, L. D. Anderson, A. Zavagno, M. T. Beltràñ, J. Bally, J. Brand, C. Brunt, R. Cesaroni, G. Joncas, D. Marshall, P. Martin, F. Massi, S. Molinari, T. Moore, A. Noriega-Crespo, L. Olmi, M. A. Thompson, M. Wienen, F. Wyrowski

Context: Hi-GAL is an open time Key-Project of the Herschel satellite that is carrying out a 5-band photometric imaging survey at 70, 160, 250, 350 and 500 μm of the Milky Way Galactic Plane in the longitude range $-60^\circ < l < 60^\circ$ and the latitude range $|b| < 1^\circ$ (Molinari et al. 2010). Our goal is to present the velocity follow-up and the distance determination method for the Hi-GAL compact sources. The field at $l=59^\circ$, observed during the science demonstration phase (SDP), is taken as an example to illustrate our method.

Hi-GAL $2^\circ \times 2^\circ$ images: PACS + SPIRE
case of $l=59^\circ$



Sources catalogue
(Elia et al. 2010)

2678 compact sources
in $l=59^\circ$ and $l=30^\circ$ SDP
fields

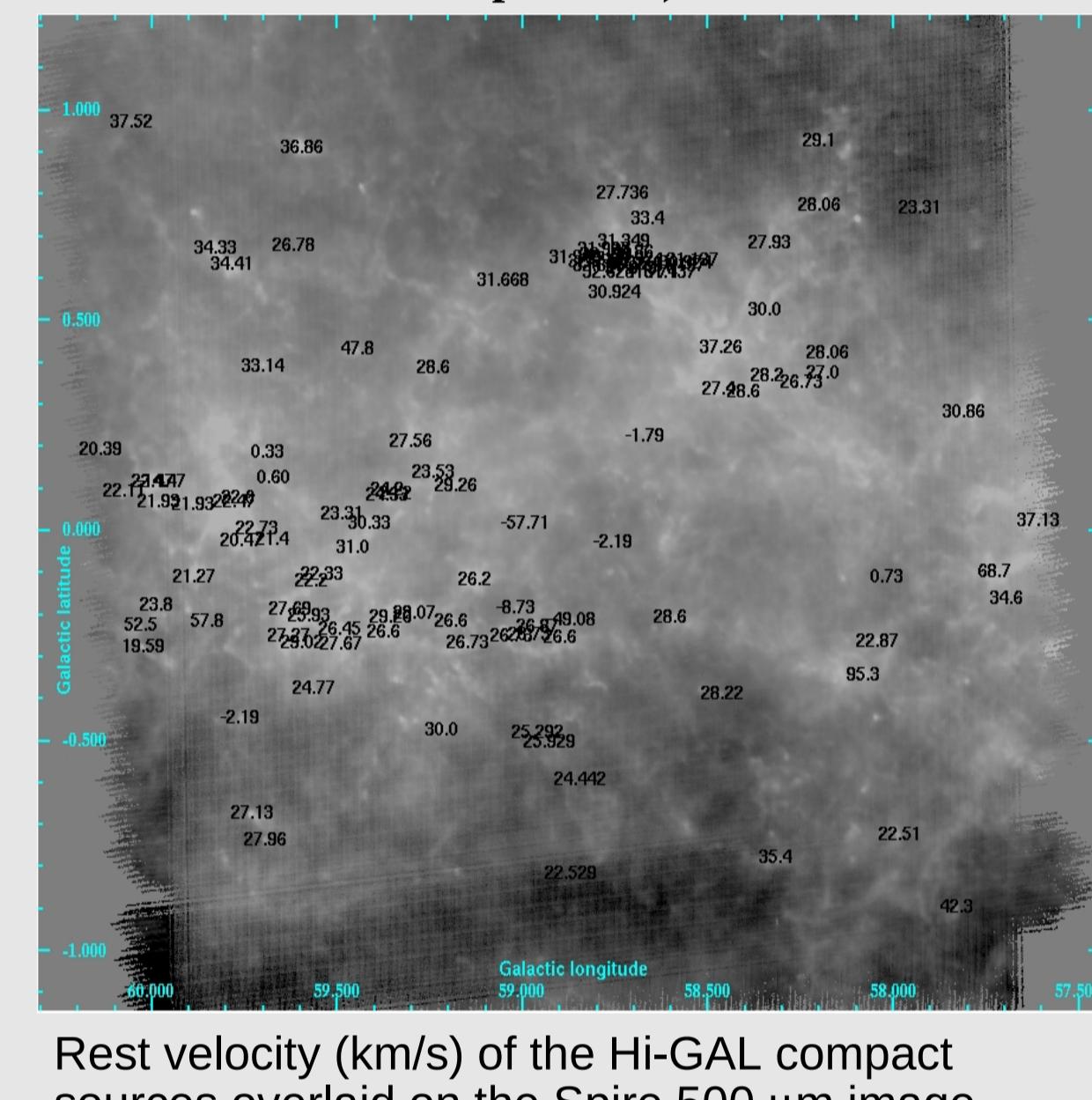
VELOCITY

NH_3 observations

(Medicina, Onsala, Effelsberg Telescopes)

$^{13}\text{CO} - \text{C}^{18}\text{O}$ line profiles extracted from datacubes (FCRAO,
Galactic Ring Survey)

709 Hi-GAL sources with an unambiguous velocity determination
(only one line detected in the spectrum)



Hi-GAL sources associated with HII regions
from velocity and spatial considerations

HII regions

=> Stellar distance for optical HII region, if exciting stars identified (e.g. Forbes 1989)

=> Maser parallax distance of star-forming region (e.g. Xu et al. 2009)

=> Near/far distance determination: HI absorption/emission, Self-absorption methods (e.g. Anderson & Bania 2009)

=> Near/far distance determination: H_2CO or OH absorption lines velocity (e.g. Sewilo et al. 2004)

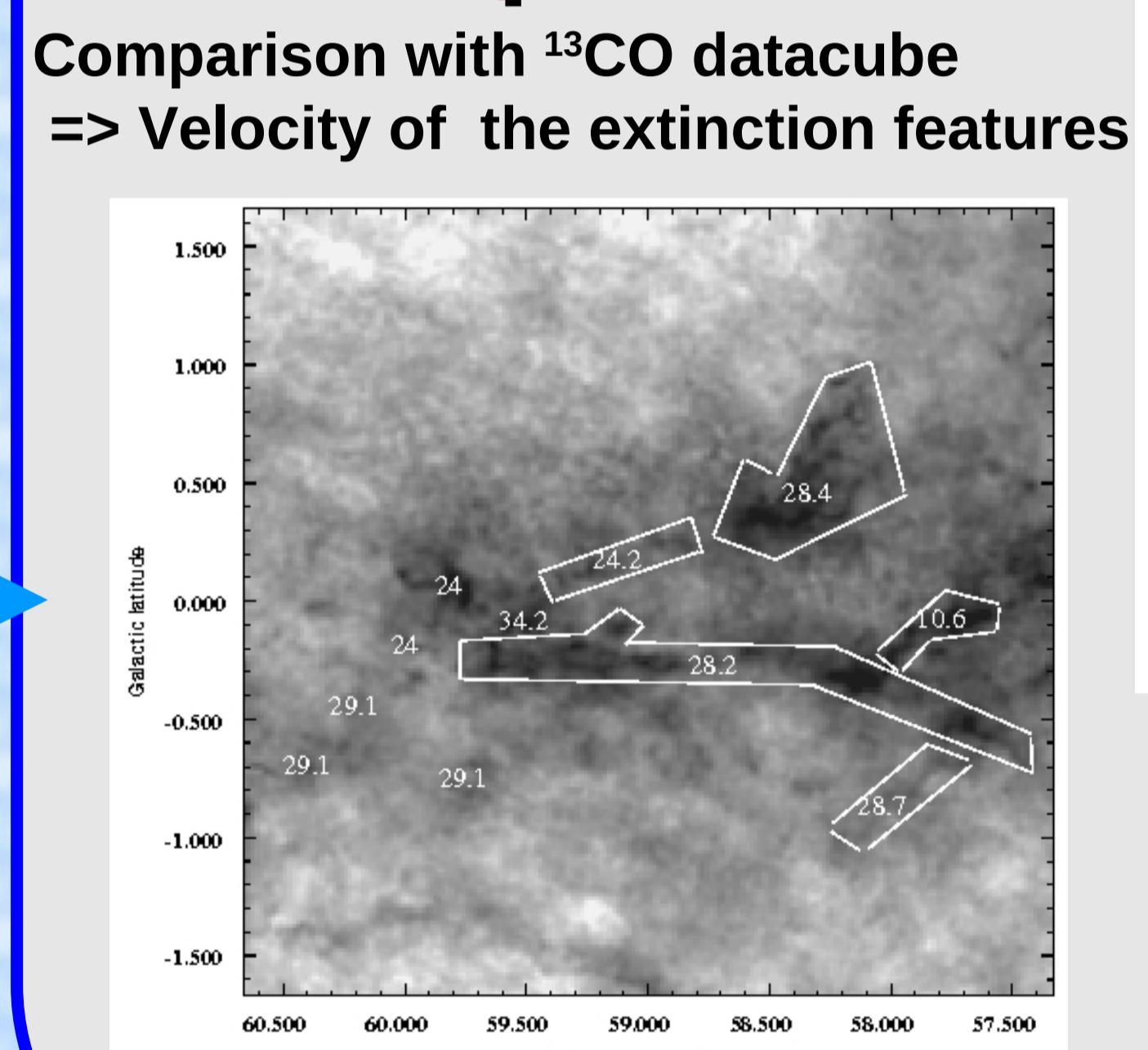


Star-forming complexes

=> Giant Molecular Clouds (e.g. Dame et al. 1986)

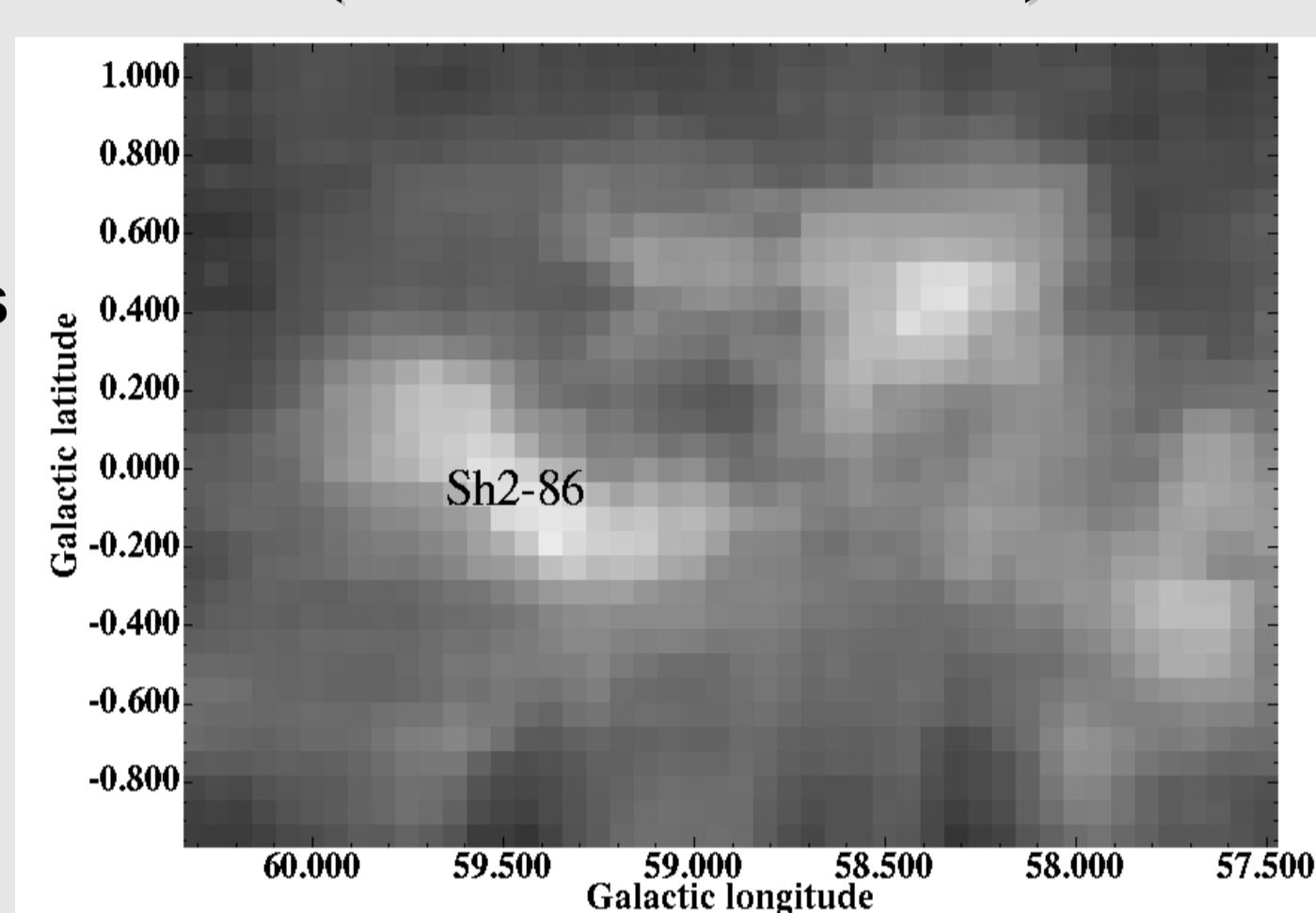
+ HII regions + molecular clouds (e.g. Roman-Duval et al. 2009)

Extinction map
2 MASS + Besançon model
(Schneider et al. 2010)



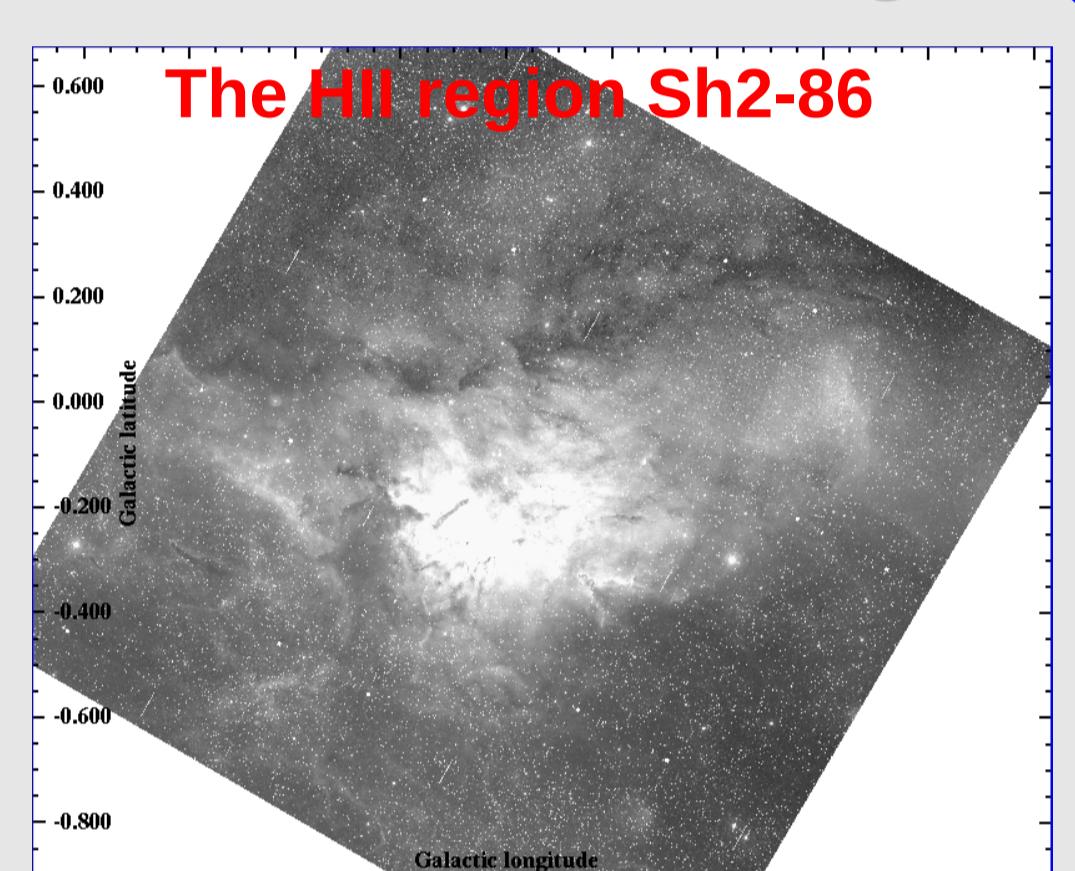
The values are velocity at the positions of the numbers or in the delimited areas. Extinction range: ~2 (white pixels) to > 15 mag (black pixels)

EXTINCTION
Extinction versus distance
2 MASS + Glimpse + Besançon model
(Marshall et al. 2009)

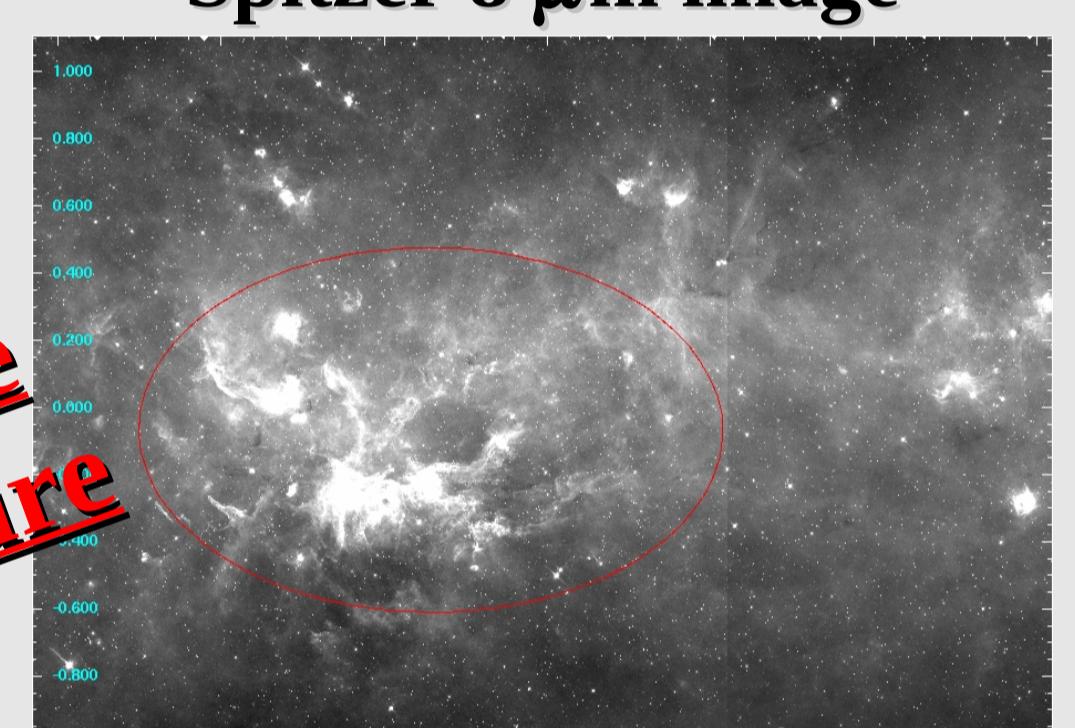


Extinction map (Av) for distances between 1.7 and 2.2 kpc. The extinction feature corresponding to the HII region Sh2-86 is identified. Extinction range: -0.4 (black pixels) to ~1.9 mag (white pixels)

Hα- IPHAS image



Spitzer 8 μm image



The red ellipse shows the extension of the Hα emission of Sh2-86.

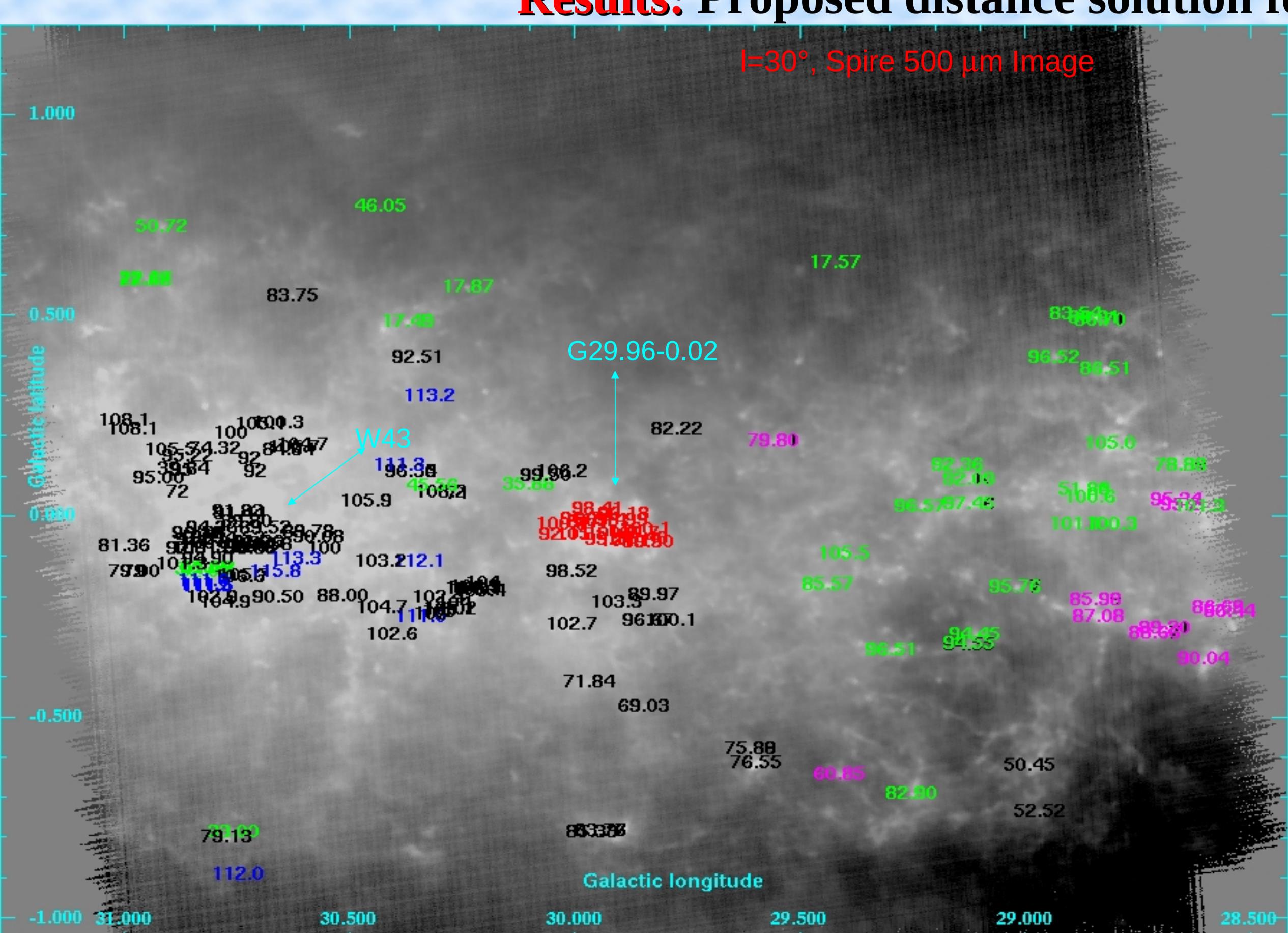
NEAR distance favoured IF source
associated with an extinction feature

Table A.2. Star-forming complexes in the $l=59^\circ$ field.

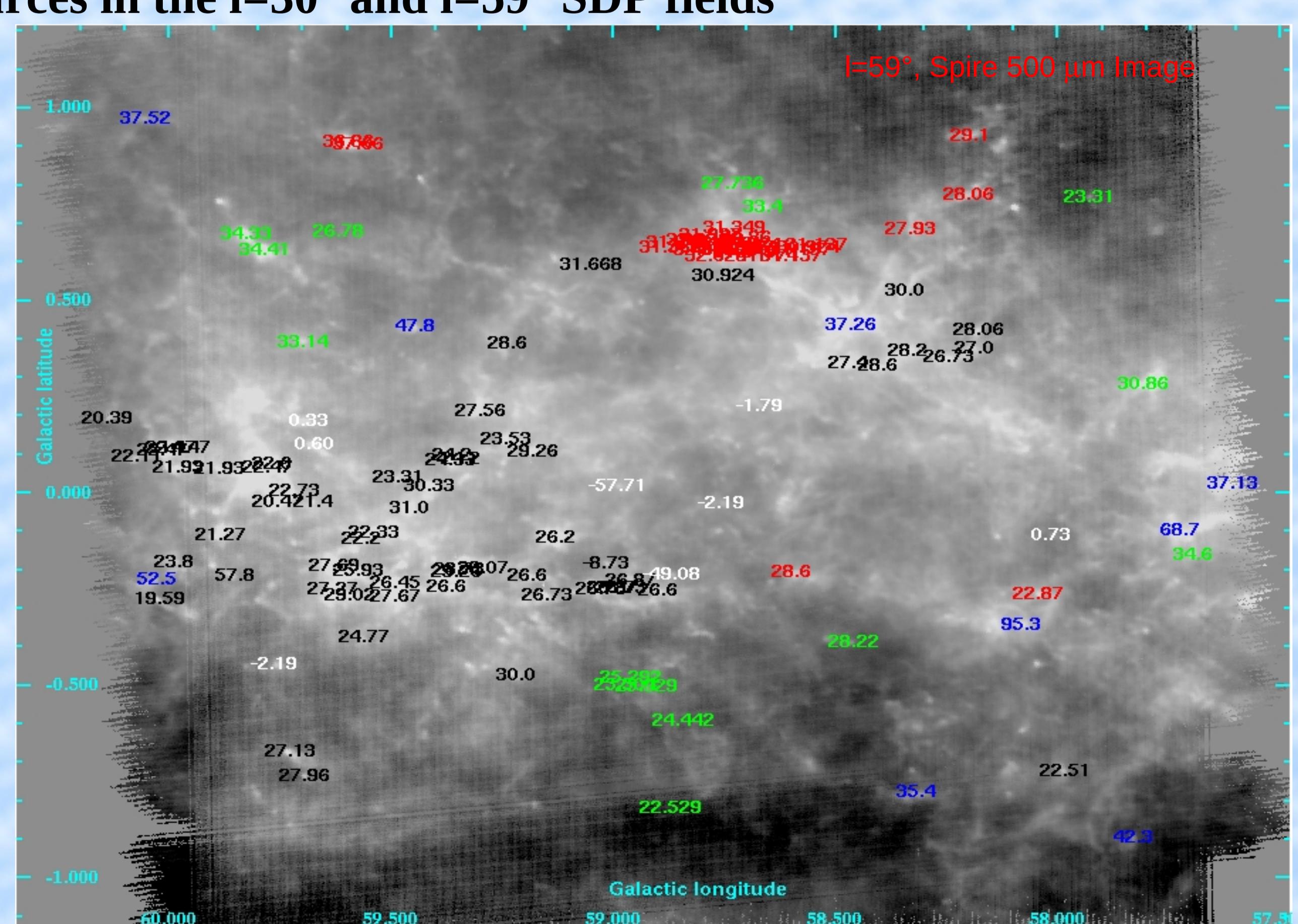
Complex number	Dist. (kpc)	optical H II reg.	source ident.	V_{Halpha} km s ⁻¹	$V_{\text{radio+comb.}}$ km s ⁻¹	$V_{\text{Molecular}}$ km s ⁻¹	V_{Abs} km s ⁻¹	Near/far
1	0.5/8.7/4.6		57.541-0.276			2.8	8.8	
	Far		Bx N 56.8+0.0(0.63,4.8)					
			Wat57.55-0.27			6.3		
			59.796+0.237			-6.0;-2.8		
			Bx W 57.2+0.0(0.61,8.2)				42.	
10							-7.	
3			Wat58.77+0.65			30.01	32.5	
/-/4.3			Wat59.60+0.92			46.4	37.3	
4	2.3/6.2/4.2	Sh2-86	59.529-0.181	26.4;27.9	29.4	26.8		F ^c
		Sh2-87	60.888-0.127	17.9	18.6	22.7		
		Sh2-88	61.470+0.90	19.3;19.8	27.3	22.9		
		Sh2-89	62.941+0.084	12.7	19.2	25.6		
		Sh2-90	63.176+0.460	22.5;18.0	15.3;19.4	22.2		
			Bx O 60.0+0.0(0.31,9)			24.5		
			Sol245 59.30-0.20			28		
			Sol246 60.00+0.10			22		
			Sol247 60.90-0.10			23		
			Sol248 61.50+0.10			22		
			Sol249 63.10+0.40			20		
			Wat59.36-0.21			26.6	29	
			Mas59.78+0.06					
2.16								

Column 9 references: a) Watson et al. (2003), b) Solomon et al. (1987), c) Kuchar & Bania (1994).

Results: Proposed distance solution for the 709 Hi-GAL sources in the $l=30^\circ$ and $l=59^\circ$ SDP fields



Velocity of the Hi-GAL sources. The colors correspond to: Black: probable association with W43 for sources around 94 km/s (~7.8 kpc) and association with Sh2-65 for sources around 51 km/s (3.3 kpc), Blue: tangent distance (~7.4 kpc), Green: far or tangent distance (e.g. ~12.4 kpc for sources >20 km/s, ~7.5 kpc for sources ~ 95 km/s), Magenta: Near distance (~5.3 kpc) and Red: probable association with G29.96-0.02 (near distance, ~ 6.2 kpc).



Velocity of the Hi-GAL sources. The colors correspond to: Black: sources probably associated with Sh2-86 (2.1 kpc), Red: near distance (~3 kpc) but not associated with Sh2-86, White: objects with negative velocity (no distance ambiguity, ~13.5 kpc) or velocity close to 0 km/s (far distance favoured, ~9.5 kpc), Blue: objects with forbidden velocity; tangent distance adopted (~ 4.3 kpc), and Green: far distance (~6 kpc).